Math 52H: Multilinear algebra, differential forms and Stokes' theorem

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Part I Multilinear Algebra

Chapter 1

Linear and multilinear functions

1.1 Dual space

Let V be a finite-dimensional real vector space. The set of all linear functions on V will be denoted by V^* .

Proposition 1.1. V^* is a vector space of the same dimension as V.

Proof. One can add linear functions and multiply them by real numbers:

$$(l_1 + l_2)(x) = l_1(x) + l_2(x)$$

 $(\lambda l)(x) = \lambda l(x) \text{ for } l, l_1, l_2 \in V^*, x \in V, \lambda \in \mathbb{R}$

It is straightforward to check that all axioms of a vector space are satisfied for V^* . Let us now check that dim $V = \dim V^*$.

Choose a basis
$$v_1 \dots v_n$$
 of V . For any $x \in V$ let $\begin{pmatrix} x_1 \\ \vdots \\ x_n \end{pmatrix}$ be its coordinates in the basis $v_1 \dots v_n$.

Notice that each coordinate x_1, \ldots, x_n can be viewed as a linear function on V. Indeed,

1) the coordinates of the sum of two vectors are equal to the sum of the corresponding coordinates;

2) when a vector is multiplied by a number, its coordinates are being multiplied by the same number.

Thus x_1, \ldots, x_n are vectors from the space V^* . Let us show now that they form a basis of V^* . Indeed, any linear function $l \in V^*$ can be written in the form $l(x) = a_1x_1 + \ldots + a_nx_n$ which means that l is a linear combination of $x_1 \ldots x_n$ with coefficients a_1, \ldots, a_n . Thus x_1, \ldots, x_n generate V^* . On the other hand, if $a_1x + \ldots + a_nx_n$ is the 0-function, then all the coefficients must be equal to 0; i.e. functions x_1, \ldots, x_n are linearly independent. Hence x_1, \ldots, x_n form a basis of V and therefore $\dim V^* = n = \dim V$.

The space V^* is called *dual* to V and the basis x_1, \ldots, x_n dual to $v_1 \ldots v_n$.

Exercise 1.2. Prove the converse: given any basis l_1, \ldots, l_n of V^* we can construct a dual basis w_1, \ldots, w_n of V so that the functions l_1, \ldots, l_n serve as coordinate functions for this basis.

Recall that vector spaces of the same dimension are isomorphic. For instance, if we fix bases in both spaces, we can map vectors of the first basis into the corresponding vectors of the second basis, and extend this map by linearity to an isomorphism between the spaces. In particular, sending a basis $S = \{v_1, \ldots, v_n\}$ of a space V into the dual basis x_1, \ldots, x_n of the dual space V^* we can establish an isomorphism $i_S : V \to V^*$. However, this isomorphism is not canonical, i.e. it depends on the choice of the basis v_1, \ldots, v_n .

If V is a Euclidean space, i.e. a space with a scalar product $\langle x, y \rangle$, then this allows us to define another isomorphism $V \to V^*$, different from the one described above. This isomorphism associates with a vector $v \in V$ a linear function $l_v(x) = \langle v, x \rangle$. We will denote the corresponding map $V \to V^*$ by \mathcal{D} . Thus we have $\mathcal{D}(v) = l_v$ for any vector $v \in V$.

Exercise 1.3. Prove that $\mathcal{D}: V \to V^*$ is an isomorphism. Show that $\mathcal{D} = i_S$ for any orthonormal basis S.

The isomorphism \mathcal{D} is independent of a choice of an orthonormal basis, but is still not completely canonical: it depends on a choice of a scalar product. However, when talking about *Euclidean spaces*, this isomorphism is canonical.

¹It is sometimes customary to denote dual bases in V and V^* by the same letters but using lower indices for V and upper indices for V^* , e.g. v_1, \ldots, v_n and v^1, \ldots, v^n . However, in these notes we do not follow this convention.

Remark 1.4. The definition of the dual space V^* also works in the infinite-dimensional case.

Exercise 1.5. Show that both maps i_S and D are injective in the infinite case as well. However, neither one is surjective if V is infinite-dimensional.

1.2 Canonical isomorphism between $(V^*)^*$ and V

The space $(V^*)^*$, dual to the dual space V, is canonically isomorphic in the finite-dimensional case to V. The word canonically means that the isomorphism is "god-given", i.e. it is independent of any additional choices.

When we write f(x) we usually mean that the function f is fixed but the argument x can vary. However, we can also take the opposite point of view, that x is fixed but f can vary. Hence, we can view the point x as a function on the vector space of functions.

If $x \in V$ and $f \in V^*$ then the above argument allows us to consider vectors of the space V as linear functions on the dual space V^* . Thus we can define a map $I: V \to V^{**}$ by the formula

$$x \mapsto I(x) \in (V^*)^*$$
, where $I(x)(l) = l(x)$ for any $l \in V^*$.

Exercise 1.6. Prove that if V is finite-dimensional then I is an isomorphism. What can go wrong in the infinite-dimensional case?

1.3 The map A^*

Given a map $\mathcal{A}: V \to W$ one can define a dual map $\mathcal{A}^*: W^* \to V^*$ as follows. For any linear function $l \in W^*$ we define the function $\mathcal{A}^*(l) \in V^*$ by the formula $\mathcal{A}^*(l)(x) = l(A(x)), x \in V$. In other words, $\mathcal{A}^*(l) = l \circ \mathcal{A}^2$.

² In fact, the above formula makes sense in much more general situation. Given any map $\Phi: X \to Y$ between two sets X and Y the formula $\Phi^*(h) = h \circ \Phi$ defines a map $\Phi^*: F(Y) \to F(X)$ between the spaces of functions on Y and X. Notice that this map goes in the direction opposite to the direction of the map Φ .

Given bases $\mathcal{B}_v = \{v_1, \dots, v_n\}$ and $\mathcal{B}_w = \{w_1, \dots, w_k\}$ in the vector spaces V and W one can associate with the map \mathcal{A} a matrix $A = M_{\mathcal{B}_v \mathcal{B}_w}(\mathcal{A})$. Its columns are coordinates of the vectors $\mathcal{A}(v_j), j = 1, \dots, n$, in the basis \mathcal{B}_w . Dual spaces V^* and W^* have dual bases $X = \{x_1, \dots, x_n\}$ and $Y = \{y_1, \dots, y_k\}$ which consist of coordinate functions corresponding to the basis \mathcal{B}_v and \mathcal{B}_w . Let us denote by A^* the matrix of the dual map A^* with respect to the bases Y and X, i. e. $A^* = M_{YX}(\mathcal{A}^*)$.

Proposition 1.7. The matrices A and A^* are transpose to each other, i.e. $A^* = A^T$.

Proof. By the definition of the matrix of a linear map we should take vectors of the basis $Y = \{y_1, \ldots, y_k\}$, apply to them the map \mathcal{A}^* , expand the images in the basis $X = \{x_1, \ldots, x_n\}$ and write the components of these vectors as *columns* of the matrix \mathcal{A}^* . Set $\tilde{y}_i = \mathcal{A}^*(y_i)$, $i = 1, \ldots, k$. For any vector $u = \sum_{j=i}^n u_j v_j \in V$, we have $\tilde{y}_i(u) = y_i(\mathcal{A}(u))$. The coordinates of the vector $\mathcal{A}(u)$ in the

basis w_1, \ldots, w_k may be obtained by multiplying the matrix A by the column $\begin{pmatrix} u_1 \\ \vdots \\ u_n \end{pmatrix}$. Hence,

$$\tilde{y}_i(u) = y_i(\mathcal{A}(u)) = \sum_{j=1}^n a_{ij}u_j.$$

But we also have

$$\sum_{j=1}^{n} a_{ij} x_j(u) = \sum_{j=1}^{n} a_{ij} u_j.$$

Hence, the linear function $\tilde{y}_i \in V^*$ has an expansion $\sum_{j=1}^n a_{ij}x_j$ in the basis $X = \{x_1, \dots x_n\}$ of the

space V^* . Hence the *i*-th column of the matrix A^* equals $\begin{pmatrix} a_{i1} \\ \vdots \\ a_{in} \end{pmatrix}$, so that the whole matrix A^*

has the form

$$A^* = \begin{pmatrix} a_{11} & \cdots & a_{k1} \\ \cdots & \cdots & \cdots \\ a_{1n} & \cdots & a_{kn} \end{pmatrix} = A^T .$$

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Exercise 1.8. Given a linear map $A: V \to W$ with a matrix A, find $A^*(y_i)$.

Answer. The map \mathcal{A}^* sends the coordinate function y_i on W to the function $\sum_{j=1}^n a_{ij}x_j$ on V, i.e. to its expression in coordinates x_i .

Proposition 1.9. Consider linear maps

$$U \stackrel{\mathcal{A}}{\rightarrow} V \stackrel{\mathcal{B}}{\rightarrow} W$$
.

Then $(\mathcal{B} \circ \mathcal{A})^* = \mathcal{A}^* \circ \mathcal{B}^*$.

Proof. For any linear function $l \in W^*$ we have

$$(\mathcal{B} \circ \mathcal{A})^*(l)(x) = l(\mathcal{B}(\mathcal{A}(x)) = \mathcal{A}^*(\mathcal{B}^*(l))(x)$$

for any $x \in U$.

Exercise 1.10. Suppose that V is a Euclidean space and A is a linear map $V \to V$. Prove that for any two vectors $X, Y \in V$ we have

$$\langle \mathcal{A}(X), Y \rangle = \langle X, \mathcal{D}^{-1} \circ \mathcal{A}^* \circ \mathcal{D}(Y) \rangle.$$
 (1.3.1)

Solution. By definition of the operator \mathcal{D} we have

$$\langle X, \mathcal{D}^{-1}(Z) \rangle = Z(X)$$

for any vector $Z \in V^*$. Applying this to $Z = \mathcal{A}^* \circ \mathcal{D}(Y)$ we see that the right-hand send of (1.3.1) is equal to to $\mathcal{A}^* \circ \mathcal{D}(Y)(X)$. On the other hand, the left-hand side can be rewritten as $\mathcal{D}(Y)(\mathcal{A}(X))$.

Let us recall that if V is an Euclidean space, then operator $\mathcal{B}:V\to V$ is called adjoint to $\mathcal{A}:V\to V$ if for any two vectors $X,Y\in V$ one has

$$\langle \mathcal{A}(X), Y \rangle = \langle X, \mathcal{B}(Y) \rangle.$$

The adjoint operator always exist and unique. It is denoted by \mathcal{A}^* . Clearly, $(\mathcal{A}^*)^* = \mathcal{A}$. In any orthonormal basis the matrices of an operator and its adjoint are transpose to each other. An

operator $A:V\to V$ is called *self-adjoint* if $A^*=A$, or equivalently, if for any two vectors $X,Y\in V$ one has

$$\langle \mathcal{A}(X), Y \rangle = \langle X, \mathcal{A}(Y) \rangle.$$

The statement of Exercise 1.10 can be otherwise expressed by saying that the adjoint operator \mathcal{A}^* is equal to $\mathcal{D}^{-1} \circ \mathcal{A}^* \circ \mathcal{D} : V \to V$. In particular, an operator $\mathcal{A} : V \to V$ is self-adjoint if and only if $\mathcal{A}^* \circ \mathcal{D} = \mathcal{D} \circ \mathcal{A}$.

Remark 1.11. As it follows from Proposition 1.7 and Exercise 1.10 the matrix of a self-adjoint operator in any *orthonormal* basis is symmetric. This is not true in an arbitrary basis.

1.4 Multilinear functions

A function $l(X_1, X_2, ..., X_k)$ of k vector arguments $X_1, ..., X_k \in V$ is called k-linear (or multilinear) if it is linear with respect to each argument when all other arguments are fixed. We say bilinear instead of 2-linear. Multilinear functions are also called tensors. Sometimes, one may also say a "k-linear form", or simply k-form instead of a "k-linear functions". However, we will reserve the term k-form for a skew-symmetric tensors which will be defined in Section 2.3 below.

If one fixes a basis $v_1
ldots v_n$ in the space V then with each bilinear function f(X,Y) one can associate a square $n \times n$ matrix as follows. Set $a_{ij} = f(v_i, v_j)$. Then $A = (a_{ij})_{i,j=1,\dots,n}$ is called the matrix of the function f in the basis v_1, \dots, v_n . For any 2 vectors

$$X = \sum_{1}^{n} x_i v_i, Y = \sum_{1}^{n} y_j v_j$$

we have

$$f(X,Y) = f\left(\sum_{i=1}^{n} x_i v_i, \sum_{j=1}^{n} y_j v_j\right) = \sum_{i,j=1}^{n} x_i y_j f(v_i, v_j) = \sum_{i,j=1}^{n} a_{ij} x_i y_j = X^T A Y.$$

Exercise 1.12. How does the matrix of a bilinear function depend on the choice of a basis?

Answer. The matrices A and \tilde{A} of the bilinear form f(x,y) in the bases v_1, \ldots, v_n and $\tilde{v}_1, \ldots, \tilde{v}_n$ are related by the formula $\tilde{A} = C^T A C$, where C is the matrix of transition from the basis $v_1 \ldots v_n$

to the basis $\tilde{v}_1 \dots \tilde{v}_n$, i.e the matrix whose columns are the coordinates of the basis $\tilde{v}_1 \dots \tilde{v}_n$ in the basis $v_1 \dots v_n$.

Similarly, with a k-linear function $f(X_1, ..., X_k)$ on V and a basis $v_1, ..., v_n$ one can associate a "k-dimensional" matrix

$$A = \{a_{i_1 i_2 \dots i_k}; \ 1 \le i_1, \dots, i_k \le n\},\$$

where

$$a_{i_1 i_2 \dots i_k} = f(v_{i_1}, \dots, v_{i_k}).$$

If
$$X_i = \sum_{j=1}^{n} x_{ij} v_j$$
, $i = 1, ..., k$, then

$$f(X_1, \dots, X_k) = \sum_{i_1, i_2, \dots i_k = 1}^n a_{i_1 i_2 \dots i_k} x_{1 i_1} x_{2 i_2} \dots x_{k i_k} ,$$

see Proposition 2.1 below.

1.5 Quotient space

Let V be a vector space and $L \subset V$ be its linear subspace. Given a vector $a \in V$ let us denote byL_a the affine subspace $a + L = \{a + x; x \in L\}$. Note that two affine subspaces L_a and L_b , $a, b \in V$ concide if $a - b \in L$ and are disjoint if $a - b \notin L$. Consider the set, denoted by V/L of all affine subspaces parallel to L. The set can be made into a vector space by defining the operations by the formulas

$$L_a + L_b := L_{a+b}, \ \lambda L_a := L_{\lambda a}; \ a, b \in V, \ \lambda \in \mathbb{R}$$

If a' and b' are other vectors in V such that $a' - a, b' - b \in L$ then $(a' + b') - (a + b) \in L$ and $\lambda a' - \lambda a \in L$, and hence $L_{a'+b'} = L_{a+b}$, $L_{\lambda a'} = \mathcal{L}_{\lambda a}$.

The vector space V/L is called the quotient space of V by L.

In other words, we can say that V/L is obtained from V by identifying vectors which differ by a vector in L. The operations in V then naturally descend to the operations on the quotient space.

If $N \subset V$ be any linear subspace such that $\dim L + \dim N = \dim V$ and $L \cap N = 0$ then the map $N \to V/L$ given by the formula $x \in N \mapsto L_x \in V/L$ is an isomorphism (why?). In particular, $\dim(V/L) = \dim V - \dim L = \operatorname{codim}_V L$.

If V is a Euclidean space then we can choose as N the orthogonal complement L^{\perp} of L in V, and thus L/V is isomorphic to L^{\perp} . The advantage of the quotient construction that it is *canonical* while L^{\perp} depends on a choice of the Euclidean structure (scalar product).

1.6 Symmetric bilinear functions and quadratic forms

A function $Q:V\to\mathbb{R}$ on a vector space V is called *quadratic* if there exists a bilinear function f(X,Y) such that

$$Q(X) = f(X, X), X \in V.$$
(1.6.1)

One also uses the term quadratic form. The bilinear function f(X,Y) is not uniquely determined by the equation (1.6.1). For instance, all the bilinear functions x_1y_2 , x_2y_1 and $\frac{1}{2}(x_1y_2 + x_2y_1)$ on \mathbb{R}^2 define the same quadratic form x_1x_2 .

On the other hand, there is a 1-1 correspondence between quadratic form and symmetric bilinear functions. A bilinear function f(X,Y) is called symmetric if f(X,Y) = f(Y,X) for all $X,Y \in V$.

Lemma 1.13. Given a quadratic form $Q: V \to \mathbb{R}$ there exists a unique symmetric bilinear form f(X,Y) such that $Q(X) = f(X,X), X \in V$.

Proof. If Q(X) = f(X, X) for a symmetric f then

$$Q(X + Y) = f(X + Y, X + Y) = f(X, X) + f(X, Y) + f(Y, X) + f(Y, Y)$$
$$= Q(X) + 2f(X, Y) + Q(Y),$$

and hence

$$f(X,Y) = \frac{1}{2} (Q(X+Y) - Q(X) - Q(Y)). \tag{1.6.2}$$

I leave it as an exercise to check that the formula (1.6.2) always defines a symmetric bilinear function.

Let $S = \{v_1, \ldots, v_n\}$ is a basis of V. The matrix $A = (a_{ij})$ of a symmetric bilinear form f(X, Y) in the basis S is called also the matrix of the corresponding quadratic form Q(X) = f(X, X). This matrix is symmetric, and

$$Q(X) = \sum_{ij} a_{ij} x_i x_j = a_{11} x_1^2 + \dots + a_{nn} x_n^2 + 2 \sum_{i < j} a_{ij} x_i x_j.$$

Thus the matrix A is diagonal if and only if the quadratic form Q is the sum of squares (with coefficients). Let us recall that if one changes the basis S to a basis $\widetilde{S} = \{\widetilde{v}_1, \dots, \widetilde{v}_n\}$ then the matrix of a bilinear form f changes to $\widetilde{C} = C^T A C$.

Exercise 1.14. (Sylvester's inertia law) Prove that there is always exists a basis $\widetilde{S} = \{\widetilde{v}_1, \dots, \widetilde{v}_n\}$ in which a quadratic form Q is reduced to a sum of squares. The number of positive, negative and zero coefficients with the squares is independent of the choice of the basis.

Thus, in some coordinate system a quadratic form can always be written as

$$-\sum_{1}^{k} x_{i}^{2} + \sum_{k+1}^{k+l} x_{j}^{2}, \ k+l \le n.$$

The number k of negative squares is called the *negative index* or simply the *index* of the quadratic form Q, the total number of k + l of non-zero squares is called the *rank* of the form. It coincides with the rank of the matrix of the form in *any* basis. A bilinear (and quadratic) form is called *non-degenerate* if its rank is maximal possible, i.e. equal to n. For a non-degenerate quadratic for Q the difference l - k between the number of positive and negative squares is called the *signature*.

A quadratic form Q is called *positive definite* if $Q(X) \geq 0$ and if Q(X) = 0 then X = 0. Equivalently, one can say that a form is positive definite if it is non-degenerate and its negative index is equal to 0.

Chapter 2

Tensor and exterior products

2.1 Tensor product

Given a k-linear function ϕ and a l-linear function ψ , one can form a (k+l)-linear function, which will be denoted by $\phi \otimes \psi$ and called the tensor product of the functions ϕ and ψ . By definition

$$\phi \otimes \psi(X_1, \dots, X_k, X_{k+1}, \dots, X_{k+l}) := \phi(X_1, \dots, X_k) \cdot \psi(X_{k+1}, \dots, X_{k+l}).$$

For instance, the tensor product two linear functions l_1 and l_2 is a bilinear function $l_1 \otimes l_2$ defined by the formula

$$l_1 \otimes l_2(U, V) = l_1(U)l_2(V).$$

Let $v_1
ldots v_n$ be a basis in V and $x_1,
ldots, x_n$ a dual basis in V^* , i.e. $x_1,
ldots, x_n$ are coordinates of a vector with respect to the basis $v_1,
ldots, v_n$.

The tensor product $x_i \otimes x_j$ is a bilinear function $x_i \otimes x_j(Y, Z) = y_i z_j$. Thus a bilinear function f with a matrix A can be written as a linear combination of the functions $x_i \otimes x_j$ as follows:

$$f = \sum_{i,j=1}^{n} a_{ij} x_i \otimes x_j,$$

where a_{ij} is the matrix of the form f in the basis $v_1 \dots v_n$. Similarly any k-linear function f with a "k-dimensional" matrix $A = \{a_{i_1 i_2 \dots i_k}\}$ can be written (see 2.1 below) as a linear combination of functions

$$x_{i_1} \otimes x_{i_2} \otimes \cdots \otimes x_{i_k}, \ 1 \leq i_1, i_2, \ldots, i_k \leq n$$
.

Namely, we have

$$f = \sum_{i_1, i_2, \dots i_k = 1}^n a_{i_1 i_2 \dots i_k} x_{i_1} \otimes x_{i_2} \otimes \dots \otimes x_{i_k}.$$

2.2 Spaces of multilinear functions

All k-linear functions, or k-tensors, on a given n-dimensional vector space V themselves form a vector space, which will be denoted by $V^{*\otimes k}$. The space $V^{*\otimes 1}$ is, of course, just the dual space V^* .

Proposition 2.1. Let $v_1, \ldots v_n$ be a basis of V, and x_1, \ldots, x_k be the dual basis of V^* formed by coordinate functions with respect to the basis V. Then n^k k-linear functions $x_{i_1} \otimes \cdots \otimes x_{i_k}$, $1 \leq i_1, \ldots, i_k \leq n$, form a basis of the space $V^{* \otimes k}$.

Proof. Take a k-linear function F from $V^{*\otimes k}$ and evaluate it on vectors v_{i_1}, \ldots, v_{i_k} :

$$F(v_{i_1},\ldots,v_{i_k})=a_{i_1\ldots i_k}$$
.

We claim that we have

$$F = \sum_{1 \leq i_1, \dots, i_k \leq n} a_{i_1 \dots i_k} x_{i_1} \otimes \dots \otimes x_{i_k}.$$

Indeed, the functions on the both sides of this equality being evaluated on any set of k basic vectors v_{i_1}, \ldots, v_{i_k} , give the same value $a_{i_1 \ldots i_k}$. The same argument shows that if $\sum_{1 \leq i_1, \ldots, i_k \leq n} a_{i_1 \ldots i_k} x_{i_1} \otimes \cdots \otimes x_{i_k} = 0$, then all coefficients $a_{i_1 \ldots i_k}$ should be equal to 0. Hence the functions $x_{i_1} \otimes \cdots \otimes x_{i_k}$, $1 \leq i_1, \ldots, i_k \leq n$, are linearly independent, and therefore form a basis of the space $V^{* \otimes k}$

Similar to the case of spaces of linear functions, a linear map $\mathcal{A}: V \to W$ induces a linear map $\mathcal{A}^*: W^{*\otimes k} \to V^{*\otimes k}$, which sends a k-linear function $F \in W^{*\otimes k}$ to a k-linear function $\mathcal{A}^*(F) \in V^{*\otimes k}$, defined by the formula

$$\mathcal{A}^*(F)(X_1,\ldots,X_k) = F(\mathcal{A}(X_1),\ldots,\mathcal{A}(X_k))$$
 for any vectors $X_1,\ldots X_k \in V$.

Exercise 2.2. Suppose V is provided with a basis v_1, \ldots, v_n and $x_{i_1} \otimes \cdots \otimes x_{i_k}$, $1 \leq i_1, \ldots, i_k \leq n$, is the corresponding basis of the space $V^{*\otimes k}$. Suppose that the map $A: V \to V$ has a matrix $A = (a_{ij})$ in the basis v_1, \ldots, v_n . Find the matrix of the map $A^*: V^{*\otimes k} \to V^{*\otimes k}$ in the basis $x_{i_1} \otimes \cdots \otimes x_{i_k}$.

2.3 Symmetric and skew-symmetric tensors

A multilinear function (tensor) is called *symmetric* if it remains unchanged under the transposition of any two of its arguments:

$$f(X_1, ..., X_i, ..., X_j, ..., X_k) = f(X_1, ..., X_j, ..., X_i, ..., X_k)$$

Equivalently, one can say that a k-tensor f is symmetric if

$$f(X_{i_1},\ldots,X_{i_k})=f(X_1,\ldots,X_k)$$

for any permutation i_1, \ldots, i_k of indices $1, \ldots, k$.

Exercise 2.3. Show that a bilinear function f(X,Y) is symmetric if and only if its matrix (in any basis) is symmetric.

Notice that the tensor product of two symmetric tensors usually is not symmetric.

Example 2.4. Any linear function is (trivially) symmetric. However, the tensor product of two functions $l_1 \otimes l_2$ is not a symmetric bilinear function unless l_1 is proportional to l_2 . On the other hand, the function $l_1 \otimes l_2 + l_2 \otimes l_1$ is symmetric.

A tensor is called *skew-symmetric* (or anti-symmetric) if it changes its sign when one transposes any two of its arguments:

$$f(X_1, ..., X_i, ..., X_j, ..., X_k) = -f(X_1, ..., X_j, ..., X_i, ..., X_k).$$

Equivalently, one can say that a k-tensor f is anti-symmetric if

$$f(X_{i_1}, \dots, X_{i_k}) = (-1)^{\operatorname{inv}(i_1 \dots i_k)} f(X_1, \dots, X_k)$$

for any permutation i_1, \ldots, i_k of indices $1, \ldots, k$, where $\operatorname{inv}(i_1 \ldots i_k)$ is the number of inversions in the permutation i_1, \ldots, i_k . Recall that two indices i_k, i_l form an *inversion* if k < l but $i_k > i_l$.

The matrix A of a bilinear skew-symmetric function is skew-symmetric, i.e.

$$A^T = -A$$
.

Example 2.5. The determinant $det(X_1, ..., X_n)$ (considered as a function of columns $X_1, ..., X_n$ of a matrix) is a skew-symmetric n-linear function.

Exercise 2.6. Prove that any n-linear skew-symmetric function on \mathbb{R}^n is proportional to the determinant.

Linear functions are trivially anti-symmetric (as well as symmetric).

As in the symmetric case, the tensor product of two skew-symmetric functions is not skew-symmetric. We will define below in Section 2.5 a new product, called an *exterior product* of skew-symmetric functions, which will again be a skew-symmetric function.

2.4 Symmetrization and anti-symmetrization

The following constructions allow us to create symmetric or anti-symmetric tensors from arbitrary tensors. Let $f(X_1, \ldots, X_k)$ be a k-tensor. Set

$$f^{\text{sym}}(X_1, \dots, X_k) := \sum_{(i_1 \dots i_k)} f(X_{i_1}, \dots, X_{i_k})$$

and

$$f^{\text{asym}}(X_1, \dots, X_k) := \sum_{(i_1 \dots i_k)} (-1)^{\text{inv}(i_1, \dots, i_k)} f(X_{i_1}, \dots, X_{i_k})$$

where the sums are taken over all permutations i_1, \ldots, i_k of indices $1, \ldots, k$. The tensors f^{sym} and f^{asym} are called, respectively, symmetrization and anti-symmetrization of the tensor f. It is now easy to see that

Proposition 2.7. The function f^{sym} is symmetric. The function f^{asym} is skew-symmetric. If f is symmetric then $f^{\text{sym}} = k!f$ and $f^{\text{asym}} = 0$. Similarly, if f is anti-symmetric then $f^{\text{asym}} = k!f$, $f^{\text{sym}} = 0$.

Exercise 2.8. Let x_1, \ldots, x_n be coordinate function in \mathbb{R}^n . Find $(x_1 \otimes x_2 \otimes \ldots \otimes x_n)^{asym}$.

Answer. The determinant.

2.5 Exterior product

For our purposes skew-symmetric functions will be more important. Thus we will concentrate on studying operations on them.

Skew-symmetric k-linear functions are also called exterior k-forms. Let ϕ be an exterior k-form and ψ an exterior l-form. We define an exterior (k+l)-form $\psi \wedge \psi$, the exterior product of ϕ and ψ , as

$$\phi \wedge \psi := \frac{1}{k! l!} (\phi \otimes \psi)^{\operatorname{asym}}.$$

In other words,

$$\phi \wedge \psi(X_1, \dots, X_k, X_{k+1}, \dots, X_{k+l}) = \frac{1}{k! l!} \sum_{i_1, \dots, i_{k+l}} (-1)^{\operatorname{inv}(i_1, \dots, i_{k+l})} \phi(X_{i_1}, \dots, X_{i_k}) \psi(X_{i_{k+1}}, \dots, X_{i_{k+l}}),$$

where the sum is taken over all permutations of indices $1, \ldots, k+l$.

Note that because the anti-symmetrization of an anti-symmetric k-tensor amounts to its multiplication by k!, we can also write

$$\phi \wedge \psi(X_1, \dots, X_{k+l}) = \sum_{i_1 < \dots < i_k, i_{k+1} < \dots < i_{k+l}} (-1)^{\operatorname{inv}(i_1, \dots, i_{k+l})} \phi(X_{i_1}, \dots, X_{i_k}) \psi(X_{i_{k+1}}, \dots, X_{i_{k+l}}),$$

where the sum is taken over all permutations i_1, \ldots, i_{k+l} of indices $1, \ldots, k+l$.

Exercise 2.9. The exterior product operation has the following properties:

- For any exterior k-form ϕ and exterior l-form ψ we have $\phi \wedge \psi = (-1)^{kl} \psi \wedge \phi$.
- Exterior product is linear with respect to each factor:

$$(\phi_1 + \phi_2) \wedge \psi = \phi_1 \wedge \psi + \phi_2 \wedge \psi$$
$$(\lambda \phi) \wedge \psi = \lambda(\phi \wedge \psi)$$

for k-forms ϕ, ϕ_1, ϕ_2 , l-form ψ and a real number λ .

• Exterior product is associative: $(\phi \wedge \psi) \wedge \omega = \phi \wedge (\psi \wedge \omega)$.

First two properties are fairly obvious. To prove associativity one can check that both sides of the equality $(\phi \wedge \psi) \wedge \omega = \phi \wedge (\psi \wedge \omega)$ are equal to

$$\frac{1}{k!l!m!}(\phi\otimes\psi\otimes\omega)^{\mathrm{asym}}.$$

In particular, if ϕ, ψ and ω are 1-forms, i.e. if k = l = m = 1 then

$$\psi \wedge \phi \wedge \omega = (\phi \otimes \psi \otimes \omega)^{\text{asym}}$$
.

This formula can be generalized for computing the exterior product of any number of 1-forms:

$$\phi_1 \wedge \dots \wedge \phi_k = (\phi_1 \otimes \dots \otimes \phi_k)^{\text{asym}}. \tag{2.5.1}$$

Example 2.10.
$$x_1 \wedge x_2 = x_1 \otimes x_2 - x_2 \otimes x_1$$
. For 2 vectors, $U = \begin{pmatrix} u_1 \\ \vdots \\ u_n \end{pmatrix}$, $V = \begin{pmatrix} v_1 \\ \vdots \\ v_n \end{pmatrix}$, we have

$$x_1 \wedge x_2(U, V) = u_1 v_2 - u_2 v_1 = \begin{vmatrix} u_1 & v_1 \\ u_2 & v_2 \end{vmatrix}.$$

For 3 vectors U, V, W we have

$$x_1 \wedge x_2 \wedge x_3(U, V, W) =$$

$$= x_1 \wedge x_2(U, V)x_3(W) + x_1 \wedge x_2(V, W)x_3(U) + x_1 \wedge x_2(W, U)x_3(V) =$$

$$(u_1v_2 - u_2v_1)w_3 + (v_1w_2 - v_2w_1)u_3 + (w_1u_2 - w_2u_1)v_3 = \begin{vmatrix} u_1 & v_1 & w_1 \\ u_2 & v_2 & w_2 \\ u_3 & v_3 & w_3 \end{vmatrix}.$$

The last equality is just the expansion formula of the determinant according to the last row.

Proposition 2.11. Any exterior 2-form f can be written as

$$f = \sum_{1 \le i < j \le n} a_{ij} \, x_i \wedge x_j$$

Proof. We had seen above that any bilinear form can be written as $f = \sum_{ij} a_{i,j} x_i \otimes x_j$. If f is skew-symmetric then the matrix $A = (a_{ij})$ is skew-symmetric, i.e. $a_{ii} = 0$, $a_{ij} = -a_{ji}$ for $i \neq j$.

Thus,
$$f = \sum_{1 \le i < j \le n} a_{ij} (x_i \otimes x_j - x_j \otimes x_i) = \sum_{1 \le i < j \le n} a_{ij} x_i \wedge x_j$$
.

Exercise 2.12. Prove that any exterior k-form f can be written as

$$f = \sum_{1 \le i, < \dots < i_k \le n} a_{i_1 \dots i_k} \, x_{i_1} \wedge x_{i_2} \wedge \dots x_{i_k}.$$

The following proposition can be proven by induction over k, similar to what has been done in Example 2.10 for the case k = 3.

Proposition 2.13. For any k 1-forms l_1, \ldots, l_k and k vectors X_1, \ldots, X_k we have

$$l_{1} \wedge \dots \wedge l_{k}(X_{1}, \dots, X_{k}) = \begin{vmatrix} l_{1}(X_{1}) & \dots & l_{1}(X_{k}) \\ \dots & \dots & \dots \\ l_{k}(X_{1}) & \dots & l_{k}(X_{k}) \end{vmatrix} .$$
 (2.5.2)

Corollary 2.14. The 1-forms l_1, \ldots, l_k are linearly dependent as vectors of V^* if and only if $l_1 \wedge \ldots \wedge l_k = 0$. In particular, $l_1 \wedge \ldots \wedge l_k = 0$ if $k > n = \dim V$.

Proof. If l_1, \ldots, l_k are dependent then for any vectors $X_1, \ldots, X_k \in V$ the rows of the determinant in the equation (2.13) are linearly dependent. Therefore, this determinant is equal to 0, and hence $l_1 \wedge \ldots \wedge l_k = 0$. In particular, when k > n then the forms l_1, \ldots, l_k are dependent (because $\dim V^* = \dim V = n$).

On the other hand, if l_1, \ldots, l_k are linearly independent, then the vectors $l_1, \ldots, l_k \in V^*$ can be completed to form a basis $l_1, \ldots, l_k, l_{k+1}, \ldots, l_n$ of V^* . According to Exercise 1.2 there exists a basis w_1, \ldots, w_n of V that is dual to the basis l_1, \ldots, l_n of V^* . In other words, l_1, \ldots, l_n can be viewed as coordinate functions with respect to the basis w_1, \ldots, w_n . In particular, we have $l_i(w_j) = 0$ if $i \neq j$ and $l_i(w_i) = 1$ for all $i, j = 1, \ldots, n$. Hence we have

$$l_1 \wedge \cdots \wedge l_k(w_1, \dots, w_k) = \begin{vmatrix} l_1(w_1) & \dots & l_1(w_k) \\ \dots & \dots & \dots \\ l_k(w_1) & \dots & l_k(w_k) \end{vmatrix} = \begin{vmatrix} 1 & \dots & 0 \\ \dots & 1 & \dots \\ 0 & \dots & 1 \end{vmatrix} = 1,$$

i.e. $l_1 \wedge \cdots \wedge l_k \neq 0$.

Proposition 2.13 can be also deduced from formula (2.5.1).

Corollary 2.14 and Exercise 2.12 imply that there are no non-zero k-forms on an n-dimensional space for k > n.

2.6 Spaces of symmetric and skew-symmetric tensors

As was mentioned above, k-tensors on a vector space V form a vector space under the operation of addition of functions and multiplication by real numbers. We denoted this space by $V^{*\otimes k}$. Symmetric and skew-symmetric tensors form subspaces of this space $V^{*\otimes k}$, which we denote, respectively, by $S^k(V^*)$ and $\Lambda^k(V^*)$. In particular, we have

$$V^* = S^1(V^*) = \Lambda^1(V^*)$$

.

Exercise 2.15. What is the dimension of the spaces $S^k(V^*)$ and $\Lambda^k(V^*)$?

Answer.

$$\dim \Lambda^k(V^*) = \binom{n}{k} = \frac{n!}{k!(n-k)!}$$
$$\dim S^k(V^*) = \frac{(n+k-1)!}{k!(n-1)!}.$$

The basis of $\Lambda^k(V^*)$ is formed by exterior k-forms $x_{i_1} \wedge x_{i_2} \wedge \ldots \wedge x_{i_k}, 1 \leq i_1 < i_2 < \ldots < i_k \leq n$.

2.7 Operator A^* on spaces of tensors

For any linear operator $\mathcal{A}: V \to W$ we introduced above in Section 1.3 the notion of a dual linear operator $\mathcal{A}^*: W^* \to V^*$. Namely $\mathcal{A}^*(l) = l \circ \mathcal{A}$ for any element $l \in V^*$, which is just a linear function on V. In this section we extend this construction to k-tensors for $k \geq 1$, i.e. we will define a map $\mathcal{A}^*: W^{*\otimes k} \to V^{*\otimes k}$.

Given a k-tensor $\phi \in W^{*\otimes k}$ and k vectors $X_1, \ldots, X_k \in V$ we define

$$\mathcal{A}^*(\phi)(X_1,\ldots,X_k) = \phi\left(\mathcal{A}(X_1),\ldots,\mathcal{A}(X_k)\right).$$

Note that if ϕ is symmetric, or anti-symmetric, so is $\mathcal{A}^*(\phi)$. Hence, the map \mathcal{A}^* also induces the maps $S^k(W^*) \to S^k(V^*)$ and $\Lambda^k(W^*) \to \Lambda^k(V^*)$. We will keep the same notation \mathcal{A}^* for both of these maps as well.

Proposition 2.16. Let $A: V \to W$ be a linear map. Then

1.
$$\mathcal{A}^*(\phi \otimes \psi) = \mathcal{A}^*(\phi) \otimes \mathcal{A}^*(\psi)$$
 for any $\phi \in W^{*\otimes k}$, $\psi \in W^{*\otimes l}$;

2.
$$\mathcal{A}^*(\phi^{\text{asym}}) = (\mathcal{A}^*(\phi))^{\text{asym}}, \ \mathcal{A}^*(\phi^{\text{sym}}) = (\mathcal{A}^*(\phi))^{\text{sym}};$$

3.
$$\mathcal{A}^*(\phi \wedge \psi) = \mathcal{A}^*(\phi) \wedge \mathcal{A}^*(\psi)$$
 for any $\phi \in \Lambda^k(W^*), \psi \in \Lambda^l(W^*)$.

If $\mathcal{B}: W \to U$ is another linear map then $(\mathcal{B} \circ \mathcal{A})^* = \mathcal{A}^* \circ \mathcal{B}^*$.

Proof.

1. Take any k+l vectors $X_1,\ldots,X_{k+l}\in V$. Then by definition of the operator \mathcal{A}^* we have

$$\mathcal{A}^*(\phi \otimes \psi)(X_1, \dots, X_{k+l}) = \phi \otimes \psi(\mathcal{A}(X_1), \dots, \mathcal{A}(X_{n+k})) =$$

$$\phi(\mathcal{A}(X_1), \dots, \mathcal{A}(X_k)\psi(\mathcal{A}(X_{k+1}), \dots, \mathcal{A}(X_{n+k})) =$$

$$\mathcal{A}^*(\phi)(X_1, \dots, X_k)\mathcal{A}^*(\psi)(X_{k+1}, \dots, X_{k+n}) =$$

$$\mathcal{A}^*(\phi) \otimes \mathcal{A}^*(\psi)(X_1, \dots, X_{k+l}).$$

2. Given k vectors $X_1, \ldots, X_k \in V$ we get

$$\mathcal{A}^{*}(\phi^{\text{a}sym})(X_{1}, \dots, X_{k}) = \phi^{\text{a}sym}(\mathcal{A}(X_{1}), \dots, \mathcal{A}(X_{k})) = \sum_{(i_{1} \dots i_{k})} (-1)^{\text{inv}(i_{1}, \dots, i_{k})} \phi(\mathcal{A}(X_{i_{1}}), \dots, \mathcal{A}(X_{i_{k}})) = \sum_{(i_{1} \dots i_{k})} (-1)^{\text{inv}(i_{1}, \dots, i_{k})} \mathcal{A}^{*}(\phi)(X_{i_{1}}, \dots, X_{i_{k}}) = (\mathcal{A}^{*}(\phi))^{\text{a}sym}(X_{1}, \dots, X_{k}),$$

where the sum is taken over all permutations i_1, \ldots, i_k of indices $1, \ldots, k$. Similarly one proves that $\mathcal{A}^*(\phi^{\text{Sym}}) = (\mathcal{A}^*(\phi))^{\text{sym}}$.

3.
$$\mathcal{A}^*(\phi \wedge \psi) = \frac{1}{k!l!} \mathcal{A}^*((\phi \otimes \psi)^{asym}) = \frac{1}{k!l!} (\mathcal{A}^*(\phi \otimes \psi))^{asym} = \mathcal{A}^*(\phi) \wedge \mathcal{A}^*(\psi).$$

The last statement of Proposition 2.16 is straightforward and its proof is left to the reader.

Let us now discuss how to compute $\mathcal{A}^*(\phi)$ in coordinates. Let us fix bases v_1, \ldots, v_m and w_1, \ldots, w_n in spaces V and W. Let x_1, \ldots, x_m and y_1, \ldots, y_n be coordinates and

$$A = \begin{pmatrix} a_{11} & \dots & a_{1m} \\ \vdots & \vdots & \vdots \\ a_{n1} & \dots & a_{nm} \end{pmatrix}$$

be the matrix of a linear map $\mathcal{A}:V\to W$ in these bases. Note that the map \mathcal{A} in these coordinates is given by n linear coordinate functions:

$$y_1 = l_1(x_1, \dots, x_m) = a_{11}x_1 + a_{12}x_2 + \dots + a_{1m}x_m$$

$$y_2 = l_2(x_1, \dots, x_m) = a_{21}x_1 + a_{22}x_2 + \dots + a_{2m}x_m$$

$$\dots$$

$$y_n = l_n(x_1, \dots, x_k) = a_{n1}x_1 + a_{n2}x_2 + \dots + a_{nm}x_n$$

We have already computed in Section 1.3 that $\mathcal{A}^*(y_k) = l_k = \sum_{j=1}^m a_{kj} x_j$, $k = 1, \ldots, n$. Indeed, the coefficients of the function $l_k = \mathcal{A}^*(y_k)$ form the k-th column of the transpose matrix A^T . Hence, using Proposition 2.16 we compute:

$$\mathcal{A}^*(y_{i_1}\otimes\cdots\otimes y_{i_k})=l_{i_1}\otimes\cdots\otimes l_{i_k}$$

and

$$\mathcal{A}^*(y_{j_1}\wedge\cdots\wedge y_{j_k})=l_{j_1}\wedge\cdots\wedge l_{j_k}.$$

Consider now the case when V = W, n = m, and we use the same basis v_1, \ldots, v_n in the source and target spaces.

Proposition 2.17.

$$\mathcal{A}^*(x_1 \wedge \cdots \wedge x_n) = \det A x_1 \wedge \cdots \wedge x_n.$$

Note that the determinant det A is independent of the choice of the basis. Indeed, the matrix of a linear map changes to a similar matrix $C^{-1}AC$ in a different basis, and det $C^{-1}AC = \det A$. Hence, we can write det A instead of det A, i.e. attribute the determinant to the linear operator A rather than to its matrix A.

Proof. We have

$$\mathcal{A}^*(x_1 \wedge \dots \wedge x_n) = l_1 \wedge \dots \wedge l_n = \sum_{i_1=1}^n a_{1i_1} x_{i_1} \wedge \dots \wedge \sum_{i_n=1}^n a_{ni_n} x_{i_n} = \sum_{i_1,\dots,i_n=1}^n a_{1i_1} \dots a_{ni_n} x_{i_1} \wedge \dots \wedge x_{i_n}.$$

Note that the in the latter sum all terms with repeating indices vanish, and hence we can replace this sum by a sum over all permutations of indices $1, \ldots, n$. Thus, we can continue

$$\mathcal{A}^*(x_1 \wedge \dots \wedge x_n) = \sum_{i_1,\dots,i_n} a_{1i_1} \dots a_{ni_n} x_{i_1} \wedge \dots \wedge x_{i_n} =$$

$$\left(\sum_{i_1,\dots,i_n} (-1)^{\operatorname{inv}(i_1,\dots,i_n)} a_{1i_1} \dots a_{ni_n}\right) x_1 \wedge \dots \wedge x_n = \det A x_1 \wedge \dots \wedge x_n.$$

Exercise 2.18. Apply the equality

$$\mathcal{A}^*(x_1 \wedge \cdots \wedge x_k \wedge x_{k+1} \wedge \cdots \wedge x_n) = \mathcal{A}^*(x_1 \wedge \cdots \wedge x_k) \wedge \mathcal{A}^*(x_{k+1} \wedge \cdots \wedge x_n)$$

for a map $A : \mathbb{R}^n \to \mathbb{R}^n$ to deduce the formula for expansion of a determinant according to its first k rows:

$$\det A = \sum_{i_1 < \dots < i_k, \ j_1 < \dots < j_{n-k}; \ i_m \neq j_l} (-1)^{\operatorname{inv}(i_1, \dots, j_{n-k})} \begin{vmatrix} a_{1,i_1} & \dots & a_{1,i_k} \\ \vdots & \vdots & \vdots \\ a_{k,i_1} & \dots & a_{k,i_k} \end{vmatrix} \begin{vmatrix} a_{k+1,j_1} & \dots & a_{k+1,j_{n-k}} \\ \vdots & \vdots & \vdots \\ a_{n,j_1} & \dots & a_{n,j_{n-k}} \end{vmatrix}.$$

Chapter 3

Orientation and Volume

3.1 Orientation

We say that two bases v_1, \ldots, v_k and w_1, \ldots, w_k of a vector space V define the same orientation of V if the matrix of transition from one of these bases to the other has a positive determinant. Clearly, if we have 3 bases, and the first and the second define the same orientation, and the second and the third define the same orientation then the first and the third also define the same orientation. Thus, one can subdivide the set of all bases of V into the two classes. All bases in each of these classes define the same orientation; two bases chosen from different classes define opposite orientation of the space. To choose an orientation of the space simply means to choose one of these two classes of bases.

There is no way to say which orientation is "positive" or which is "negative"—it is a question of convention. For instance, the so-called counter-clockwise orientation of the plane depends from which side we look at the plane. The positive orientation of our physical 3-space is a physical, not mathematical, notion.

Suppose we are given two oriented spaces V, W of the same dimension. An invertible linear map (= isomorphism) $\mathcal{A}: V \to W$ is called *orientation preserving* if it maps a basis which defines the given orientation of V to a basis which defines the given orientation of W.

Any non-zero exterior n-form η on V induces an orientation of the space V. Indeed, the preferred set of bases is characterized by the property $\eta(v_1, \ldots, v_n) > 0$.

3.2 Orthogonal transformations

Let V be a Euclidean vector space. Recall that a linear operator $\mathcal{U}: V \to V$ is called *orthogonal* if it preserves the scalar product, i.e. if

$$\langle \mathcal{U}(X), \mathcal{U}(Y) \rangle = \langle X, Y \rangle,$$
 (3.2.1)

for any vectors $X, Y \in V$. Recall that we have

$$\langle \mathcal{U}(X), \mathcal{U}(Y) \rangle = \langle X, \mathcal{U}^* \left(\mathcal{U}(Y) \right) \rangle,$$

where $\mathcal{U}^{\star}: V \to V$ is the adjoint operator to \mathcal{U} , see Section 1.3 above.

Hence, the orthogonality of an operator \mathcal{U} is equivalent to the identity $\mathcal{U}^* \circ \mathcal{U} = \mathrm{Id}$, or $\mathcal{U}^* = \mathcal{U}^{-1}$. Here we denoted by Id the identity operator, i.e. $\mathrm{Id}(X) = X$ for any $X \in V$.

Let us recall, see Exercise 1.10, that the adjoint operator U^* is related to the dual operator $U^*: V^* \to V^*$ by the formula

$$U^{\star} = \mathcal{D}^{-1} \circ U^* \circ \mathcal{D}.$$

Hence, for an orthogonal operator \mathcal{U} , we have $\mathcal{D}^{-1} \circ U^* \circ \mathcal{D} = \mathcal{U}^{-1}$, i.e.

$$\mathcal{U}^* \circ \mathcal{D} = \mathcal{D} \circ \mathcal{U}^{-1}. \tag{3.2.2}$$

Let v_1, \ldots, v_n be an orthonormal basis in V and U be the matrix of \mathcal{U} in this basis. The matrix of the adjoint operator in an orthonormal basis is the transpose of the matrix of this operator. Hence, the equation $\mathcal{U}^* \circ \mathcal{U} = \text{Id}$ translates into the equation $U^T U = E$, or equivalently $U U^T = E$, or $U^{-1} = U^T$ for its matrix. Matrices, which satisfy this equation are called *orthogonal*. If we write

$$U = \begin{pmatrix} u_{11} & \dots & u_{1n} \\ \dots & \dots & \dots \\ u_{n1} & \dots & u_{nn} \end{pmatrix} ,$$

then the equation $U^TU = E$ can be rewritten as

$$\sum_{i} u_{ki} u_{ji} = \begin{cases} 1, & \text{if } k = j, \\ 0, & \text{if } k \neq j, \end{cases}$$

Similarly, the equation $UU^T = E$ can be rewritten as

$$\sum_{i} u_{ik} u_{ij} = \begin{cases} 1, & \text{if } k = j, \\ 0, & \text{if } k \neq j, \end{cases}$$

The above identities mean that columns (and rows) of an orthogonal matrix U form an orthonormal basis of \mathbb{R}^n with respect to the dot-product.

In particular, we have

$$1 = \det(U^T U) = \det(U^T) \det U = (\det U)^2,$$

and hence $\det U = \pm 1$. In other words, the determinant of any orthogonal matrix is equal ± 1 . We can also say that the determinant of an orthogonal *operator* is equal to ± 1 because the determinant of the matrix of an operator is independent of the choice of a basis. Orthogonal transformations with $\det = 1$ preserve the orientation of the space, while those with $\det = -1$ reverse it.

Composition of two orthogonal transformations, or the inverse of an orthogonal transformation is again an orthogonal transformation. The set of all orthogonal transformations of an n-dimensional Euclidean space is denoted by O(n). Orientation preserving orthogonal transformations sometimes called special, and the set of special orthogonal transformations is denoted by SO(n). For instance O(1) consists of two elements and SO(1) of one: $O(1) = \{1, -1\}$, $SO(1) = \{1\}$. SO(2) consists of rotations of the plane, while O(2) consists of rotations and reflections with respect to lines.

3.3 Determinant and Volume

We begin by recalling some facts from Linear Algebra. Let V be an n-dimensional Euclidean space with an inner product $\langle \ , \ \rangle$. Given a linear subspace $L \subset V$ and a point $x \in V$, the projection proj L(x) is a vector $y \in L$ which is uniquely characterized by the property $x - y \perp L$, i.e. $\langle x - y, z \rangle = 0$ for any $z \in L$. The length $||x - \operatorname{proj}_L(x)||$ is called the distance from x to L; we denote it by $\operatorname{dist}(x, L)$.

Let $U_1, \ldots, U_k \in V$ be linearly independent vectors. The k-dimensional parallelepiped spanned by vectors U_1, \ldots, U_k is, by definition, the set

$$P(U_1, \dots, U_k) = \left\{ \sum_{1}^{k} \lambda_j U_j; \ 0 \le \lambda_1, \dots, \lambda_k \le 1 \right\} \subset \text{Span}(U_1, \dots, U_k).$$

Given a k-dimensional parallelepiped $P = P(U_1, ..., U_k)$ we will define its k-dimensional volume by the formula

$$Vol P = ||U_1|| dist(U_2, Span(U_1)) dist(U_3, Span(U_1, U_2)) \dots dist(U_k, Span(U_1, \dots, U_{k-1})).$$
(3.3.1)

Of course we can write $dist(U_1, 0)$ instead of $||U_1||$. This definition agrees with the definition of the area of a parallelogram, or the volume of a 3-dimensional parallelepiped in the elementary geometry.

Proposition 3.1. Let v_1, \ldots, v_n be an orthonormal basis in V. Given n vectors U_1, \ldots, U_n let us denote by U the matrix whose columns are coordinates of these vectors in the basis v_1, \ldots, v_n :

$$U := \begin{pmatrix} u_{11} & \dots & u_{1n} \\ & \vdots & \\ u_{n1} & \dots & u_{nn} \end{pmatrix}$$

Then

$$\operatorname{Vol} P(U_1, \dots, U_n) = |\det U|.$$

Proof. If the vectors U_1, \ldots, U_n are linearly dependent then $\operatorname{Vol} P(U_1, \ldots, U_n) = \det U = 0$. Suppose now that the vectors U_1, \ldots, U_n are linearly independent, i.e. form a basis. Consider first the case where this basis is orthonormal. Then the matrix U is orthogonal. i.e. $UU^T = E$, and hence $\det U = \pm 1$. But in this case $\operatorname{Vol} P(U_1, \ldots, U_n) = 1$, and hence $\operatorname{Vol} P(U_1, \ldots, U_n) = |\det U|$.

Now let the basis U_1, \ldots, U_n be arbitrary. Let us apply to it the Gram-Schmidt orthonormalization process. Recall that this process consists of the following steps. First, we normalize the vector U_1 , then subtract from U_2 its projection to $\mathrm{Span}(U_1)$, Next, we normalize the new vector U_2 , then subtract from U_3 its projection to $\mathrm{Span}(U_1, U_2)$, and so on. At the end of this process we obtain an orthonormal basis. It remains to notice that each of these steps affected $\mathrm{Vol}\,P(U_1,\ldots,U_n)$ and $|\det U|$ in a similar way. Indeed, when we multiplied the vectors by a positive number, both the volume and the determinant were multiplied by the same number. When we subtracted from a vector U_k its projection to $\mathrm{Span}(U_1,\ldots,U_{k-1})$, this affected neither the volume nor the determinant.

Corollary 3.2. 1. Let x_1, \ldots, x_n be a Cartesian coordinate system. 1 Then

$$\operatorname{Vol} P(U_1, \dots, U_n) = |x_1 \wedge \dots x_n(U_1, \dots, U_n)|.$$

¹i.e. a coordinate system with respect to an orthonormal basis

2. Let $A: V \to V$ be a linear map. Then

$$\operatorname{Vol} P(\mathcal{A}(U_1), \dots, \mathcal{A}(U_n)) = |\det \mathcal{A}| \operatorname{Vol} P(U_1, \dots, U_n).$$

Proof.

- 1. According to 2.13, $x_1 \wedge \ldots x_n(U_1, \ldots, U_n) = \det U$.
- 2. $x_1 \wedge \ldots x_n(\mathcal{A}(U_1), \ldots, \mathcal{A}(U_n)) = \mathcal{A}^*(x_1 \wedge \cdots \wedge x_n)(U_1, \ldots, U_n) = \det \mathcal{A}x_1 \wedge \ldots x_n(U_1, \ldots, U_n)$. In view of Proposition 3.1 and the first part of Corollary 3.2 the value

$$x_1 \wedge \dots x_n(U_1, \dots, U_n) = \det U$$

is called sometimes the *signed volume* of the parallelepiped $P(U_1, \ldots, U_n)$. It is positive when the basis U_1, \ldots, U_n defines the given orientation of the space V, and it is negative otherwise.

Note that $x_1 \wedge \ldots x_k(U_1, \ldots, U_k)$ for $0 \leq k \leq n$ is the signed k-dimensional volume of the orthogonal projection of the parallelepiped $P(U_1, \ldots, U_k)$ to the coordinate subspace $\{x_{k+1} = \cdots = x_n = 0\}$.

For instance, let ω be the 2-form $x_1 \wedge x_2 + x_3 \wedge x_4$ on \mathbb{R}^4 . Then for any two vectors $U_1, U_2 \in \mathbb{R}^4$ the value $\omega(U_1, U_2)$ is the sum of signed areas of projections of the parallelogram $P(U_1, U_2)$ to the coordinate planes spanned by the two first and two last basic vectors.

3.4 Volume and Gram matrix

In this section we will compute the $Vol_k P(v_1, \ldots, v_k)$ in the case when the number k of vectors is less than the dimension n of the space.

Let V be an Euclidean space. Given vectors $v_1, \ldots, v_k \in V$ we can form a $k \times k$ -matrix

$$G(v_1, \dots, v_k) = \begin{pmatrix} \langle v_1, v_1 \rangle & \dots & \langle v_1, v_k \rangle \\ \dots & \dots & \dots \\ \langle v_k, v_1 \rangle & \dots & \langle v_k, v_k \rangle \end{pmatrix},$$
(3.4.1)

which is called the *Gram matrix* of vectors v_1, \ldots, v_k .

Suppose we are given Cartesian coordinate system in V and let us form a matrix C whose columns are coordinates of vectors v_1, \ldots, v_k . Thus the matrix C has n rows and k columns. Then

$$G(v_1, \dots, v_k) = C^T C,$$

because in Cartesian coordinates the scalar product looks like the dot-product.

We also point out that if k = n and vectors v_1, \ldots, v_n form a basis of V, then $G(v_1, \ldots, v_k)$ is just the matrix of the bilinear function $\langle X, Y \rangle$ in the basis v_1, \ldots, v_n . It is important to point out that while the matrix C depends on the choice of the basis, the matrix G does not.

Proposition 3.3. Given any k vectors v_1, \ldots, v_k in an Euclidean space V the volume $Vol_k P(v_1, \ldots, v_k)$ can be computed by the formula

$$\operatorname{Vol}_k P(v_1, \dots, v_k)^2 = \det G(v_1, \dots, v_k) = \det C^T C,$$
 (3.4.2)

where $G(v_1, ..., v_k)$ is the Gram matrix and C is the matrix whose columns are coordinates of vectors $v_1, ..., v_k$ in some orthonormal basis.

Proof. Suppose first that k = n. Then according to Proposition 3.1 we have $Vol_k P(v_1, \ldots, v_k) = |\det C|$. But $\det C^T C = \det C^2$, and the claim follows.

Let us denote vectors of our orthonormal basis by w_1, \ldots, w_n . Consider now the case when

$$\operatorname{Span}(v_1, \dots, v_k) \subset \operatorname{Span}(w_1, \dots, w_k). \tag{3.4.3}$$

In this case the elements in the j-th row of the matrix C are zero if j > k. Hence, if we denote by \widetilde{C} the square $k \times k$ matrix formed by the first k rows of the matrix C, then $C^TC = \widetilde{C}^T\widetilde{C}$ and thus $\det C^TC = \det \widetilde{C}^T\widetilde{C}$. But $\det \widetilde{C}^T\widetilde{C} = \operatorname{Vol}_k P(v_1, \dots, v_k)$ in view of our above argument in the equi-dimensional case applied to the subspace $\operatorname{Span}(w_1, \dots, w_k) \subset V$, and hence

$$\operatorname{Vol}_k^2 P(v_1, \dots, v_k) = \det C^T C = \det G(v_1, \dots, v_k).$$

But neither $Vol_k P(v_1, ..., v_k)$, nor the Gram matrix $G(v_1, ..., v_k)$ depends on the choice of an orthonormal basis. On the other hand, using Gram-Schmidt process one can always find an orthonormal basis which satisfies condition (3.4.3).

Remark 3.4. Note that $\det G(v_1, \ldots, v_k) \geq 0$ and $\det G(v_1, \ldots, v_k) = 0$ if an and only if the vectors v_1, \ldots, v_k are linearly dependent.

Chapter 4

Dualities

4.1 Duality between k-forms and (n-k)-forms on a n-dimensional Euclidean space V

Let V be an n-dimensional vector space. As we have seen above, the space $\Lambda^k(V^*)$ of k-forms, and the space $\Lambda^{n-k}(V^*)$ of (n-k)-forms have the same dimension $\frac{n!}{k!(n-k)!}$; these spaces are therefore isomorphic. Suppose that V is an *oriented Euclidean* space, i.e. it is supplied with an orientation and an inner product $\langle \ , \ \rangle$. It turns out that in this case there is a canonical way to establish this isomorphism which will be denoted by

$$\star: \Lambda^k(V^*) \to \Lambda^{n-k}(V^*)$$
.

Definition 4.1. Let α be a k-form. Then given any vectors U_1, \ldots, U_{n-k} , the value $\star \alpha(U_1, \ldots, U_{n-k})$ can be computed as follows. If U_1, \ldots, U_{n-k} are linearly dependent then $\star \alpha(U_1, \ldots, U_{n-k}) = 0$. Otherwise, let S^{\perp} denote the orthogonal complement to the space $S = \operatorname{Span}(U_1, \ldots, U_{n-k})$. Choose a basis Z_1, \ldots, Z_k of S^{\perp} such that

$$\operatorname{Vol}_k(Z_1,\ldots,Z_k) = \operatorname{Vol}_{n-k}(U_1,\ldots,U_{n-k})$$

and the basis $Z_1, \ldots, Z_k, U_1, \ldots, U_{n-k}$ defines the given orientation of the space V. Then

$$\star \alpha(U_1, \dots, U_{n-k}) = \alpha(Z_1, \dots, Z_k). \tag{4.1.1}$$

Let us first show that

Lemma 4.2. $\star \alpha$ is a (n-k)-form, i.e. $\star \alpha$ is skew-symmetric and multilinear.

Proof. To verify that $\star \alpha$ is skew-symmetric we note that for any $1 \leq i < j \leq n-q$ the bases

$$Z_1, Z_2, \ldots, Z_k, U_1, \ldots, U_i, \ldots, U_i, \ldots, U_{n-k}$$

and

$$-Z_1, Z_2, \ldots, Z_k, U_1, \ldots, U_i, \ldots, U_i, \ldots, U_{n-k}$$

define the same orientation of the space V, and hence

$$\star \alpha(U_1, \dots, U_j, \dots, U_i, \dots, U_{n-k}) = \alpha(-Z_1, Z_2, \dots, Z_k)$$
$$= -\alpha(Z_1, Z_2, \dots, Z_k) = -\star \alpha(U_1, \dots, U_i, \dots, U_j, \dots, U_{n-k}).$$

Hence, in order to check the *multi*-linearity it is sufficient to prove the linearity of α with respect to the first argument only. It is also clear that

$$\star \alpha(\lambda U_1, \dots, U_{n-k}) = \star \lambda \alpha(U_1, \dots, U_{n-k}). \tag{4.1.2}$$

Indeed, multiplication by $\lambda \neq 0$ does not change the span of the vectors U_1, \ldots, U_{n-q} , and hence if $\star \alpha(U_1, \ldots, U_{n-k}) = \alpha(Z_1, \ldots, Z_k)$ then $\star \alpha(\lambda U_1, \ldots, U_{n-k}) = \alpha(\lambda Z_1, \ldots, Z_k) = \lambda \alpha(Z_1, \ldots, Z_k)$.

Thus it remains to check that

$$\star \alpha(U_1 + \widetilde{U}_1, U_2, \dots, U_{n-k}) = \star \alpha(U_1, U_2, \dots, U_{n-k}) + \star \alpha(\widetilde{U}_1, U_2, \dots, U_{n-k})).$$

Let us denote $L := \operatorname{Span}(U_2, \dots, U_{n-k})$ and observe that $\operatorname{proj}_L(U_1 + \widetilde{U}_1) = \operatorname{proj}_L(U_1) + \operatorname{proj}_L(\widetilde{U}_1)$. Denote $N := U_1 - \operatorname{proj}_L(U_1)$ and $\widetilde{N} := \widetilde{U}_1 - \operatorname{proj}_L(\widetilde{U}_1)$. The vectors N and \widetilde{N} are normal components of U_1 and \widetilde{U}_1 with respect to the subspace L, and the vector $N + \widetilde{N}$ is the normal component of $U_1 + \widetilde{U}_1$ with respect to L. Hence, we have

$$\star \alpha(U_1, \dots, U_{n-k}) = \star \alpha(N, \dots, U_{n-k}), \quad \star \alpha(\widetilde{U}_1, \dots, U_{n-k}) = \star \alpha(\widetilde{N}, \dots, U_{n-k}),$$

and

$$\star \alpha(U_1 + \widetilde{U}_1, \dots, U_{n-k}) = \star \alpha(N + \widetilde{N}, \dots, U_{n-k}).$$

Indeed, in each of these three cases,

- vectors on both side of the equality span the same space;
- the parallelepiped which they generate have the same volume, and
- the orientation which these vectors define together with a basis of the complementary space remains unchanged.

Hence, it is sufficient to prove that

$$\star \alpha(N + \widetilde{N}, U_2, \dots, U_{n-k}) = \star \alpha(N, U_2, \dots, U_{n-k}) + \star \alpha(\widetilde{N}, U_2, \dots, U_{n-k}). \tag{4.1.3}$$

If the vectors N and \widetilde{N} are linearly dependent, i.e. one of them is a multiple of the other, then (4.1.3) follows from (4.1.2).

Suppose now that N and \widetilde{N} are linearly independent. Let L^{\perp} denote the orthogonal complement of $L = \operatorname{Span}(U_2, \ldots, U_{n-k})$. Then $\dim L^{\perp} = k+1$ and we have $N, \widetilde{N} \in L^{\perp}$. Let us denote by M the plane in L^{\perp} spanned by the vectors N and \widetilde{N} , and by M^{\perp} its orthogonal complement in L^{\perp} . Note that $\dim M^{\perp} = k-1$.

Choose any orientation of M so that we can talk about counter-clockwise rotation of this plane. Let $Y, \widetilde{Y} \in M$ be vectors obtained by rotating N and \widetilde{N} in M counter-clockwise by the angle $\frac{\pi}{2}$. Then $Y + \widetilde{Y}$ can be obtained by rotating $N + \widetilde{N}$ in M counter-clockwise by the same angle $\frac{\pi}{2}$. Let us choose in M^{\perp} a basis Z_2, \ldots, Z_k such that

$$Vol_{k-1}P(Z_2,...,Z_k) = Vol_{n-k-1}P(U_2,...,U_{n-k}).$$

Note that the orthogonal complements to $\operatorname{Span}(N, U_2, \dots, U_{n-k})$, $\operatorname{Span}(\widetilde{N}, U_2, \dots, U_{n-k})$, and to $\operatorname{Span}(N+\widetilde{N}, U_2, \dots, U_{n-k})$ in V coincide, respectively, with the orthogonal complements to the the vectors N, \widetilde{N} and to $N+\widetilde{N}$ in L^{\perp} . In other words, we have

$$(\operatorname{Span}(N, U_2, \dots, U_{n-k}))_V^{\perp} = \operatorname{Span}(Y, Z_2, \dots, Z_k),$$

$$\left(\operatorname{Span}(\widetilde{N}, U_2, \dots, U_{n-k})\right)_V^{\perp} = \operatorname{Span}(\widetilde{Y}, Z_2, \dots, Z_k) \text{ and }$$

$$\left(\operatorname{Span}(N + \widetilde{N}, U_2, \dots, U_{n-k})\right)_V^{\perp} = \operatorname{Span}(Y + \widetilde{Y}, Z_2, \dots, Z_k).$$

Next, we observe that

$$Vol_{n-k}P(N, U_2, \dots, U_{n-k}) = Vol_kP(Y, Z_2, \dots, Z_k),$$

$$\operatorname{Vol}_{n-k}P(\widetilde{N},U_2,\ldots,U_{n-k}) = \operatorname{Vol}_kP(\widetilde{Y},Z_2,\ldots,Z_k)$$
 and
$$\operatorname{Vol}_{n-k}P(N+\widetilde{N},U_2,\ldots,U_{n-k}) = \operatorname{Vol}_kP(Y+\widetilde{Y},Z_2,\ldots,Z_k).$$

Consider the following 3 bases of V:

$$Y, Z_2, \dots, Z_k, N, U_2, \dots, U_{n-k},$$

$$\widetilde{Y}, Z_2, \dots, Z_k, \widetilde{N}, U_2, \dots, U_{n-k},$$

$$Y + \widetilde{Y}, Z_2, \dots, Z_k, N + \widetilde{N}, U_2, \dots, U_{n-k},$$

and observe that all three of them define the same a priori given orientation of V. Thus, by definition of the operator \star we have:

$$\star \alpha(N + \widetilde{N}, U_2, \dots, U_{n-k}) = \alpha(Y + \widetilde{Y}, Z_2, \dots, Z_k)$$

$$= \alpha(Y, Z_2, \dots, Z_k) + \alpha(\widetilde{Y}, Z_2, \dots, Z_k) = \star \alpha(N, U_2, \dots, U_{n-k}) + \star \alpha(\widetilde{N}, U_2, \dots, U_{n-k}).$$

This completes the proof that $\star \alpha$ is an (n-k)-form.

Thus the map $\alpha \mapsto \star \alpha$ defines a map $\star : \Lambda^k(V^*) \to \Lambda^{n-k}(V^*)$. Clearly, this map is linear. In order to check that \star is an isomorphism let us choose an orthonormal basis in V and consider the coordinates $x_1, \ldots, x_n \in V^*$ corresponding to that basis.

Let us recall that the forms $x_{i_1} \wedge x_{i_2} \wedge \cdots \wedge x_{i_k}$, $1 \leq i_1 < i_2 < \cdots < i_k \leq n$, form a basis of the space $\Lambda^k(V^*)$.

Lemma 4.3.

$$\star x_{i_1} \wedge x_{i_2} \wedge \dots \wedge x_{i_k} = (-1)^{\text{inv}(i_1, \dots, i_k, j_1, \dots, j_{n-k})} x_{j_1} \wedge x_{j_2} \wedge \dots \wedge x_{j_{n-k}}, \qquad (4.1.4)$$

where $j_1 < \cdots < j_{n-k}$ is the set of indices, complementary to i_1, \ldots, i_k . In other words, $i_1, \ldots, i_k, j_1, \ldots, j_{n-k}$ is a permutation of indices $1, \ldots, n$.

Proof. Evaluating $\star(x_{i_1} \wedge \cdots \wedge x_{i_k})$ on basic vectors $v_{j_1}, \ldots, v_{j_{n-k}}, 1 \leq j_1 < \cdots < j_{n-q} \leq n$, we get 0 unless all the indices j_1, \ldots, j_{n-k} are all different from i_1, \ldots, i_k , while in the latter case we get

$$\star (x_{i_1} \wedge \dots \wedge x_{i_k})(v_{j_1}, \dots, v_{j_{n-k}}) = (-1)^{\operatorname{inv}(i_1, \dots, i_k, j_1, \dots, j_{n-k})}.$$

Hence,

$$\star x_{i_1} \wedge x_{i_2} \wedge \dots \wedge x_{i_k} = (-1)^{\operatorname{inv}(i_1,\dots,i_k,j_1,\dots,j_{n-k})} x_{j_1} \wedge x_{j_2} \wedge \dots \wedge x_{j_{n-k}}.$$

Thus \star establishes a 1 to 1 correspondence between the bases of the spaces $\Lambda^k(V^*)$ and the space $\Lambda^{n-k}(V^*)$, and hence it is an isomorphism. Note that by linearity for any form $\alpha = \sum_{1 \leq i_1 < \dots < i_k \leq n} a_{i_1 \dots i_k} x_{i_1} \wedge \dots \wedge x_{i_k}$ we have

$$\star \alpha = \sum_{1 \le i_1 < \dots < i_k \le n} a_{i_1 \dots i_k} \star (x_{i_1} \wedge \dots \wedge x_{i_k}).$$

Examples.

- 1. $\star C = Cx_1 \wedge \cdots \wedge x_n$; in other words the isomorphism \star acts on constants (= 0-forms) by multiplying them by the volume form.
- 2. In \mathbb{R}^3 we have

$$\star x_1 = x_2 \wedge x_3, \star x_2 = -x_1 \wedge x_3 = x_3 \wedge x_1, \star x_3 = x_1 \wedge x_2,$$

$$\star (x_1 \wedge x_2) = x_3, \star (x_3 \wedge x_1) = x_2, \star (x_2 \wedge x_3) = x_1.$$

3. More generally, given a 1-form $l = a_1x_1 + \cdots + a_nx_n$ we have

$$\star l = a_1 x_2 \wedge \cdots \wedge x_n - a_2 x_1 \wedge x_3 \wedge \cdots \wedge x_n + \cdots + (-1)^{n-1} a_n x_1 \wedge \cdots \wedge x_{n-1}.$$

In particular for n = 3 we have

$$\star (a_1x_1 + a_2x_2 + a_3x_3) = a_1x_2 \wedge x_3 + a_2x_3 \wedge x_1 + a_3x_1 \wedge x_2.$$

Proposition 4.4.

$$\star^2 = (-1)^{k(n-k)} \mathrm{Id}, \quad i.e. \quad \star \left(\star \omega\right) = (-1)^{k(n-k)} \omega \quad \textit{for any k-form} \quad \omega \,.$$

In particular, if dimension $n = \dim V$ is odd then $*^2 = \operatorname{Id}$. If n is even and ω is a k-form then $\star(\star\omega) = \omega$ if k is even, and $\star(\star\omega) = -\omega$ if k is odd.

Proof. It is sufficient to verify the equality

$$\star(\star\omega) = (-1)^{k(n-k)}\omega$$

for the case when ω is a basic form, i.e.

$$\omega = x_{i_1} \wedge \cdots \wedge x_{i_k}, \quad 1 \leq i_1 < \cdots < i_k \leq n.$$

We have

$$\star (x_{i_1} \wedge \dots \wedge x_{i_k}) = (-1)^{\operatorname{inv}(i_1, \dots, i_k, j_1, \dots, j_{n-k})} x_{j_1} \wedge x_{j_2} \wedge \dots \wedge x_{j_{n-k}}$$

and

$$\star (x_{i_1} \wedge x_{i_2} \wedge \cdots \wedge x_{i_{n-k}}) = (-1)^{\text{inv}(j_1, \dots, j_{n-k}, i_1, \dots, i_k)} * x_{i_1} \wedge \cdots \wedge x_{i_k}.$$

But the permutations $i_1 \dots i_k j_1 \dots j_{n-k}$ and $j_1 \dots j_{n-k} i_1 \dots i_k$ differ by k(n-k) transpositions of pairs of its elements. Hence, we get

$$(-1)^{\operatorname{inv}(i_1,\dots,i_k,j_1,\dots,j_{n-k})} = (-1)^{k(n-k)}(-1)^{\operatorname{inv}(j_1,\dots,j_{n-k},i_1,\dots,i_k)},$$

and, therefore,

$$\star (\star (x_{i_1} \wedge \cdots \wedge x_{i_k})) = \star ((-1)^{\operatorname{inv}(i_1, \dots, i_k, j_1, \dots, j_{n-k})} x_{j_1} \wedge x_{j_2} \wedge \cdots \wedge x_{j_{n-k}})$$

$$= (-1)^{\operatorname{inv}(i_1, \dots, i_k, j_1, \dots, j_{n-k})} \star (x_{j_1} \wedge x_{j_2} \wedge \cdots \wedge x_{j_{n-k}})$$

$$= (-1)^{\operatorname{inv}(i_1, \dots, i_k, j_1, \dots, j_{n-k}) + \operatorname{inv}(j_1, \dots, j_{n-k}, i_1, \dots, i_k)} x_{i_1} \wedge \cdots \wedge x_{i_k}$$

$$= (-1)^{k(n-k)} x_{i_1} \wedge \cdots \wedge x_{i_k}.$$

Exercise 4.5. (a) For any special orthogonal operator A the operators A^* and \star commute, i.e.

$$\mathcal{A}^* \circ \star = \star \circ \mathcal{A}^*$$
.

(b) Let A be an orthogonal matrix of order n with det A = 1. Prove that for any $k \in \{1, ..., n\}$ the absolute value of each k-minor M of A is equal to the absolute value of its complementary minor of order (n - k). (Hint: Apply (a) to the form $x_{i_1} \wedge \cdots \wedge x_{i_k}$).

(c) Let V be an oriented 3-dimensional Euclidean space. Prove that for any two vectors $X, Y \in V$, their cross-product can be written in the form

$$X \times Y = \mathcal{D}^{-1} \left(\star (\mathcal{D}(X) \wedge \mathcal{D}(Y)) \right).$$

4.2 Euclidean structure on the space of exterior forms

Suppose that the space V is oriented and Euclidean, i.e. it is endowed with an inner product $\langle \, , \, \rangle$ and an orientation.

Given two forms $\alpha, \beta \in \Lambda^k(V^*)$, k = 0, ..., n, let us define

$$\langle \langle \alpha, \beta \rangle \rangle = \star (\alpha \wedge \star \beta).$$

Note that $\alpha \wedge \star \beta$ is an *n*-form for every k, and hence, $\langle \langle \alpha, \beta \rangle \rangle$ is a 0-form, i.e. a real number.

Proposition 4.6. 1. The operation $\langle \langle , \rangle \rangle$ defines an inner product on $\Lambda^k(V^*)$ for each $k = 0, \ldots, n$.

2. If $A: V \to V$ is a special orthogonal operator then the operator $A^*: \Lambda^k(V^*) \to \Lambda^k(V^*)$ is orthogonal with respect to the inner product $\langle \langle , \rangle \rangle$.

Proof. 1. We need to check that $\langle \langle \alpha, \beta \rangle \rangle$ is a symmetric bilinear function on $\Lambda^k(V^*)$ and $\langle \langle \alpha, \alpha \rangle \rangle > 0$ unless $\alpha = 0$. Bilinearity is straightforward. Hence, it is sufficient to verify the remaining properties for basic vectors $\alpha = x_{i_1} \wedge \cdots \wedge x_{i_k}, \beta = x_{j_1} \wedge \cdots \wedge x_{j_k}$, where $1 \leq i_1 < \cdots < i_k \leq n, 1 \leq j_1 < \cdots < j_k \leq n$. Here (x_1, \ldots, x_n) is any Cartersian coordinates in V which define its given orientation.

Note that $\langle \langle \alpha, \beta \rangle \rangle = 0 = \langle \langle \beta, \alpha \rangle \rangle$ unless $i_m = j_m$ for all $m = 1, \dots, k$, and in the latter case we have $\alpha = \beta$. Furthermore, we have

$$\langle \langle \alpha, \alpha \rangle \rangle = \star (\alpha \wedge \star \alpha) = \star (x_1 \wedge \cdots \wedge x_n) = 1 > 0.$$

2. The inner product $\langle \langle , \rangle \rangle$ is defined only in terms of the Euclidean structure and the orientation of V. Hence, for any special orthogonal operator \mathcal{A} (which preserves these structures) the induced operator $\mathcal{A}^* : \Lambda^k(V^*) \to \Lambda^k(V^*)$ preserves the inner product $\langle \langle , \rangle \rangle$.

Note that we also proved that the basis of k-forms $x_{i_1} \wedge \cdots \wedge x_{i_k}$, $1 \leq i_1 < \cdots < i_k \leq n$, is orthonormal with respect to the scalar product $\langle \langle , \rangle \rangle$. Hence, we get

Corollary 4.7. Suppose that a k-form α can be written in Cartesian coordinates as

$$\alpha = \sum_{1 \le i_1 < \dots < i_k \le n} a_{i_1 \dots i_k} x_{i_1} \wedge \dots \wedge x_{i_k}.$$

Then

$$||\alpha||^2 = \langle \langle \alpha, \alpha \rangle \rangle = \sum_{1 \le i_1 < \dots < i_k \le n} a_{i_1 \dots i_k}^2.$$

Corollary 4.8. For any two exterior k-forms we have

$$\star \alpha \wedge \beta = (-1)^{k(n-k)} \alpha \wedge \star \beta.$$

Exercise 4.9. 1. Show that for any k-forms we have

$$\langle \langle \alpha, \beta \rangle \rangle = \langle \langle \star \alpha, \star \beta \rangle \rangle.$$

2. Show that if α, β are 1-forms on an Euclidean space V. Then

$$\langle \langle \alpha, \beta \rangle \rangle = \langle \mathcal{D}^{-1}(\alpha), \mathcal{D}^{-1}(\beta) \rangle,$$

 $i.e \ the \ scalar \ product \ \langle \langle \ , \ \rangle \rangle \ \ on \ V^* \ \ is \ the \ push-forward \ by \ \mathcal{D} \ \ of \ the \ scalar \ product \ \langle \ , \ \rangle \ \ on \ V.$

Corollary 4.10. Let V be a Euclidean n-dimensional space. Choose an orthonormal basis e_1, \ldots, e_n in V. Then for any vectors $Z_1 = (z_{11}, \ldots, z_{n1}), \ldots, Z_k = (z_{1k}, \ldots, z_{nk}) \in V$ we have

$$(\operatorname{Vol}_k P(Z_1, \dots, Z_k))^2 = \sum_{1 \le i_1 < \dots < i_k \le n} Z_{i_1, \dots, i_k}^2,$$
(4.2.1)

where

$$Z_{i_1,\ldots,i_k} = \begin{vmatrix} z_{i_11} & \ldots & z_{i_1k} \\ \ldots & \ldots & \ldots \\ z_{i_k1} & \ldots & z_{i_kk} \end{vmatrix}.$$

Proof. Consider linear functions $l_j = \mathcal{D}(Z_j) = \sum_{i=1}^n z_{ij} x_i \in V^*, j = 1, \dots, k$. Then

$$l_{1} \wedge \dots \wedge l_{k} = \sum_{i_{1}=1}^{n} z_{i_{1}j} x_{i_{1}} \wedge \dots \wedge \sum_{i_{k}=1}^{n} z_{i_{k}j} x_{i_{k}} = \sum_{i_{1},\dots,i_{k}} z_{i_{1}} \dots z_{i_{k}} x_{i_{1}} \wedge x_{i_{k}} = \sum_{1 \leq i_{1} < \dots < i_{k} \leq n} Z_{i_{1},\dots,i_{k}} x_{i_{1}} \wedge \dots x_{i_{k}}.$$

$$(4.2.2)$$

In particular, if one has $Z_1, \ldots, Z_k \in \operatorname{Span}(e_1, \ldots, e_k)$ then $Z_{1\ldots k} = \operatorname{Vol} P(Z_1, \ldots, Z_k)$ and hence

$$l_1 \wedge \cdots \wedge l_k = Z_{1\dots k} x_1 \wedge \cdots \wedge x_k = \operatorname{Vol} P(Z_1, \dots, Z_k) x_1 \wedge \cdots \wedge x_k,$$

which yields the claim in this case.

In the general case, according to Proposition 4.7 we have

$$|| l_1 \wedge \dots \wedge l_k ||^2 = \sum_{1 \le i_1 < \dots < i_k \le n} Z_{i_1, \dots, i_k}^2,$$
 (4.2.3)

which coincides with the right-hand side of (4.2.1). Thus it remains to check to that

$$|| l_1 \wedge \cdots \wedge l_k || = \operatorname{Vol}_k P(Z_1, \dots, Z_k).$$

Given any orthogonal transformation $\mathcal{A}: V \to V$ we have, according to Proposition 4.6, the equality

$$||l_1 \wedge \cdots \wedge l_k|| = ||\mathcal{A}^* l_1 \wedge \cdots \wedge \mathcal{A}^* l_k||.$$
 (4.2.4)

We also note that any orthogonal transformation $\mathcal{B}: V \to V$ preserves k-dimensional volume of all k-dimensional parallelepipeds:

$$|\operatorname{Vol}_k P(Z_1, \dots, Z_k)| = |\operatorname{Vol}_k P(\mathcal{B}(Z_1), \dots, \mathcal{B}(Z_k))|. \tag{4.2.5}$$

On the other hand, there exists an orthogonal transformation $\mathcal{A}: V \to V$ such that $\mathcal{A}^{-1}(Z_1), \dots, \mathcal{A}^{-1}(Z_k) \in \text{Span}(e_1, \dots, e_k)$. Denote $\widetilde{Z}_j := \mathcal{A}^{-1}(Z_j), j = 1, \dots, k$. Then, according to (3.2.2) we have

$$\widetilde{l}_j := \mathcal{D}(\widetilde{Z}_j) = \mathcal{D}(\mathcal{A}^{-1}(Z_j)) = \mathcal{A}^*(\mathcal{D}(Z_j)) = \mathcal{A}^*l_j.$$

As was pointed out above we then have

$$|\operatorname{Vol}_k P(\widetilde{Z}_1, \dots, \widetilde{Z}_k)| = ||\widetilde{l}_1 \wedge \dots \wedge \widetilde{l}_k|| = ||\mathcal{A}^* l_1 \wedge \dots \wedge \mathcal{A}^* l_k||, \tag{4.2.6}$$

and hence, the claim follows from (4.2.4) and (4.2.5) applied to $\mathcal{B} = \mathcal{A}^{-1}$.

We recall that an alternative formula for computing $\operatorname{Vol}_k P(Z_1, \dots, Z_k)$ was given earlier in Proposition 3.3.

Remark 4.11. Note that the above proof also shows that for any k vectors v_1, \ldots, v_k we have

$$\operatorname{Vol}_k P(v_1, \dots, v_k) = ||l_1 \wedge \dots \wedge l_k||,$$

where $l_j = \mathcal{D}(v_i), i = 1, \dots, k$.

4.3 Contraction

Let V be a vector space and $\phi \in \Lambda^k(V^*)$ a k-form. Define a (k-1)-form $\psi = v \, \lrcorner \, \phi$ by the formula

$$\psi(X_1,\ldots,X_{k-1}) = \phi(v,X_1,\ldots,X_{k-1})$$

for any vectors $X_1, \ldots, X_{k-1} \in V$. We say that the form ψ is obtained by a *contraction* of ϕ with the vector v. Sometimes, this operation is called also an *interior product* of ϕ with v and denoted by $i(v)\phi$ instead of $v \perp \psi$. In these notes we will not use this notation.

Proposition 4.12. Contraction \exists is a bilinear operation, i.e.

$$(v_1 + v_2) \, \lrcorner \, \phi = v_1 \, \lrcorner \, \phi + v_2 \, \lrcorner \, \phi$$
$$(\lambda v) \, \lrcorner \, \phi = \lambda (v \, \lrcorner \, \phi)$$
$$v \, \lrcorner \, (\phi_1 + \phi_2) = v \, \lrcorner \, \phi_1 + v \, \lrcorner \, \phi_2$$
$$v \, \lrcorner \, (\lambda \phi) = \lambda (v \, \lrcorner \, \phi).$$

Here $v, v_1, v_2 \in V$; $\phi, \phi_1, \phi_2 \in \Lambda^k(V^*)$; $\lambda \in \mathbb{R}$.

The proof is straightforward.

Let ϕ be a non-zero *n*-form. Then we have

Proposition 4.13. The map $\exists: V \to \Lambda^{n-1}(V^*)$, defined by the formula $\exists(v) = v \, \exists \, \phi$ is an isomorphism between the vector spaces V and $\Lambda^{n-1}(V^*)$.

Proof. Take a basis v_1, \ldots, v_n . Let $x_1, \ldots, x_n \in V^*$ be the dual basis, i.e. the corresponding coordinate system. Then $\phi = ax_1 \wedge \ldots \wedge x_n$, where $a \neq 0$. To simplify the notation let us assume that a = 1, so that

$$\phi = x_1 \wedge \ldots \wedge x_n.$$

Let us compute the images $v_i \, \lrcorner \, \phi$, $i = 1, \ldots, k$ of the basic vectors. Let us write

$$v_i \, \lrcorner \, \phi = \sum_{1}^n a_j x_1 \wedge \dots \wedge x_{j-1} \wedge x_{j+1} \wedge \dots \wedge x_n.$$

Then

$$v_{i} \, \, \, \, \, \, \, \phi(v_{1}, \dots, v_{l-1}, v_{l+1}, \dots, v_{n})$$

$$= \sum_{1}^{n} a_{j} x_{1} \wedge \dots \wedge x_{j-1} \wedge x_{j+1} \wedge \dots \wedge x_{n} (v_{1}, \dots, v_{l-1}, v_{l+1}, \dots, v_{n}) = a_{l}, \qquad (4.3.1)$$

but on the other hand,

$$v_{i} \, \, \downarrow \, \phi(v_{1}, \dots, v_{l-1}, v_{l+1}, \dots, v_{n}) = \phi(v_{i}, v_{1}, \dots, v_{l-1}, v_{l+1}, \dots, v_{n})$$

$$= (-1)^{i-1} \, \phi(v_{1}, \dots, v_{l-1}, v_{i}, v_{l+1}, \dots, v_{n}) = \begin{cases} (-1)^{i-1} &, l = i; \\ 0 &, \text{ otherwise} \end{cases}$$

$$(4.3.2)$$

Thus,

$$v_i \, \lrcorner \, \phi = (-1)^{i-1} x_1 \wedge \cdots \wedge x_{i-1} \wedge x_{i+1} \wedge \cdots \wedge x_n.$$

Hence, the map \bot sends a basis of V^* into a basis of $\Lambda^{n-1}(V^*)$, and therefore it is an isomorphism.

Take a vector $v = \sum_{1}^{n} a_{j}v_{j}$. Then we have

$$v \perp (x_1 \wedge \dots \wedge x_n) = \sum_{1}^{n} (-1)^{i-1} a_i x_1 \wedge \dots \wedge x_{i-1} \wedge x_{i+1} \wedge \dots \wedge x_n.$$
 (4.3.3)

This formula can be interpreted as the formula of expansion of a determinant according to the first column (or the first row). Indeed, for any vectors U_1, \ldots, U_{n-1} we have

$$v \, \lrcorner \, \phi(U_1, \dots, U_{n-1}) = \det(v, U_1, \dots, U_{n-1}) = \begin{vmatrix} a_1 & u_{1,1} & \dots & u_{1,n-1} \\ \dots & \dots & \dots & \dots \\ a_n & u_{n,1} & \dots & u_{n,n-1} \end{vmatrix},$$

where $\begin{pmatrix} u_{1,i} \\ \vdots \\ u_{n,i} \end{pmatrix}$ are coordinates of the vector $U_i \in V$ in the basis v_1, \ldots, v_n . On the other hand,

Suppose that dim V=3. Then the formula (4.3.3) can be rewritten as

$$v \perp (x_1 \wedge x_2 \wedge x_3) = a_1 x_2 \wedge x_3 + a_2 x_3 \wedge x_1 + a_3 x_1 \wedge x_2$$

where $\begin{pmatrix} a_1 \\ a_2 \\ a_3 \end{pmatrix}$ are coordinates of the vector V. Let us describe the geometric meaning of the

operation \bot . Set $\omega = v \bot (x_1 \land x_2 \land x_3)$. Then $\omega(U_1, U_2)$ is the volume of the parallelogram defined by the vectors U_1, U_2 and v. Let ν be the unit normal vector to the plane $L(U_1, U_2) \subset V$. Then we have

$$\omega(U_1, U_2) = \text{Area } P(U_1, U_2) \cdot \langle v, \nu \rangle$$
.

If we interpret v as the velocity of a fluid flow in the space V then $\omega(U_1, U_2)$ is just an amount of fluid flown through the parallelogram Π generated by vectors U_1 and U_2 for the unit time. It is called the flux of v through the parallelogram Π .

Let us return back to the case $\dim V = n$.

Exercise 4.14. Let

$$\alpha = x_{i_1} \wedge \cdots \wedge x_{i_k}, \quad 1 < i_1 < \cdots < i_k < n$$

and $v = (a_1, \ldots, a_n)$. Show that

$$v \, \lrcorner \, \alpha = \sum_{i=1}^{k} (-1)^{j+1} a_{i_j} x_{i_1} \wedge \dots x_{i_{j-1}} \wedge x_{i_{j+1}} \wedge \dots x_{i_k} \, .$$

The next proposition establishes a relation between the isomorphisms \star , \bot and \mathcal{D} .

Proposition 4.15. Let V be a Euclidean space, and x_1, \ldots, x_n be coordinates in an orthonormal basis. Then for any vector $v \in V$ we have

$$\star \mathcal{D}v = v \, \rfloor \, (x_1 \wedge \cdots \wedge x_n) \, .$$

Proof. Let $v = (a_1, \ldots, a_n)$. Then $\mathcal{D}v = a_1x_1 + \cdots + a_nx_n$ and

$$\star \mathcal{D}v = a_1 x_2 \wedge \cdots \wedge x_n - a_2 x_1 \wedge x_3 \wedge \cdots \wedge x_n \wedge x_1 + \cdots + (-1)^{n-1} a_n x_1 \wedge \cdots \wedge x_{n-1}.$$

But according to Proposition 4.13 the (n-1)-form $v \perp (x_1 \wedge \cdots \wedge x_n)$ is defined by the same formula.

We finish this section by the proposition which shows how the contraction operation interacts with the exterior product.

Proposition 4.16. Let α, β be forms of order k and l, respectively, and v a vector. Then

$$v \, \rfloor (\alpha \wedge \beta) = (v \, \rfloor \alpha) \wedge \beta + (-1)^k \alpha \wedge (v \, \rfloor \beta).$$

Proof. Note that given any indices $k_1, \ldots k_m$ (not necessarily ordered) we have $e_i \, \exists x_{k_1} \wedge \cdots \wedge x_{k_m} = 0$ if $i \notin \{k_1, \ldots, k_m\}$ and $e_i \, \exists x_{k_1} \wedge \cdots \wedge x_{k_m} = (-1)^J x_{k_1} \wedge \cdots \wedge x_{k_m}$, where $J = \operatorname{inv}(i; k_1, \ldots, k_m)$ is the number of variables ahead of x_i .

By linearity it is sufficient to consider the case when v, α, β are basic vector and forms, i.e.

$$v = e_i, \alpha = x_{i_1} \wedge \cdots \wedge x_{i_k}, \beta = x_{j_1} \dots x_{j_\ell}$$

We have $v \, \lrcorner \, \alpha \neq 0$ if and only if the index i is among the indices i_1, \ldots, i_k . In that case

$$v \, \lrcorner \, \alpha = (-1)^J x_{i_1} \wedge \dots \overset{i}{\vee} \dots \wedge x_{i_k}$$

and if $v \, \exists \, \beta \neq 0$ then

$$v \, \exists \, \beta = (-1)^{J'} x_{j_1} \wedge \dots \overset{i}{\vee} \dots \wedge x_{j_l},$$

where $J = \text{inv}(i; i_1, ..., i_k), J' = \text{inv}(i; j_1, ..., j_l).$

If it enters α bot not β then

$$v \, \rfloor \, (\alpha \wedge \beta) = (-1)^J x_{i_1} \wedge \dots \vee \cdots \wedge x_{i_k} \wedge x_{i_1} \wedge \dots \wedge x_{i_l} = (v \, \rfloor \, \alpha) \wedge \beta,$$

while $\alpha \wedge (v \, \lrcorner \, \beta) = 0$.

Similarly, if it enters β but not α then

$$v \perp (\alpha \wedge \beta) = (-1)^{J'+m} x_{i_1} \wedge \dots \wedge x_{i_k} \wedge x_{j_1} \wedge \dots \vee \dots \wedge x_{j_l} = (-1)^k \alpha \wedge (v \perp \beta),$$

while $v \, \lrcorner \, \alpha \wedge \beta = 0$. Hence, in both these cases the formula holds.

If x_i enters both products then $\alpha \wedge \beta = 0$, and hence $v \perp (\alpha \wedge \beta) = 0$.

On the other hand,

$$(v \, \exists \, \alpha) \wedge \beta + (-1)^k \alpha \wedge (v \, \exists \, \beta) = (-1)^J x_{i_1} \wedge \dots \overset{i}{\vee} \dots \wedge x_{i_k} \wedge x_{j_1} \dots \wedge x_{j_l}$$
$$+ (-1)^{k+J'} x_{i_1} \wedge \dots \wedge x_{i_k} \wedge x_{j_1} \wedge \dots \overset{i}{\vee} \dots \wedge x_{j_l} = 0,$$

because the products $x_{i_1} \wedge ... \vee ... \wedge x_{i_k} \wedge x_{j_1} \cdots \wedge x_{j_l}$ and $x_{i_1} \wedge ... \wedge x_{i_k} \wedge x_{j_1} \wedge ... \vee ... \wedge x_{j_l}$ differ only in the position of x_i . In the first product it is at the (k + J')-s position, and in the second at (J + 1)-st position. Hence, the difference in signs is $(-1)^{J+J'+k+1}$, which leads to the required cancellation.

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Chapter 5

Complex vector spaces

5.1 Complex numbers

The space \mathbb{R}^2 can be endowed with an associative and commutative multiplication operation. This operation is uniquely determined by three properties:

- it is a bilinear operation;
- the vector (1,0) is the unit;
- the vector (0,1) satisfies $(0,1)^2 = (0,-1)$.

The vector (0,1) is usually denoted by i, and we will simply write 1 instead of the vector (1,0). Hence, any point $(a,b) \in \mathbb{R}^2$ can be written as a+bi, where $a,b \in \mathbb{R}$, and the product of a+bi and c+di is given by the formula

$$(a+bi)(c+di) = ac - bd + (ad+bc)i.$$

The plane \mathbb{R}^2 endowed with this multiplication is denoted by \mathbb{C} and called the set of *complex numbers*. The real line generated by 1 is called the *real axis*, the line generated by *i* is called the *imaginary axis*. The set of real numbers \mathbb{R} can be viewed as embedded into \mathbb{C} as the real axis. Given a complex number z = x + iy, the numbers x and y are called its *real* and *imaginary* parts, respectively, and denoted by $\operatorname{Re} z$ and $\operatorname{Im} z$, so that $z = \operatorname{Re} z + i \operatorname{Im} z$.

For any non-zero complex number z = a + bi there exists an inverse z^{-1} such that $z^{-1}z = 1$. Indeed, we can set

$$z^{-1} = \frac{a}{a^2 + b^2} - \frac{b}{a^2 + b^2}i.$$

The commutativity, associativity and existence of the inverse is easy to check, but it should not be taken for granted: it is impossible to define a similar operation any \mathbb{R}^n for n > 2.

Given $z = a + bi \in \mathbb{C}$ its conjugate is defined as $\bar{z} = a - bi$. The conjugation operation $z \mapsto \bar{z}$ is the reflection of \mathbb{C} with respect to the real axis $\mathbb{R} \subset \mathbb{C}$. Note that

$$Rez = \frac{1}{2}(z + \bar{z}), \quad Imz = \frac{1}{2i}(z - \bar{z}).$$

Let us introduce the polar coordinates (r,ϕ) in $\mathbb{R}^2=\mathbb{C}$. Then a complex number z=x+yi can be written as $r\cos\phi+ir\sin\phi=r(\cos\phi+i\sin\phi)$. This form of writing a complex number is called, sometimes, ttrigonometric. The number $r=\sqrt{x^2+y^2}$ is called the *modulus* of z and denoted by |z| and ϕ is called the *argument* of ϕ and denoted by $\arg z$. Note that the argument is defined only $\mod 2\pi$. The value of the argument in $[0,2\pi)$ is sometimes called the *principal value* of the argument. When z is real than its modulus |z| is just the absolute value. We also not that $|z|=\sqrt{z\overline{z}}$.

An important role plays the triangle inequality

$$||z_1| - |z_2|| \le |z_1 + z_2| \le |z_1| + |z_2|.$$

Exponential function of a complex variable

Recall that the exponential function e^x has a Taylor expansion

$$e^x = \sum_{n=0}^{\infty} \frac{x^n}{n!} = 1 + x + \frac{x^2}{2} + \frac{x^3}{6} + \dots$$

We then define for a *complex* the exponential function by the same formula

$$e^z := 1 + z + \frac{z^2}{2!} + \dots + \frac{z^n}{n!} + \dots$$

One can check that this power series absolutely converging for all z and satisfies the formula

$$e^{z_1 + z_2} = e^{z_1} e^{z_2}.$$



Figure 5.1: Leonhard Euler (1707-1783)

In particular, we have

$$e^{iy} = 1 + iy - \frac{y^2}{2!} - i\frac{y^3}{3!} + \frac{y^4}{4!} + \dots + \dots$$
 (5.1.1)

$$= \sum_{k=0}^{\infty} (-1)^k \frac{y^{2k}}{2k!} + i \sum_{k=0}^{\infty} (-1)^k \frac{y^{2k+1}}{(2k+1)!}.$$
 (5.1.2)

But $\sum_{k=0}^{\infty} (-1)^k \frac{y^{2k}}{2k!} = \cos y$ and $\sum_{k=0}^{\infty} (-1)^k \frac{y^{2k+1}}{(2k+1)!} = \sin y$, and hence we get Euler's formula

$$e^{iy} = \cos y + i \sin y$$
.

and furthermore,

$$e^{x+iy} = e^x e^{iy} = e^x (\cos y + i \sin y),$$

i.e. $|e^{x+iy}| = e^x, \arg(e^z) = y$. In particular, any complex number $z = r(\cos \phi + i \sin \phi)$ can be rewritten in the form $z = re^{i\phi}$. This is called the *exponential form* of the complex number z.

Note that

$$\left(e^{i\phi}\right)^n = e^{in\phi},$$

and hence if $z = re^{i\phi}$ then $z^n = r^n e^{in\phi} = r^n (\cos n\phi + i \sin n\phi)$.

Note that the operation $z \mapsto iz$ is the rotation of \mathbb{C} counterclockwise by the angle $\frac{\pi}{2}$. More generally a multiplication operation $z \mapsto zw$, where $w = \rho e^{i\theta}$ is the composition of a rotation by the angle θ and a radial dilatation (homothety) in ρ times.

Exercise 5.1. 1. Compute $\sum_{0}^{n} \cos k\theta$ and $\sum_{1}^{n} \sin k\theta$.

2. Compute $1 + \binom{n}{4} + \binom{n}{8} + \binom{n}{12} + \dots$

5.2 Complex vector space

In a real vector space one knows how to multiply a vector by a real number. In a complex vector space there is defined an operation of multiplication by a complex number. Example is the space \mathbb{C}^n whose vectors are n-tuples $z=(z_1,\ldots,z_n)$ of complex numbers, and multiplication by any complex number $\lambda=\alpha+i\beta$ is defined component-wise: $\lambda(z_1,\ldots,z_n)=(\lambda z_1,\ldots,\lambda z_n)$. Complex vector space can be viewed as an upgrade of a real vector space, or better to say as a real vector space with an additional structure.

In order to make a real vector space V into a complex vector space, one just needs to define how to multiply a vector by i. This operation must be a linear map $\mathcal{J}:V\to V$ which should satisfy the condition $\mathcal{J}^2=-\mathrm{Id}$, i.e $\mathcal{J}(\mathcal{J}(v))=i(iv)=-v$.

Example 5.2. Consider \mathbb{R}^{2n} with coordinates $(x_1, y_1, \dots, x_n, y_n)$. Consider a $2n \times 2n$ -matrix

$$J = \begin{pmatrix} 0 & -1 & 0 & 0 \\ 1 & 0 & 0 & 0 \\ 0 & 0 & 0 & -1 \\ 0 & 0 & 1 & 0 \\ & & & \dots \\ & & & 0 & -1 \\ & & & & 1 & 0 \end{pmatrix}$$

Then $J^2 = -I$. Consider a linear operator $\mathcal{J}: \mathbb{R}^{2n} \to \mathbb{R}^{2n}$ with this matrix, i.e. $\mathcal{J}(Z) = JZ$ for any vector $Z \in \mathbb{R}^{2n}$ which we view as a column-vector. Then $\mathcal{J}^2 = -\mathrm{Id}$, and hence we can define on \mathbb{R}^{2n} a complex structure (i.e.the multiplication by i by the formula

$$iZ = \mathcal{J}(Z), \ Z \in \mathbb{R}^{2n}.$$

This complex vector space is canonically isomorphic to \mathbb{C}^n , where we identify the real vector $(x_1, y_1, \ldots, x_n, y_n) \in \mathbb{R}^{2n}$ with a complex vector $(z_1 = x_1 + iy_1, \ldots, z_n x_n + iy_n)$.

On the other hand, any complex vector space can be viewed as a real vector space. In order to do that we just need to "forget" how to multiply by i. This procedure is called the *realification* of a complex vector space. For example, the realification of \mathbb{C}^n is \mathbb{R}^{2n} . Sometimes to emphasize the realification operation we will denote the realification of a complex vector space V by $V_{\mathbb{R}}$. As the sets these to objects coincide.

Given a complex vector space V we can define linear combinations $\sum \lambda_i v_i$, where $\lambda_i \in \mathbb{C}$ are complex numbers, and thus similarly to the real case talk about really dependent, really independent vectors. Given vectors $v_1, \ldots v_n \in V$ we define its $complex \ span \ \mathrm{Span}_{\mathbb{C}}(V_1, \ldots, v_n)$ by the formula

$$\operatorname{Span}_{\mathbb{C}}(v_1,\ldots,v_n) = \left\{ \sum_{1}^{n} \lambda_i v_i, \ \lambda_i \in \mathbb{C}. \right\}$$

. A basis of a complex vector space is a system of complex linear independent vectors v_1, \ldots, v_n such that $\operatorname{Span}_{\mathbb{C}}(v_1, \ldots, v_n) = V$. The number of vectors in a complex basis is called the *complex dimension* of V and denoted $\dim_{\mathbb{C}} V$.

For instance dim $\mathbb{C}^n = n$. On the other hand, its realification \mathbb{R}^{2n} has real dimension 2n. In particular, \mathbb{C} is a complex vector space of dimension 1, and therefore it is called a complex *line* rather than a plane.

Exercise 5.3. Let v_1, \ldots, v_n be a complex basis of a complex vector space V. Find the real basis of its realification $V_{\mathbb{R}}$.

Answer. $v_1, iv_1, v_2, iv_2, \dots, v_n, iv_n$.

There is another important operation which associates with a real vector space V of real dimension n a complex vector space $V_{\mathbb{C}}$ of complex dimension n. It is done in a way similar to how we made complex numbers out of real numbers. As a real vector space the space $V_{\mathbb{C}}$ is just the direct sum $V \oplus V = \{(v, w); v, w \in V\}$. This is a real space of dimension 2n. We then make $V \oplus V$ into a complex vector space by defining the multiplication by i by the formula:

$$i(v, w) = (-w, v).$$

We will write vector (v,0) simply by v and (0,v)=i(v,0) by iv. Hence, every vector of $V_{\mathbb{C}}$ can be written as v+iw, where $v,w\in V$. If v_1,\ldots,v_n is a real basis of V, then the same vectors form a complex basis of $V_{\mathbb{C}}$.

Complexification of a real vector space

Given a real space V one can associate with it a complex vector space $V_{\mathbb{C}}$, called the *complexification* of V, as follows. It is made of vectors of V in exactly the same way as complex numbers made of reals. Namely, $V_{\mathbb{C}}$ consists of expressions X + iY, where $X, Y \in V$. We define a multiplication of a complex number a + ib by a vector X + iY by a formula

$$(a+ib)(X+iY) = aX - bY + i(aY + bX).$$

For instance, \mathbb{C}^n is canonically isomorphic to $(\mathbb{R}^n)_{\mathbb{C}}$, because any vector $Z = (z_1 = x_1 + iy_1, \dots, z_n = x_n + iy_n) \in \mathbb{C}^n$ can be uniquely written as Z = X + iY, where $X = (x_1, \dots, x_n), Y = (y_1, \dots, y_n) \in \mathbb{R}^n$. Note that the *realification* of the space $V_{\mathbb{C}}$, i.e. the space $(V_{\mathbb{C}})_{\mathbb{R}}$ is canonically is just $V \oplus V = \{(X, Y), X, Y \in \mathbb{R}^n\}$.

If v_1, \ldots, v_n is a basis of a V over \mathbb{R} , then the same vectors form a basis of the complexified space $V_{\mathbb{C}}$ over \mathbb{C} . Thus $\dim_{\mathbb{R}} V = \dim_{\mathbb{C}} V_{\mathbb{C}}$.

5.3 Complex linear maps

Complex linear maps and their realifications

Given two complex vector spaces V, W a map $\mathcal{A}: V \to W$ is called *complex linear* (or \mathbb{C} -linear) if $\mathcal{A}(X+Y) = \mathcal{A}(X) + \mathcal{A}(Y)$ and $\mathcal{A}(\lambda X) = \lambda \mathcal{A}(X)$ for any vectors $X, Y \in V$ and any *complex* number $\lambda \in \mathbb{C}$. Thus complex linearity is stronger condition than the real linearity. The difference is in the additional requirement that $\mathcal{A}(iX) = i\mathcal{A}(X)$. In other words, the operator \mathcal{A} must commute with the operation of multiplication by i.

Any linear map $\mathcal{A}: \mathbb{C} \to \mathbb{C}$ is a multiplication by a complex number a=c+id. If we view \mathbb{C} as \mathbb{R}^2 and right the real matrix of this map in the standard basis 1, i we get the matrix $\begin{pmatrix} c & -d \\ d & c \end{pmatrix}$. Indeed, $\mathcal{A}(1)=a=c+di$ and $\mathcal{A}(i)=ai=-d+ci$, so the first column of the matrix is equal to $\begin{pmatrix} c \\ d \end{pmatrix}$ and the second one is equal to $\begin{pmatrix} -d \\ c \end{pmatrix}$.

If we have bases v_1, \ldots, v_n of V and $w_1, \ldots w_m$ of W then one can associate with \mathcal{A} an $m \times n$ complex matrix A by the same rule as in the real case.

Recall (see Exercise 5.3) that vectors $v_1, v_1' = iv_1, v_2, v_2' = iv_2, \dots, v_n, v_n' = iv_n$ and $w_1, w_1' = iw_1, w_2, w_2' = iw_2, \dots w_m, w_m' = iw_m$) form real bases of the realifications $V_{\mathbb{R}}$ and $W_{\mathbb{R}}$ of the spaces V and W. if

$$A = \begin{pmatrix} a_{11} & \dots & a_{1n} \\ & \dots & \\ a_{m1} & \dots & a_{mn} \end{pmatrix}$$

is the complex matrix of \mathcal{A} then the real matrix $A_{\mathbb{R}}$ of the map \mathcal{A} is the real basis has order $2n \times 2n$ and is obtained from A by replacing each complex element $a_{kl} = c_{kl} + id_{kl}$ by a 2×2 matrix $\begin{pmatrix} c_{kl} & -d_{kl} \\ d_{kl} & c_{kl} \end{pmatrix}$.

Exercise 5.4. Prove that

$$\det A_{\mathbb{R}} = |\det A|^2.$$

Complexification of real linear maps

Given a real linear map $\mathcal{A}:V\to W$ one can define a complex linear map $\mathcal{A}_{\mathbb{C}}:V_{\mathbb{C}}\to W_{\mathbb{C}}$ by the formula

$$\mathcal{A}_{\mathbb{C}}(v+iw) = \mathcal{A}(v) + i\mathcal{A}(w).$$

. If A is the matrix of \mathcal{A} in a basis v_1, \ldots, v_n then $\mathcal{A}_{\mathbb{C}}$ has the same matrix in the same basis viewed as a complex basis of $V_{\mathbb{C}}$. The operator $\mathcal{A}_{\mathbb{C}}$ is called the *complexification* of the operator \mathcal{A} .

In particular, one can consider \mathbb{C} -linear functions $V \to \mathbb{C}$ on a complex vector space V. Complex coordinates z_1, \ldots, z_n in a complex basis are examples of \mathbb{C} -linear functions, and any other \mathbb{C} -linear function on V has a form $c_1z_1 + \ldots + c_nz_n$, where $c_1, \ldots, c_n \in \mathbb{C}$ are complex numbers.

Complex-valued \mathbb{R} -linear functions

It is sometimes useful to consider also \mathbb{C} -valued \mathbb{R} -linear functions on a complex vector space V, i.e. \mathbb{R} -linear maps $V \to \mathbb{C}$ (i.e. a linear map $V_{\mathbb{R}} \to \mathbb{R}^2$). Such a \mathbb{C} -valued function has the form $\lambda = \alpha + i\beta$, where α, β are usual real linear functions. For instance the function \bar{z} on \mathbb{C} is a \mathbb{C} -valued \mathbb{R} -linear function which is not \mathbb{C} -linear.

If z_1, \ldots, z_n are complex coordinates on a complex vector space V then any \mathbb{R} -linear complexvalued function can be written as $\sum_{i=1}^{n} a_i z_i + b_i \bar{z}_i$, where $a_i, b_i \in \mathbb{C}$ are complex numbers.

We can furthermore consider complex-valued tensors and, in particular complex-valued exterior forms. A \mathbb{C} -valued k-form λ can be written as $\alpha + i\beta$ where α and β are usual \mathbb{R} -valued k-forms. For instance, we can consider on \mathbb{C}^n the 2-form $\omega = \frac{i}{2} \sum_{1}^{n} z_k \wedge \bar{z}_k$. It can be rewritten as $\omega = \frac{i}{2} \sum_{1}^{n} (x_k + iy_k) \wedge (x_k - iy_k) = \sum_{1}^{n} x_k \wedge y_k$.

Part II

Calculus of differential forms

Chapter 6

Topological preliminaries

6.1 Elements of topology in a vector space

We recall in this section some basic topological notions in a finite-dimensional vector space and elements of the theory of continuous functions. The proofs of most statements are straightforward and we omit them.

Let us choose in V a scalar product.

Notation $B_r(p) := \{x \in V; ||x - p|| < r\}, D_r(p) := \{x \in V; ||x - p|| \le r\} \text{ and } S_r(p) := \{x \in V; ||x - p|| = r\} \text{ stand for open, closed balls and the sphere of radius } r \text{ centered at a point } p \in V.$

Open and closed sets

A set $U \subset V$ is called *open* if for any $x \in U$ there exists $\epsilon > 0$ such that $B_{\epsilon}(x) \subset U$.

A set $A \subset V$ is called *closed* if its complement $V \setminus A$ is open. Equivalently,

Lemma 6.1. The set A is closed if and only if for any sequence $x_n \in A$, n = 1, 2, ... which converges to a point $a \in V$, the limit point a belongs to A.

Remark 6.2. It is important to note that the notion of open and closed sets are independent of the choice of the auxiliary Euclidean structure in the space V.

Points which appear as limits of sequences of points $x_n \in A$ are called *limit points* of A.

There are only two subsets of V which are simultaneously open and closed: V and \varnothing .

Lemma 6.3. 1. For any family U_{λ} , $\lambda \in \Lambda$ of open sets the union $\bigcup_{\lambda \in \Lambda} U_{\lambda}$ is open.

- 2. For any family A_{λ} , $\lambda \in \Lambda$ of closed sets the intersection $\bigcap_{\lambda \in \Lambda} A_{\lambda}$ is closed.
- 3. The union $\bigcup_{i=1}^{n} A_i$ of a finite family of closed sets is closed.
- 4. The intersection $\bigcap_{i=1}^{n} U_i$ of a finite family of open sets is open.

By a neighborhood of a point $a \in V$ we understand any open set $U \ni p$.

Given any subset $X \subset V$ a point $a \in V$ is called

- interior point for A if there is a neighborhood $U \ni p$ such that $U \subset A$;
- boundary point if it is not an interior point neither for A nor for its complement $V \setminus A$.

We emphasize that a boundary point of A may or may not belong to A. Equivalently, a point $a \in V$ is a boundary point of A if it is a limit point both for A and $V \setminus A$.

The set of all interior points of A is called the *interior of* A and denoted Int A. The set of all boundary points of A is called the *boundary* of A and denoted ∂A . The union of all limit points of A is called the *closure* of A and denoted \overline{A} .

Lemma 6.4. 1. We have $\overline{A} = A \cup \partial A$, Int $A = A \setminus \partial A$.

- 2. \overline{A} is equal to the intersection of all closed sets containg A
- 3. Int U is the union of all open sets contained in A.

Given a subset $X \subset V$

- a subset $Y \subset X$ is called *relatively open* in Y if there exists an open set $U \subset V$ such that $Y = X \cap U$.
- a subset $Y \subset X$ is called *relatively closed* in Y if there exists a closed set $A \subset V$ such that $Y = X \cap A$.

One also call relatively and open and closed subsets of X just open and closed in X.

Exercise 6.5. Prove that though we defined open sets using a Euclidean structure on the vector $space\ V$ the definition of open and closed sets is independent of this choice.



Figure 6.1: Bernard Bolzano (1781-1848)

6.2 Everywhere and nowhere dense sets

A closed set A is called nowhere dense if $\operatorname{Int} A = \emptyset$. For instance any finite set is nowhere dense. Any linear subspace $L \subset V$ is nowhere dense in V if $\dim L < \dim V$. Here is a more interesting example of a nowhere dense set.

Fix some number $\epsilon < 1$. For any interval $\Delta = [a,b]$ we denote by Δ_{ϵ} the *open* interval centered at the point $c = \frac{a+b}{2}$, the middle point of Δ , of the total length equal to $\epsilon(b-a)$. We denote by $C(\Delta) := \Delta \setminus \Delta_{\epsilon}$. Thus $C(\Delta)$ consists of two disjoint smaller closed intervals. Let I = [0,1]. Take $C(I) = I_1 \cup I_2$. Take again $C(I_1) \cup C(I_2)$ then again apply the operation C to four new closed intervals. Continue the process, and take the intersection of all sets arising on all steps of this construction. The resulted closed set $K_{\epsilon} \subset I$ is nowhere dense. It is called a *Cantor set*.

A subset $B \subset A$ is called *everywhere dense* in A if $\overline{B} \supset A$. For instance the set $\mathbb{Q} \cap I$ of rational points in the interval I = [0, 1] is everywhere dense in I.

6.3 Compactness and connectedness

A set $A \subset V$ is called *compact* if one of the following equivalent conditions is satisfied:

COMP1. A is closed and bounded.



Figure 6.2: Karl Weierstrass (1815-1897)

COMP2. from any infinite sequence of points $x_n \in A$ one can choose a subsequence x_{n_k} converging to a point $a \in A$.

COMP3. from any family $U_{\lambda}, \lambda \in \Lambda$ of open sets covering A, i.e. $\bigcup_{\lambda \in \Lambda} U_{\lambda} \supset A$, one can choose finitely many sets $U_{\lambda_1}, \dots, U_{\lambda_k}$ which cover A, i.e. $\bigcup_{1}^{k} U_{\lambda_k} \supset A$.

The equivalence of these definitions is a combination of theorems of Bolzano-Weierstrass and Émile Borel.

A set A is called *path-connected* if for any two points $a_0, a_1 \in A$ there is a continuous path $\gamma: [0,1] \to A$ such that $\gamma(0) = a_0$ and $\gamma(1) = a_1$.

A set A is called *connected* if one cannot present A as a union $A = A_1 \cup A_2$ such that $A_1 \cap A_2 = \emptyset$, $A_1, A_2 \neq \emptyset$ and both A_1 and A_2 are simultaneously relatively closed and open in A.

Lemma 6.6. Any path-connected set is connected.

Proof. Suppose that A is disconnected. Then it can be presented as a union $A = A_0 \cup A_1$ of two non-empty relatively open (and hence relatively closed) subsets. Consider the function $\phi : A \to \mathbb{R}$ defined by the formula

$$\phi(x) = \begin{cases} 0, & x \in A_0, \\ 1, & x \in A_1. \end{cases}$$



Figure 6.3: Émile Borel (1871-1956)

We claim that the function ϕ is is continuous. Indeed, For each i=0,1 and any point $a \in A_i$ there exists $\epsilon > 0$ such that $B_{\epsilon}(x) \cap A \subset A_i$. Hence the function ϕ is constant on $B_{\epsilon}(x) \cap A$, and hence continuous at the point x. Now take points $x_0 \in A_0$ and $x_1 \in A_1$ and connect them by a path $\gamma : [0,1] \to A$ (this path exists because A is path-connected). Consider the function $\psi := \phi \circ \gamma : [0,1] \to \mathbb{R}$. This function is continuous (as a composition of two continuous maps). Furthermore, $\psi(0) = 0, \psi(1) = 1$. Hence, by an intermediate value theorem of Cauchy the function ψ must take all values in the interval [0,1]. But this is a contradiction because by construction the function ψ takes no other values except 0 and 1.

Lemma 6.7. Any open connected subset $U \subset \mathbb{R}^n$ is path connected.

Proof. Take any point $a \in U$. Denote by C_a the set of all points in U which can be connected with a by a path. We need to prove that $C_a = U$.

First, we note that C_a is open. Indeed, if $b \in C_a$ then using openness of U we can find $\epsilon > 0$ such the ball $B_{\epsilon}(b) \in U$. Any point of $c \in B_{\epsilon}(b)$ can be connected by a straight interval $I_{bc} \subset B_{\epsilon}(b)$ with b, and hence it can be connected by a path with a, i.e. $c \in C_a$. Thus $B_{\epsilon}(b) \subset C_a$, and hence C_a is open. Similarly we prove that the complement $U \setminus C_a$ is open. Indeed, take $b \notin C_a$. As above, there exists an open ball $B_{\epsilon}(b) \subset U$. Then $B_{\epsilon}(b) \subset U \setminus C_a$. Indeed, if it were possible to connect a point $c \in B_{\epsilon}(b)$ with a by a path, then the same would be true for b, because b and c are connected by the interval I_{bc} . Thus, we have $U = C_a \cup (U \setminus C_a)$, both sets C_a and $U \setminus C_a$ are open and C_a is

non-empty. Hence, $U \setminus C_a$ is to be empty in view of connectedness of U. Thus, $C_a = U$, i.e. U is path-connected.

Exercise 6.8. In general a connected set need not to be path-connected. A canonical example is the closure of the graph of the function $\sin \frac{1}{x}, x \in \mathbb{R} \setminus 0$.. Prove it.

Exercise 6.9. Prove that any non-empty connected (= path-connected) open subset of \mathbb{R} is equal to an interval (a,b) (we allow here $a=-\infty$ and $b=\infty$). If one drops the condition of openness, then one needs to add a closed and semi-closed intervals and a point.

Remark 6.10. One of the corollaries of this exercise is that in \mathbb{R} any connected set is path-connected.

Solution. Let $A \subset \mathbb{R}$ be a non-empty connected subset. Let a < b be two points of A. Suppose that a point $c \in (a,b)$ does not belong to A. Then we can write $A = A_0 \cup A_1$, where $A_0 = A \cap (-\infty,0)$, $A_1 = A \cap (0,\infty)$. Both sets A_0 and A_1 are relatively open and non-empty, which contradicts connectedness of A. Hence if two points a and b, a < b, are in A, then the whole interval [a,b] is also contained in A. Denote $m := \inf A$ and $M := \sup A$ (we assume that $m = -\infty$ if A is unbounded from below and $M = +\infty$ if A is unbounded from above). Then the above argument shows that the open interval (m, M) is contained in A. Thus, there could be 5 cases:

- $m, M \notin A$; in this case A = (m, M);
- $m \in A, M \notin A$; in this case A = [m, M);
- $m \notin A, M \in A$; in this case A = (m, M];
- $m, M \in A$ and m < M; in this case A = [m, M];
- $m, M \in A$ and m = M; in this case A consists of one point.

6.4 Connected and path-connected components

Lemma 6.11. Let A_{λ} , $\lambda \in \Lambda$ be any family of connected (resp. path-connected) subsets of a vector space V. Suppose $\bigcap_{\lambda \in \Lambda} A_{\lambda} \neq \emptyset$. Then $\bigcup_{\lambda \in \Lambda} A_{\lambda}$ is also connected (resp. path-connected)

Proof. Pick a point $a \in \bigcap_{\lambda \in \Lambda} A_{\lambda}$. Consider first the case when A_{λ} are path connected. Pick a point $a \in \bigcap_{\lambda \in \Lambda} A_{\lambda}$. Then a can be connected by path with all points in A_{λ} for any points in $\lambda \in \Lambda$. Hence, all points of A_{λ} and $A_{\lambda'}$ can be connected with each other for any $\lambda, \lambda' \in \Lambda$.

Suppose now that A_{λ} are connected. Denote $A:=\bigcup_{\lambda\in\Lambda}A_{\lambda}$. Suppose A can be presented as a union $A=U\cup U'$ of disjoint relatively open subsets, where we denoted by U the set which contains the point $a\in\bigcap_{\lambda\in\Lambda}A_{\lambda}$. Then for each $\lambda\in\Lambda$ the intersections $U_{\lambda}:=U\cap A_{\lambda}$ and $U'_{\lambda}:=U'\cap A_{\lambda}$ are relatively open in A_{λ} . We have $A_{\lambda}=U_{\lambda}\cup U'_{\lambda}$. By assumption, $U_{\lambda}\ni a$, and hence $U_{\lambda}\neq\varnothing$. Hence, connectedness of A_{λ} implies that $U'_{\lambda}=\varnothing$. But then $U'=\bigcup_{\lambda\in\Lambda}U'_{\lambda}=\varnothing$, and therefore A is connected.

Given any set $A \subset V$ and a point $a \in A$ the connected component (resp. path-connected component $C_a \subset A$ of the point $a \in A$ is the union of all connected (resp. path-connected) subsets of A which contains the point a. Due to Lemma 6.11 the (path-)connected component C_a is itself (path-)connected, and hence it is the biggest (path-)connected subset of A which contains the point a. The path-connected component of a can be equivalently defined as the set of all points of A one can connect with a by a path in A.

Note that (path-)connected components of different points either coincide or do not intersect, and hence the set A can be presented as a disjoint union of (path-)-connected components.

Lemma 6.7 shows that for open sets in a vector space V the notions of connected and path-connected components coincide, and due to Exercise 6.9 the same is true for any subsets in \mathbb{R} . In particular, any open set $U \subset \mathbb{R}$ can be presented as a union of disjoint open intervals, which are its connected (= path-connected) components. Note that the number of these intervals can be infinite, but always countable.

6.5 Continuous maps and functions

Let V, W be two Euclidean spaces and A is a subset of V. A map $f: A \to W$ is called *continuous* if one of the three equivalent properties hold:

- 1. For any $\epsilon > 0$ and any point $x \in A$ there exists $\delta > 0$ such that $f(B_{\delta}(x) \cap A) \subset B_{\epsilon}(f(x))$.
- 2. If for a sequence $x_n \in A$ there exists $\lim x_n = x \in A$ then the sequence $f(x_n) \in W$ converges

to f(x).

- 3. For any open set $U \subset W$ the pre-image $f^{-1}(U)$ is relatively open in A.
- 4. For any closed set $B \subset W$ the pre-image $f^{-1}(B)$ is relatively closed in A.

Let us verify equivalence of 3 and 4. For any open set $U \subset W$ its complement $B = W \setminus U$ is closed and we have $f^{-1}(U) = A \setminus f^{-1}(B)$. Hence, if $f^{-1}(U)$ is relatively open, i.e. $f^{-1}(U) = U' \cap A$ for an open set $U' \subset V$, then $f^{-1}(B) = A \cap (V \setminus U')$, i.e. $f^{-1}(B)$ is relatively closed. The converse is similar.

Let us deduce 1 from 3. The ball $B_{\epsilon}(f(x))$ is open. Hence $f^{-1}(B_{\epsilon}(f(x)))$ is relatively open in A. Hence, there exists $\delta > 0$ such that $B_{\delta}(x) \cap A \subset f^{-1}(B_{\epsilon}(f(x)))$, i.e. $f(B_{\delta}(x) \cap A) \subset B_{\epsilon}(f(x))$. We leave the converse and the equivalence of definition 2 to the reader.

Remark 6.12. Consider a map $f: A \to W$ and denote B := f(A). Then definition 3 can be equivalently stated as follows:

3'. For any set $U \subset B$ relatively open in B its pre-image $f^{-1}(U)$ is relatively open in A.

Definition 4 can be reformulated in a similar way.

Indeed, we have $U = U' \cap A$ for an open set $U' \subset W$, while $f^{-1}(U) = f^{-1}(U')$.

The following theorem summarize properties of continuous maps.

Theorem 6.13. Let $f: A \to W$ be a continuous map. Then

- 1. if A is compact then f(A) is compact;
- 2. if A is connected then f(A) is connected;
- 3. if A is path connected then f(A) is path-connected.

Proof. 1. Take any infinite sequence $y_n \in f(A)$. Then there exist points $x_n \in A$ such that $y_n = f(x_n), n = 1, \ldots$ Then there exists a converging subsequence $x_{n_k} \to a \in A$. Then by continuity $\lim_{k \to \infty} f(x_{n_k}) = f(a) \in f(A)$, i.e. f(A) is compact.

2. Suppose that f(A) can be presented as a union $B_1 \cup B_2$ of simultaneously relatively open and closed disjoint non-empty sets. Then $f^{-1}(B_1), f^{-1}(B_2) \subset A$ are simultaneously relatively open

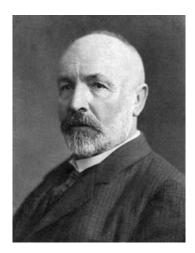


Figure 6.4: George Cantor (1845-1918)

and closed in A, disjoint and non-empty. We also have $f^{-1}(B_1) \cup f^{-1}(B_2) = f^{-1}(B_1 \cup B_2) = f^{-1}(f(A)) = A$. Hence A is disconnected which is a contradiction.

3. Take any two points $y_0, y_1 \in f(A)$. Then there exist $x_0, x_1 \in A$ such that $f(x_0) = y_0, f(x_1) = y_1$. But A is path-connected. Hence the points x_0, x_1 can be connected by a path $\gamma : [0, 1] \to A$. Then the path $f \circ \gamma : [0, 1] \to f(A)$ connects y_0 and y_1 , i.e. f(A) is path-connected.

Note that in the case $W = \mathbb{R}$ Theorem 6.13.1 is just the Weierstrass theorem: a continuos function on a compact set is bounded and achieves its maximal and minimal values.

We finish this section by a theorem of George Cantor about uniform continuity.

Theorem 6.14. Let A be compact and $f: A \to W$ a continuous map. Then for any $\epsilon > 0$ there exists $\delta > 0$ such that for any $x \in A$ we have $f(B_{\delta}(x)) \subset B_{\epsilon}(f(x))$.

Proof. Choose $\epsilon > 0$. By continuity of f for every point $x \in A$ there exists $\delta(x) > 0$ such that

$$f(B_{\delta(x)}(x)) \subset B_{\frac{\epsilon}{4}}(f(x)).$$

We need to prove that $\inf_{x\in A}\delta(x)>0$. Note that for any point in $y\in B_{\frac{\delta(x)}{2}}(x)$ we have $B_{\frac{\delta(x)}{2}}(y)\subset B_{\delta(x)}(x)$, and hence $f(B_{\frac{\delta(x)}{2}}(y))\subset B_{\epsilon}(f(y))$. By compactness, from the covering $\bigcup_{x\in A}B_{\frac{\delta(x)}{2}}(x)$ we can choose a finite number of balls $B_{\frac{\delta(x_j)}{2}}(x_j), j=1,\ldots,N$ which still cover A. Then $\delta=\min_k\frac{\delta(x_j)}{2}$ satisfy the condition of the theorem, i.e. $f(B_{\delta}(x))\subset B_{\epsilon}(f(x))$ for any $x\in A$.

Chapter 7

Vector fields and differential forms

7.1 Differential and gradient

Given a vector space V we will denote by V_x the vector space V with the origin translated to the point $x \in V$. One can think of V_x as that tangent space to V at the point x. Though the parallel transport allows one to identify spaces V and V_x it will be important for us to think about them as different spaces.

Let $f: U \to \mathbb{R}$ be a function on a domain $U \subset V$ in a vector space V. The function f is called differentiable at a point $x \in U$ if there exists a linear function $l: V_x \to \mathbb{R}$ such that

$$f(x+h) - f(x) = l(h) + o(||h||)$$

for any sufficiently small vector h, where the notation o(t) stands for any function such that $\frac{o(t)}{t} \underset{t \to 0}{\to} 0$. The linear function l is called the differential of the function f at the point x and is denoted by $d_x f$. In other words, f is differentiable at $x \in U$ if for any $h \in V_x$ there exists a limit

$$l(h) = \lim_{t \to 0} \frac{f(x+th) - f(x)}{t},$$

and the limit l(h) linearly depends on h. The value $l(h) = d_x f(h)$ is called the directional derivative of f at the point x in the direction h. The function f is called differentiable on the whole domain U if it is differentiable at each point of U.

Simply speaking, the differentiability of a function means that at a small scale near a point x the function behaves approximately like a linear function, the differential of the function at the

point x. However this linear function varies from point to point, and we call the family $\{d_x f\}_{x \in U}$ of all these linear functions the differential of the function f, and denote it by df (without a reference to a particular point x).

Let us summarize the above discussion. Let $f:U\to\mathbb{R}$ be a differentiable function. Then for each point $x\in U$ there exists a linear function $d_xf:V_x\to\mathbb{R}$, the differential of f at the point x defined by the formula

$$d_x f(h) = \lim_{t \to 0} \frac{f(x+th) - f(x)}{t}, x \in U, h \in V_x.$$

We recall that existence of partial derivatives at a point $a \in U$ does not guarantee the differentiability of f at the point a. On the other hand if partial derivatives exists in a neighborhood of a and continuous at the point a then f is differentiable at this point. The functions whose first partial derivatives are continuous in u are called C^1 -smooth, or sometimes just smooth. Equivalently, we can say that f is smooth if the differential $d_x f$ continuously depends on the point $x \in U$.

If v_1, \ldots, v_n are vectors of a basis of V, parallel transported to the point x, then we have

$$d_x f(v_i) = \frac{\partial f}{\partial x_i}(x), x \in U, i = 1, \dots, n,$$

where x_1, \ldots, x_n are coordinates with respect to the chosen basis v_1, \ldots, v_n .

Notice that if f is a linear function,

$$f(x) = a_1 x_1 + \dots + a_n x_n \,,$$

then for each $x \in V$ we have

$$d_x f(h) = a_1 h_1 + \dots + a_n h_n$$
, $h = (h_1, \dots, h_n) \in V_x$.

Thus the differential of a linear function f at any point $x \in V$ coincides with this function, parallel transported to the space V_x . This observation, in particular, can be applied to linear coordinate functions x_1, \ldots, x_n with respect to a chosen basis of V.

In Section 7.7 below we will define the differential for maps $f: U \to W$, where W is a vector space and not just the real line \mathbb{R} .

7.2 Smooth functions

We recall that existence of partial derivatives at a point $a \in U$ does not guarantee the differentiability of f at the point a. On the other hand if partial derivatives exists in a neighborhood of a and continuous at the point a then f is differentiable at this point. The functions whose first partial derivatives are continuous in u are called C^1 -smooth. Equivalently, we can say that f is smooth if the differential $d_x f$ continuously depends on the point $x \in U$.

More generally, for $k \geq 1$ a function $f: U \to \mathbb{R}$ is called C^k -smooth all its partial derivatives up to order k are continuous in U. The space of C^k -smooth functions is denoted by $C^k(U)$. We will also use the notation $C^0(U)$ and $C^\infty(U)$ which stands, respectively, for the spaces of continuous functions and functions with continuous derivatives of all orders. In this notes we will often speak of smooth functions without specifying the class of smoothness, assuming that functions have as many continuous derivatives as necessary to justify our computations.

Remark 7.1. We will often need to consider smooth maps, functions, vectors fields, differential forms, etc. defined on a closed subset A of a vector space V. We will always mean by that the these objects are defined on some open neighborhood $U \supset A$. It will be not important for us how exactly these objects are extended to U but to make sense of differentiability we need to assume that they are extended. In fact, one can define what differentiability means without any extension, but this would go beyond the goals of these lecture notes.

Moreover, a theorem of Hassler Whitney asserts that any function smooth on a closed subset $A \subset V$ can be extended to a smooth function to a neighborhood $U \supset A$.

7.3 Gradient vector field

If V is an Euclidean space, i.e. a vector space with an inner product \langle , \rangle , then there exists a canonical isomorphism $\mathcal{D}: V \to V^*$, defined by the formula $\mathcal{D}(v)(x) = \langle v, x \rangle$ for $v, x \in V$. Of course, \mathcal{D} defines an isomorphism $V_x \to V_x^*$ for each $x \in V$. Set

$$\nabla f(x) = \mathcal{D}^{-1}(d_x f).$$

The vector $\nabla f(x)$ is called the *gradient* of the function f at the point $x \in U$. We will also use the notation $\operatorname{grad} f(x)$.

By definition we have

$$\langle \nabla f(x), h \rangle = d_x f(h)$$
 for any vector $h \in V$.

If ||h|| = 1 then $d_x f(h) = ||\nabla f(x)|| \cos \varphi$, where φ is the angle between the vectors $\nabla f(x)$ and h. In particular, the directional derivative $d_x f(h)$ has its maximal value when $\varphi = 0$. Thus the direction of the gradient is the direction of the maximal growth of the function and the length of the gradient equals this maximal value.

As in the case of a differential, the gradient varies from point to point, and the family of vectors $\{\nabla f(x)\}_{x\in U}$ is called the *gradient vector field* ∇f .

We discuss the general notion of a vector field in Section 7.4 below.

7.4 Vector fields

A vector field v on a domain $U \subset V$ is a function which associates to each point $x \in U$ a vector $v(x) \in V_x$, i.e. a vector originated at the point x.

A gradient vector field ∇f of a function f provides us with an example of a vector field, but as we shall see, gradient vector fields form only a small very special class of vector fields.

Let v be a vector field on a domain $U \in V$. If we fix a basis in V, and parallel transport this basis to all spaces $V_x, x \in V$, then for any point $x \in V$ the vector $v(x) \in V_x$ is described by its coordinates $(v_1(x), v_2(x), \dots, v_n(x))$. Therefore, to define a vector field on U is the same as to define n functions v_1, \dots, v_n on U, i.e. to define a map $(v_1, \dots, v_n) : U \to \mathbb{R}^n$. We call a vector field v C^k -smooth if the functions v_1, \dots, v_n are smooth on U.

Thus, if a basis of V is fixed, then the difference between the maps $U \to \mathbb{R}^n$ and vector fields on U is just a matter of geometric interpretation. When we speak about a vector field v we view v(x) as a vector in V_x , i.e. originated at the point $x \in U$. When we speak about a map $v: U \to \mathbb{R}^n$ we view v(x) as a point of the space V, or as a vector with its origin at $\mathbf{0} \in V$.

Vector fields naturally arise in a context of Physics, Mechanics, Hydrodynamics, etc. as force, velocity and other physical fields.

There is another very important interpretation of vector fields as first order differential operators.

Let $C^{\infty}(U)$ denote the vector space of infinitely differentiable functions on a domain $U \subset V$. Let v be a C^{∞} -smooth vector field on V. We associate with v a linear operator

$$D_v: C^{\infty}(U) \to C^{\infty}(U),$$

given by the formula

$$D_v(f) = df(v), f \in C^{\infty}(U).$$

In other words, we compute at any point $x \in U$ the directional derivative of f in the direction of the vector v(x). Clearly, the operator D_v is linear: $D_v(af+bg) = aD_v(f) + bD_v(g)$ for any functions $f, g \in C^{\infty}(U)$ and any real numbers $a, b \in \mathbb{R}$. It also satisfies the *Leibniz rule*:

$$D_v(fg) = D_v(f)g + fD_v(g).$$

In view of the above correspondence between vector fields and first order differential operators it is sometimes convenient just to view a vector field as a differential operator. Hence, when it will not be confusing we may drop the notation D_v and just directly apply the vector v to a function f (i.e. write v(f) instead of $D_v(f)$).

Let v_1, \ldots, v_n be a basis of V, and x_1, \ldots, x_n be the coordinate functions in this basis. We would like to introduce the notation for the vector field obtained from vectors v_1, \ldots, v_n by parallel transporting them to all points of the domain U. To motivate the notation which we are going to introduce, let us temporarily denote these vector fields by $\mathbf{v}_1, \ldots, \mathbf{v}_n$. Observe that $D_{\mathbf{v}_i}(f) = \frac{\partial f}{\partial x_i}$, $i = 1, \ldots, n$. Thus the operator $D_{\mathbf{v}_i}$ is just the operator $\frac{\partial}{\partial x_i}$ of taking i-th partial derivative. Hence, viewing the vector field \mathbf{v}_i as a differential operator we will just use the notation $\frac{\partial}{\partial x_i}$ instead of \mathbf{v}_i . Given any vector field v with coordinate functions $a_1, a_2, \ldots, a_n : U \to \mathbb{R}$ we have

$$D_v(f)(x) = \sum_{i=1}^n a_i(x) \frac{\partial f}{\partial x_i}(x)$$
, for any $f \in C^{\infty}(U)$,

and hence we can write $v = \sum_{i=1}^{n} a_i \frac{\partial}{\partial x_i}$. Note that the coefficients a_i here are functions and not constants.

7.4.1 Gradient vector field

Suppose that V, \langle , \rangle is a Euclidean vector space. Choose a (not necessarily orthonormal) basis v_1, \ldots, v_n . Let us find the coordinate description of the gradient vector field ∇f , i.e. find the coefficients a_j in the expansion $\nabla f(x) = \sum_{1}^{n} a_i(x) \frac{\partial}{\partial x_i}$. By definition we have

$$\langle \nabla f(x), h \rangle = d_x f(h) = \sum_{1}^{n} \frac{\partial f}{\partial x_j}(x) h_j$$
 (7.4.1)

for any vector $h \in V_x$ with coordinates (h_1, \ldots, h_n) in the basis v_1, \ldots, v_n parallel transported to V_x . Let us denote $g_{ij} = \langle v_i, v_j \rangle$. Thus $G = (g_{ij})$ is a symmetric $n \times n$ matrix, which is called the Gram matrix of the basis v_1, \ldots, v_n . Then the equation (7.4.1) can be rewritten as

$$\sum_{i,j=1}^{n} g_{ji} a_i h_j = \sum_{1}^{n} \frac{\partial f}{\partial x_j}(x) h_j.$$

Because h_j are arbitrarily numbers it implies that the coefficients with h_j in the right and left sides should coincide for all j = 1, ..., n. Hence we get the following system of linear equations:

$$\sum_{i=1}^{n} g_{ij} a_i = \frac{\partial f}{\partial x_j}(x), \quad j = 1, \dots, n,$$
(7.4.2)

or in matrix form

$$G\begin{pmatrix} a_1 \\ \vdots \\ a_n \end{pmatrix} = \begin{pmatrix} \frac{\partial f}{\partial x_1}(x) \\ \vdots \\ \frac{\partial f}{\partial x_n}(x) \end{pmatrix},$$

and thus

$$\begin{pmatrix} a_1 \\ \vdots \\ a_n \end{pmatrix} = G^{-1} \begin{pmatrix} \frac{\partial f}{\partial x_1}(x) \\ \vdots \\ \frac{\partial f}{\partial x_n}(x) \end{pmatrix}, \tag{7.4.3}$$

i.e.

$$\nabla f = \sum_{i,j=1}^{n} g^{ij} \frac{\partial f}{\partial x_i}(x) \frac{\partial}{\partial x_j}, \tag{7.4.4}$$

where we denote by g^{ij} the entries of the inverse matrix $G^{-1} = (g_{ij})^{-1}$

If the basis v_1, \ldots, v_n is orthonormal then G is the unit matrix, and thus in this case

$$\nabla f = \sum_{1}^{n} \frac{\partial f}{\partial x_{j}}(x) \frac{\partial}{\partial x_{j}}, \tag{7.4.5}$$

i.e. ∇f has coordinates $(\frac{\partial f}{\partial x_1}, \dots, \frac{\partial f}{\partial x_n})$. However, simple expression (7.4.5) for the gradient holds only in the orthonormal basis. In the general case one has a more complicated expression (7.4.4).

7.5 Differential forms

Similarly to vector fields, we can consider *fields of exterior forms*, i.e. functions on $U \subset V$ which associate to each point $x \in U$ a k-form from $\Lambda^k(V_x^*)$. These fields of exterior k-forms are called differential k-forms.

Thus the relation between k-forms and differential k-forms is exactly the same as the relation between vectors and vector-fields. For instance, a differential 1-form α associates with each point $x \in U$ a linear function $\alpha(x)$ on the space V_x . Sometimes we will write α_x instead of $\alpha(x)$ to leave space for the arguments of the function $\alpha(x)$.

- **Example 7.2.** 1. Let $f: V \to \mathbb{R}$ be a smooth function. Then the differential df is a differential 1-form. Indeed, with each point $x \in V$ it associates a linear function $d_x f$ on the space V_x . As we shall see, most differential 1-form $are\ not$ differentials of functions (just as most vector fields are not gradient vector fields).
- 2. A differential 0-form f on U associates with each point $x \in U$ a 0-form on V_x , i. e. a number $f(x) \in \mathbb{R}$. Thus differential 0-forms on U are just functions $U \to \mathbb{R}$.

7.6 Coordinate description of differential forms

Let x_1, \ldots, x_n be coordinate linear functions on V, which form the basis of V^* dual to a chosen basis v_1, \ldots, v_n of V. For each $i = 1, \ldots, n$ the differential dx_i defines a linear function on each space $V_x, x \in V$. Namely, if $h = (h_1, \ldots, h_n) \in V_x$ then $dx_i(h) = h_i$. Indeed

$$d_x x_i(h) = \lim_{t \to 0} \frac{x_i + th_i - x_i}{t} = h_i,$$

independently of the base point $x \in V$. Thus differentials dx_1, \ldots, dx_n form a basis of the space V_x^* for each $x \in V$. In particular, any differential 1-form α on v can be written as

$$\alpha = f_1 dx_1 + \ldots + f_n dx_n,$$

where f_1, \ldots, f_n are functions on V. In particular,

$$df = \frac{\partial f}{\partial x_1} dx_1 + \ldots + \frac{\partial f}{\partial x_n} dx_n. \tag{7.6.1}$$

Let us point out that this simple expression of the differential of a function holds in an **arbitrary** coordinate system, while an analogous simple expression (7.4.5) for the gradient vector field is valid only in the case of Cartesian coordinates. This reflects the fact that while the notion of differential is intrinsic and independent of any extra choices, one needs to have a background inner product to define the gradient.

Similarly, any differential 2-form w on a 3-dimensional space can be written as

$$\omega = b_1(x)dx_2 \wedge dx_3 + b_2(x)dx_3 \wedge dx_1 + b_3(x)dx_1 \wedge dx_2$$

where b_1, b_2 , and b_3 are functions on V. Any differential 3-form Ω on a 3-dimensional space V has the form

$$\Omega = c(x)dx_1 \wedge dx_2 \wedge dx_3$$

for a function c on V.

More generally, any differential k-form α can be expressed as

$$\alpha = \sum_{1 \le i_1 < i_2 \dots < i_k \le n} a_{i_1 \dots i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}$$

for some functions $a_{i_1...i_k}$ on V.

7.7 Smooth maps and their differentials

Let V, W be two vector spaces of arbitary (not, necessarily, equal) dimensions and $U \subset V$ be an open domain in V.

Recall that a map $f: U \to W$ is called differentiable if for each $x \in U$ there exists a linear map

$$l: V_x \to W_{f(x)}$$

such that

$$l(h) = \lim_{t \to 0} \frac{f(x+th) - f(x)}{t}$$

for any $h \in V_x$. In other words,

$$f(x+th) - f(x) = tl(h) + o(t)$$
, where $\frac{o(t)}{t} \underset{t\to 0}{\to} 0$.

The map l is denoted by $d_x f$ and is called the differential of the map f at the point $x \in U$. Thus, $d_x f$ is a linear map $V_x \to W_{f(x)}$.

The space $W_{f(x)}$ can be identified with W via a parallel transport, and hence sometimes it is convenient to think about the differential as a map $V_x \to W$, In particular, in the case of a linear function. i.e. when $W = \mathbb{R}$ it is customary to do that, and hence we defined earlier in Section 7.1 the differential of a function $f: U \to \mathbb{R}$ at a point $x \in U$ as a linear function $V_x \to \mathbb{R}$, i.e. an element of V_x^* , rather than a linear map $V_x \to W_{f(x)}$.

Let us pick bases in V and W and let (x_1, \ldots, x_k) and (y_1, \ldots, y_n) be the corresponding coordinate functions. Then each of the spaces V_x and W_y , $x \in V$, $y \in W$ inherits a basis obtained by parallel transport of the bases of V and W. In terms of these bases, the differential $d_x f$ is given by the Jacobi matrix

$$\begin{pmatrix} \frac{\partial f_1}{\partial x_1} & \cdots & \frac{\partial f_1}{\partial x_k} \\ \cdots & \cdots & \cdots \\ \frac{\partial f_n}{\partial x_1} & \cdots & \frac{\partial f_n}{\partial x_k} \end{pmatrix}$$

In what follows we will consider only sufficiently smooth maps, i.e. we assume that all maps and their coordinate functions are differentiable as many times as we need it.

7.8 Operator f^*

Let U be a domain in a vector space V and $f:U\to W$ a smooth map. Then the differential df defines a linear map

$$d_x f: V_x \to W_{f(x)}$$

for each $x \in V$.

Let ω be a differential k-form on W. Thus ω defines an exterior k-form on the space W_y for each $y \in W$.

Let us define the differential k-form $f^*\omega$ on U by the formula

$$(f^*\omega)|_{V_x} = (d_x f)^*(\omega|_{W_{f(x)}}).$$

Here the notation $\omega|_{W_y}$ stands for the exterior k-form defined by the differential form ω on the space W_y .

In other words, for any k vectors, $H_1, \ldots, H_k \in V_x$ we have

$$f^*\omega(H_1,\ldots,H_k)=\omega(d_xf(H_1),\ldots,d_xf(H_k)).$$

We say that the differential form $f^*\omega$ is induced from ω by the map f, or that $f^*\omega$ is the pull-back of ω by f.

Example 7.3. Let $\Omega = h(x)dx_1 \wedge \cdots \wedge dx_n$. Then formula (3.2) implies

$$f^*\Omega = h \circ f \det Df dx_1 \wedge \cdots \wedge dx_n$$
.

Here

$$\det Df = \begin{vmatrix} \frac{\partial f_1}{\partial x_1} & \cdots & \frac{\partial f_1}{\partial x_n} \\ \vdots & \ddots & \vdots \\ \frac{\partial f_n}{\partial x_1} & \cdots & \frac{\partial f_n}{\partial x_n} \end{vmatrix}$$

is the determinant of the Jacobian matrix of $f = (f_1, \ldots, f_n)$.

Similarly to Proposition 1.9 we get

Proposition 7.4. Given 2 maps

$$U_1 \stackrel{f}{\to} U_2 \stackrel{g}{\to} U_3$$

and a differential k form ω on U_3 we have

$$(g \circ f)^*(\omega) = f^*(g^*\omega).$$

An important special case of the pull-back operator f^* is the *restriction* operator. Namely Let $L \subset V$ be an affine subspace. Let $j: L \hookrightarrow V$ be the inclusion map. Then given a differential k-form α on a domain $U \subset V$ we can consider the form $j^*\alpha$ on the domain $U' := L \cap U$. This form is called the *restriction* of the form α to U' and it is usually denoted by $\alpha|_U'$. Thus the restricted form $\alpha|_U'$ is the same form α but viewed as function of a point $a \in U'$ and vectors $T_1, \ldots, T_k \in L_a$.

7.9 Coordinate description of the operator f^*

Consider first the linear case. Let \mathcal{A} be a linear map $V \to W$ and $\omega \in \Lambda^p(W^*)$. Let us fix coordinate systems x_1, \ldots, x_k in V and y_1, \ldots, y_n in W. If A is the matrix of the map \mathcal{A} then we already have seen in Section 2.7 that

$$A^*y_j = l_j(x_1, \dots, x_k) = a_{j1}x_1 + a_{j2}x_2 + \dots + a_{jk}x_k, \ j = 1, \dots, n,$$

and that for any exterior k-form

$$\omega = \sum_{1 \le i_1 < \dots < i_p \le n} A_{i_1, \dots, i_p} y_{i_1} \wedge \dots \wedge y_{i_p}$$

we have

$$\mathcal{A}^*\omega = \sum_{1 \le i, < \dots < i_p \le n} A_{i_1 \dots i_p} l_{i_1} \wedge \dots \wedge l_{i_p}.$$

Now consider the non-linear situation. Let ω be a differential p-form on W. Thus it can be written in the form

$$\omega = \sum A_{i_1 \dots i_p}(y) dy_{i_1} \wedge \dots dy_{i_p}$$

for some functions $A_{i_1...i_p}$ on W.

Let U be a domain in V and $f: U \to W$ a smooth map.

Proposition 7.5. $f^*\omega = \sum A_{i_1...,i_p}(f(x))df_{i_1} \wedge ... \wedge df_{i_p}$, where $f_1,...,f_n$ are coordinate functions of the map f.

Proof. For each point $x \in U$ we have, by definition,

$$f^*\omega|_{V_x} = (d_x f)^*(\omega|_{W_{f_x}})$$

But the coordinate functions of the linear map $d_x f$ are just the differentials $d_x f_i$ of the coordinate functions of the map f. Hence the desired formula follows from the linear case proven in the previous proposition.

7.10 Examples

1. Consider the domain $U=\{r>0, 0\leq \varphi<2\pi\}$ on the plane $V=\mathbb{R}^2$ with cartesian coordinates (r,φ) . Let $W=\mathbb{R}^2$ be another copy of \mathbb{R}^2 with cartesian coordinates (x,y). Consider a map $P:V\to W$ given by the formula

$$P(r,\varphi) = (r\cos\varphi, r\sin\varphi).$$

This map introduces (r, φ) as *polar coordinates* on the plane W. Set $\omega = dx \wedge dy$. It is called the area form on W. Then

$$P^*\omega = d(r\cos\varphi) \wedge d(r\sin\varphi) = (\cos\varphi dr + rd(\cos\varphi)) \wedge (\sin\varphi dr + rd(\sin\varphi) =$$

$$(\cos\varphi dr - r\sin\varphi d\varphi) \wedge (\sin\varphi dr + r\cos\varphi d\varphi) =$$

$$\cos\varphi \cdot \sin\varphi dr \wedge dr - r\sin^2\varphi d\varphi \wedge dr + r\cos^2\varphi dr \wedge d\varphi - r^2\sin\varphi\cos\varphi d\varphi \wedge d\varphi =$$

$$r\cos^2\varphi dr \wedge d\varphi + r\sin^2 dr \wedge d\varphi = rdr \wedge d\varphi.$$

2. Let $f: \mathbb{R}^2 \to \mathbb{R}$ be a smooth function and the map $F: \mathbb{R}^2 \to \mathbb{R}^3$ be given by the formula

$$F(x,y) = (x, y, f(x,y))$$

Let

$$\omega = P(x, y, z)dy \wedge dz + Q(x, y, z)dz \wedge dx + R(x, y, z)dx \wedge dy$$



Figure 7.1: Johann Friedrich Pfaff (1765–1825)

be a differential 2-form on \mathbb{R}^3 . Then

$$F^*\omega = P(x, y, f(x, y))dy \wedge df + \\ + Q(x, y, f(x, y))df \wedge dx + R(x, y, f(x, y))dx \wedge dy \\ = P(x, y, f(x, y))dy \wedge (f_x dx + f_y dy) + \\ + Q(x, y, f(x, y))(f_x dx + f_y dy) \wedge dx + \\ + R(x, y, f(x, y))dx \wedge dy = \\ = (R(x, y, f(x, y)) - P(x, y, f(x, y))f_x - Q(x, y, f(x, y))f_y)dx \wedge dy$$

where f_x, f_y are partial derivatives of f.

3. If p > k then the pull-back $f^*\omega$ of a p-form ω on U to a k-dimensional space V is equal to 0.

7.11 Pfaffian equations

Given a non-zero linear function l on an n-dimensional vector space V the equation l = 0 defines a hyperplane, i.e. an (n-1)-dimensional subspace of V.

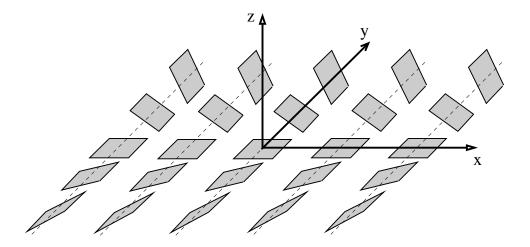


Figure 7.2: Contact structure

Suppose we are given a differential 1-form λ on a domain $U \subset V$. Suppose that $\lambda_x \neq 0$ for each $x \in U$.

Then the equation

$$\lambda = 0 \tag{7.11.1}$$

defines a hyperplane field ξ on U, i.e. a family of hyperplanes $\xi_x = \{\lambda_x = 0\} \subset V_x, \ x \in U$.

The equation of this type is called *Pfaffian* in honor of a German mathematician Johann Friedrich Pfaff (1765–1825).

Example 7.6. Let $V = \mathbb{R}^3$ with coordinates (x, y, z)

1. Let $\lambda = dz$. Then $\xi = \{dz = 0\}$ is the horizontal plane field which is equal to $\mathrm{Span}(\frac{\partial}{\partial x}, \frac{\partial}{\partial y})$. 2. Let $\lambda = dz - ydx$. Then the plane field dz - ydx is shown on Fig. 7.11. This plane is non-integrable in the following sense. There are no surfaces in \mathbb{R}^3 tangent to ξ . This plane field is called a *contact structure*. It plays an important role in symplectic and contact geometry, which is, in turn, the geometric language for Mechanics and Geometric Optics.

Chapter 8

Exterior differential

8.1 Coordinate definition of the exterior differential

Let us denote by by $\Omega^k(U)$ the space of all differential k-forms on U. When we will need to specify the class of smoothness of coefficients of the form we will use the notation $\Omega^k_l(U)$ for the space of differential k-forms with C^l -smooth coefficients, i.e. which depend C^l -smoothly on a point of U.

We will define a map

$$d: \Omega^k(U) \to \Omega^{k+1}(U),$$

or more precisely

$$d: \Omega_l^k(U) \to \Omega_{l-1}^{k+1}(U),$$

(thus assuming that $l \geq 1$), which is called the *exterior differential*. In the current form it was introduced by Élie Cartan, but essentially it was known already to Henri Poincaré.

We first define it in coordinates and then prove that the result is independent of the choice of the coordinate system. Let us fix a coordinate system x_1, \ldots, x_n in $V \supset U$. As a reminder, a differential k-form $w \in \Omega^k(U)$ has the form $w = \sum_{i_1 < \ldots < i_k} a_{i_1 \ldots , i_k} dx_{i_1} \wedge \ldots \wedge dx_{i_k}$ where $a_{i_1 \ldots i_k}$ are functions on the domain U. Define

$$dw := \sum_{i_1 < \dots < i_k} da_{i_1 \dots i_k} \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k}.$$

Examples. 1. Let $w \in \Omega^1(U)$, i.e. $w = \sum_{i=1}^n a_i dx_i$. Then

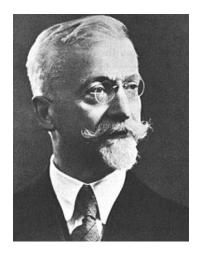


Figure 8.1: Élie Cartan (1869–1951)

$$dw = \sum_{i=1}^{n} da_i \wedge dx_i = \sum_{i=1}^{n} \left(\sum_{j=1}^{n} \frac{\partial a_i}{\partial x_j} dx_j \right) \wedge dx_i = \sum_{1 \le i < j \le n} \left(\frac{\partial a_j}{\partial x_i} - \frac{\partial a_i}{\partial x_j} \right) dx_i \wedge dx_j.$$

For instance, when n=2 we have

$$d(a_1dx_1 + a_2dx_2) = \left(\frac{\partial a_2}{\partial x_1} - \frac{\partial a_1}{\partial x_2}\right) dx_1 \wedge dx_2$$

For n=3, we get

$$d(a_1dx_1 + a_2dx_2 + a_3dx_3) = \left(\frac{\partial a_3}{\partial x_2} - \frac{\partial a_2}{\partial x_3}\right)dx_2 \wedge dx_3 + \left(\frac{\partial a_1}{\partial x_3} - \frac{\partial a_3}{\partial x_1}\right)dx_3 \wedge dx_1 + \left(\frac{\partial a_2}{\partial x_1} - \frac{\partial a_1}{\partial x_2}\right)dx_1 \wedge dx_2.$$

2. Let n=3 and $w\in\Omega^2(U)$. Then

$$w = a_1 dx_2 \wedge dx_3 + a_2 dx_3 \wedge dx_1 + a_3 dx_1 \wedge dx_2$$

and

$$dw = da_1 \wedge dx_2 \wedge dx_3 + da_2 \wedge dx_3 \wedge dx_1 + da_3 \wedge dx_1 \wedge dx_2$$
$$= \left(\frac{\partial a_1}{\partial x_1} + \frac{\partial a_2}{\partial x_2} + \frac{\partial a_3}{\partial x_3}\right) dx_1 \wedge dx_2 \wedge dx_3$$



Figure 8.2: Henri Poincaré (1854–1912)

3. For 0-forms, i.e. functions the exterior differential coincides with the usual differential of a function.

8.2 Properties of the operator d

Proposition 8.1. For any two differential forms, $\alpha \in \Omega^k(U)$, $\beta \in \Omega^l(U)$ we have

$$d(\alpha \wedge \beta) = d\alpha \wedge \beta + (-1)^k \alpha \wedge d\beta.$$

Proof. We have

$$\alpha = \sum_{i_1 < \dots < i_k} a_{i_1 \dots i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}$$

$$\beta = \sum_{j_1 < \dots < j_l} b_{j_1 \dots j_l} dx_{j_1} \wedge \dots \wedge dx_{j_l}$$

$$\alpha \wedge \beta = \left(\sum_{i_1 < \dots < i_k} a_{i_1 \dots i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k} \right) \wedge \left(\sum_{j_i < \dots < j_l} b_{j_1 \dots j_l} dx_{j_1} \wedge \dots \wedge dx_{j_l} \right)$$

$$= \sum_{\substack{i_1 < \dots < i_k \\ j_1 < \dots < j_l}} a_{i_1 \dots i_k} b_{j_1 \dots j_l} dx_{i_l} \wedge \dots \wedge dx_{i_k} \wedge dx_{j_l} \wedge \dots \wedge dx_{j_l}$$

$$d(\alpha \wedge \beta) = \sum_{\substack{i_1 < \dots < i_k \\ j_1 < \dots < j_l}} (b_{j_1 \dots j_l} da_{i_1 \dots i_k} + a_{i_1 \dots i_k} db_{j_1 \dots j_l}) \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k} \wedge dx_{j_1} \wedge \dots \wedge dx_{j_l}$$

$$= \sum_{\substack{i_1 < \dots < i_k \\ j_1 < \dots < j_l}} b_{j_1 \dots j_l} da_{i_1 \dots i_k} \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k} \wedge dx_{j_l} \wedge \dots \wedge dx_{j_l}$$

$$+ \sum_{\substack{i_1 < \dots < i_k \\ j_1 < \dots < j_l}} a_{i_1 \dots i_k} db_{j_1 \dots j_l} \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k} \wedge dx_{j_1} \wedge \dots \wedge dx_{j_l}$$

$$= \left(\sum_{\substack{i_1 < \dots < i_k \\ j_1 < \dots < j_l}} da_{i_1 \dots i_k} \wedge dx_{i_1 \dots} \wedge dx_{i_k} \right) \wedge \left(\sum b_{j_1 \dots j_l} dx_{j_1} \wedge \dots \wedge dx_{j_l} \right)$$

$$+ (-1)^k \left(\sum_{\substack{i_1 < \dots < i_k \\ i_1 < \dots < i_k}} a_{i_1 \dots i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k} \right) \wedge \left(\sum_{\substack{j_1 < \dots < j_l \\ j_1 < \dots < j_l}} db_{j_1 \dots j_l} \wedge dx_{j_1} \wedge \dots \wedge dx_{j_l} \right)$$

$$= d\alpha \wedge \beta + (-1)^k \alpha \wedge d\beta.$$

Notice that the sign $(-1)^k$ appeared because we had to make k transposition to move $db_{j_1...j_l}$ to its place.

Proposition 8.2. For any differential k-form w we have

$$ddw = 0.$$

Proof. Let
$$w = \sum_{i_1 < \dots < i_k} a_{i_1 \dots i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}$$
. Then we have
$$dw = \sum_{i_1 < \dots < i_k} da_{i_1 \dots i_k} \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k}.$$

Applying Proposition 8.1 we get

$$ddw = \sum_{i_1 < \dots < i_k} dda_{i_1 \dots i_k} \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k} - da_{i_1 \dots i_k} \wedge ddx_{i_1} \wedge \dots \wedge dx_{i_k} + \dots$$
$$+ (-1)^k da_{i_1 \dots i_k} \wedge dx_{i_1} \wedge \dots \wedge ddx_{i_k}.$$

But ddf = 0 for any function f as was shown above. Hence all terms in this sum are equal to 0, i.e. ddw = 0.

Definition. A k-form ω is called *closed* if $d\omega = 0$. It is called *exact* if there exists a (k-1)-form θ such that $d\theta = \omega$. The form θ is called the *primitive* of the form ω . The previous theorem can be reformulated as follows:

Corollary 8.3. Every exact form is closed.

The converse is not true in general. For instance, take a differential 1-form

$$\omega = \frac{xdy - ydx}{x^2 + y^2}$$

on the punctured plane $U = \mathbb{R}^2 \setminus 0$ (i.e the plane \mathbb{R}^2 with the deleted origin). It is easy to calculate that $d\omega = 0$, i.e ω is closed. On the other hand it is not exact. Indeed, let us write down this form in polar coordinates (r, φ) . We have

$$x = r \cos \varphi, \ y = r \sin \varphi.$$

Hence,

$$\omega = \frac{1}{r^2} \left(r \cos \varphi (\sin \varphi dr + r \cos \varphi d\varphi) - r \sin \varphi (\cos \varphi dr - r \sin \varphi d\varphi) \right) = d\varphi.$$

If there were a function H on U such that $dH = \omega$, then we would have to have $H = \varphi + \text{const}$, but this is impossible because the polar coordinate φ is not a continuous univalent function on U. Hence ω is not exact.

However, as we will see later, a closed form is exact if it is defined on the whole vector space V.

Proposition 8.4. Operators f^* and d commute, i.e. for any differential k-form $w \in \Omega^k(W)$, and a smooth map $f: U \to W$ we have

$$df^*w = f^*dw$$

.

Proof. Suppose first that k=0, i.e. w is a function $\varphi:W\to\mathbb{R}$. Then $f^*\varphi=\varphi\circ f$. Then $d(\varphi\circ f)=f^*d\varphi$. Indeed, for any point $x\in U$ and a vector $X\in V_x$ we have

$$d(\varphi \circ f)(X) = d\varphi(d_x f(X))$$
(chain rule)

But $d\varphi(d_x f(X)) = f^*(d\varphi(X))$.

Consider now the case of arbitrary k-form w,

$$w = \sum_{i_1 < \dots < i_k} a_{i_1 \dots i_k} dx_{i_1} \wedge \dots \wedge dx_{i_k}.$$

Then

$$f^*w = \sum_{i_1 < \dots < i_k} a_{i_1 \dots i_k} \circ f df_{i_1} \wedge \dots \wedge df_{i_k}$$

where f_1, \ldots, f_n are coordinate functions of the map f. Using the previous theorem and taking into account that $d(df_i) = 0$, we get

$$d(f^*w) = \sum_{i_1 < \dots < i_k} d(a_{i_1 \dots i_k} \circ f) \wedge df_{i_1} \wedge \dots \wedge df_{i_k}.$$

On the other hand

$$dw = \sum_{i_1 < \dots < i_k} da_{i_1 \dots i_k} \wedge dx_{i_1} \wedge \dots \wedge dx_{i_k}$$

and therefore

$$f^*dw = \sum_{i_1 < \dots < i_k} f^*(da_{i_1 \dots i_k}) \wedge df_{i_1} \wedge \dots \wedge df_{i_k}.$$

But according to what is proven above, we have

$$f^*da_{i_1...i_k} = d(a_{i_1...i_k} \circ f)$$

Thus,

$$f^*dw = \sum_{i_1 < \dots < i_k} d(a_{i_1 \dots i_k} \circ f) \wedge df_{i_1} \wedge \dots \wedge df_{i_k} = df^*w$$

The above theorem shows, in particular, that the definition of the exterior differential is independent of the choice of the coordinate. Moreover, one can even use non-linear (curvilinear) coordinate systems, like polar coordinates on the plane.

8.3 Curvilinear coordinate systems

A (non-linear) coordinate system on a domain U in an n-dimensional space V is a smooth map $f = (f_1, \ldots, f_n) : U \to \mathbb{R}^n$ such that

- 1. For each point $x \in U$ the differentials $d_x f_1, \ldots, d_x f_n \in (V_x)^*$ are linearly independent.
- 2. f is injective, i.e. $f(x) \neq f(y)$ for $x \neq y$.

Thus a coordinate map f associates n coordinates $y_1 = f_1(x), \ldots, y_n = f_n(x)$ with each point $x \in U$. The inverse map $f^{-1}: U' \to U$ is called the parameterization. Here $U' = f(U) \subset \mathbb{R}^n$ is the image of U under the map f. If one already has another set of coordinates $x_1 \ldots x_n$ on U, then the coordinate map f expresses new coordinates $y_1 \ldots y_n$ through the old one, while the parametrization map expresses the old coordinate through the new one. Thus the statement

$$g^*dw = dg^*w$$

applied to the parametrization map g just tells us that the formula for the exterior differential is the same in the new coordinates and in the old one.

Consider a space \mathbb{R}^n with coordinates (u_1, \ldots, u_n) . The j-th coordinate line is given by equations $u_i = c_i, i = 1, \ldots, n; i \neq j$. Given a domain $U' \subset \mathbb{R}^n$ consider a parameterization map $g: U' \to U \subset V$. The images $g\{u_i = c_i, i \neq j\} \subset U$ of coordinates lines $\{u_i = c_i, i \neq j\} \subset U'$ are called coordinate lines in U with respect to the curvilinear coordinate system (u_1, \ldots, u_n) . For instance, coordinate lines for polar coordinates in \mathbb{R}^2 are concentric circles and rays, while coordinate lines for spherical coordinates in \mathbb{R}^3 are rays from the origin, and latitudes and meridians on concentric spheres.

8.4 Geometric definition of the exterior differential

We will show later (see Lemma 10.2) that one can give another equivalent definition of the operator d) without using any coordinates at all. But as a first step we give below an equivalent definition of the exterior differential which is manifestly invariant of a choice of affine coordinates.

Given a point $a \in V$ and vectors $h_1, \ldots, h_k \in V_a$ we will denote by $P_a(h_1, \ldots, h_k)$ the k-dimensional parallelepiped

$$\left\{ a + \frac{1}{2} \sum_{1}^{k} u_j h_j; \ |u_1|, \dots, |u_k| \le 1 \right\}$$

centered at the point a with its sides parallel to the vectors h_1, \ldots, h_k , see Fig. ??. For instance, for k = 1 the parallelepiped $P_a(h)$ is an interval centered at a of length |a|.

As it was explained above in Section 8.1 for a 0-form $f \in \Omega^0(U)$, i.e. a function $f : U \to \mathbb{R}$ its exterior differential is just its usual differential, and hence it can be defined by the formula

$$(df)_a(h) = d_a f(h) = \lim_{t \to 0} \frac{1}{t} \left(f\left(a + \frac{th}{2}\right) - f\left(a - \frac{th}{2}\right) \right), t \in \mathbb{R}, h \in V_a, a \in U$$

Proposition 8.5. Let $\alpha \in \Omega_1^{k-1}(U)$ be a differential (k-1)-form on a domain $U \subset \mathbb{R}^n$. Then for any point $a \in U$ and any vectors $h_1, \ldots, h_k \in V_a$ we have

$$(d\alpha)_{a}(h_{1},\ldots,h_{k}) = \lim_{t \to 0} \frac{1}{t^{k}} \sum_{j=1}^{k} (-1)^{k-1} \left(\alpha_{a+\frac{th_{j}}{2}}(th_{1},\ldots,\overset{j}{\vee},\ldots,th_{k}) - \alpha_{a-\frac{th_{j}}{2}}(th_{1},\ldots,\overset{j}{\vee},\ldots,th_{k}) \right)$$

$$= \lim_{t \to 0} \frac{1}{t} \sum_{j=1}^{k} (-1)^{k-1} \left(\alpha_{a+\frac{th_{j}}{2}}(h_{1},\ldots,\overset{j}{\vee},\ldots,h_{k}) - \alpha_{a-\frac{th_{j}}{2}}(h_{1},\ldots,\overset{j}{\vee},\ldots,h_{k}) \right)$$

$$(8.4.1)$$

For instance, for a 1-form $\alpha \in \Omega^1(U)$, we have

$$\begin{split} (d\alpha)_a(h_1,h_2) &:= \lim_{t\to 0} \frac{1}{t^2} \left(\alpha_{a+\frac{th_1}{2}}(th_2) + \alpha_{a-\frac{th_1}{2}}(-th_2) + \alpha_{a+\frac{th_2}{2}}(-th_1) + \alpha_{a-\frac{th_2}{2}}(th_1) \right) \\ &= \lim_{t\to 0} \frac{1}{t} \left(\alpha_{a+\frac{th_1}{2}}(h_2) - \alpha_{a-\frac{th_1}{2}}(h_2) - \alpha_{a+\frac{th_2}{2}}(h_1) + \alpha_{a-\frac{th_2}{2}}(h_1) \right) \end{split}$$

Proof. First we observe that for a C^1 -smooth form α the limit in the right-hand side of formula (8.4.1) exists and define a k-form (i.e. for each a it is a multilinear skewsymmetric function of vectors $h_1, \ldots, h_k \in \mathbb{R}^n_a$). We can also assume that the vectors h_1, \ldots, h_k are linearly independent. Otherwise, both definitions give 0. Denote $L_a = \operatorname{Span}(h_1, \ldots, h_k) \subset V_a$ and consider the restriction $\alpha' := \alpha|_{L \cap U}$. Let (y_1, \ldots, y_k) be coordinates in L_a corresponding to the basis h_1, \ldots, h_k . Then we

can write $\alpha' := \sum P_i dy_1 \wedge \ldots \vee^i \cdots \wedge dy_k$ and thus

$$\lim_{t \to 0} \frac{1}{t} \left(\sum_{j=1}^{k} \alpha_{a + \frac{th_j}{2}}(h_1, \dots, \overset{i}{\vee}, \dots, h_k) - \alpha_{a + \frac{th_j}{2}}(h_1, \dots, \overset{i}{\vee}, \dots, h_k) \right)$$

$$= \sum_{j=1}^{k} \sum_{i=1}^{k} \lim_{t \to 0} \frac{P_j(a + \frac{th_j}{2}) - P_j(a - \frac{th_j}{2})}{t} dy_1 \wedge \dots \overset{i}{\vee} \dots \wedge dy_k(h_1, \dots, \overset{j}{\vee}, \dots, h_k)$$

$$= \sum_{j=1}^{k} \sum_{i=1}^{k} \frac{\partial P_j}{\partial y_j}(a) dy_1 \wedge \dots \overset{i}{\vee} \dots \wedge dy_k(h_1, \dots, \overset{i}{\vee}, \dots, h_k).$$
(8.4.2)

But

$$dy_1 \wedge \dots \overset{i}{\vee} \dots \wedge dy_k(h_1, \dots, \overset{j}{\vee}, \dots, h_k) = \begin{cases} (-1)^{i-1}, & \text{if } i = j \\ 0, & \text{if } i \neq j. \end{cases}$$

Hence, the expression in (8.4.2) is equal to $\sum_{1}^{k} (-1)^{i-1} \frac{\partial P_i}{\partial y_i}(a)$. But

$$d\alpha' = \left(\sum_{1}^{k} (-1)^{i-1} \frac{\partial P_i}{\partial y_i}\right) dy_1 \wedge \dots \wedge dy_k,$$

and therefore $(d\alpha')_a(h_1,\ldots,h_k) = \sum_{1}^{k} (-1)^{i-1} \frac{\partial P_i}{\partial y_i}(a)$, i.e. the two definitions of the exterior differential coincide.

8.5 More about vector fields

Similarly to the case of linear coordinates, given any curvilinear coordinate system (u_1, \ldots, u_n) in U, one denotes by

$$\frac{\partial}{\partial u_1}, \dots, \frac{\partial}{\partial u_n}$$

the vector fields which correspond to the partial derivatives with respect to the coordinates u_1, \ldots, u_n . In other words, the vector field $\frac{\partial}{\partial u_i}$ is tangent to the u_i -coordinate lines and represents the the velocity vector of the curves $u_1 = \text{const}_1, \ldots, u_{i-1} = \text{const}_{i-1}, u_{i+1} = \text{const}_{i+1}, \ldots, u_n = \text{const}_n$, parameterized by the coordinate u_i .

For instance, suppose we are given spherical coordinates $(r, \theta, \varphi), r \geq 0, \phi \in [0, \pi], \theta \in [0, 2\pi)$ in

 \mathbb{R}^3 . The spherical coordinates are related to the cartesian coordinates (x,y,z) by the formulas

$$x = r\sin\varphi\cos\theta,\tag{8.5.1}$$

$$y = r\sin\varphi\sin\theta,\tag{8.5.2}$$

$$z = r\cos\varphi. \tag{8.5.3}$$

Then the vector fields

$$\frac{\partial}{\partial r}$$
, $\frac{\partial}{\partial \varphi}$, and $\frac{\partial}{\partial \theta}$

are mutually orthogonal and define the same orientation of the space as the standard Cartesian coordinates in \mathbb{R}^3 . We also have $||\frac{\partial}{\partial r}|| = 1$. However the length of vector fields $\frac{\partial}{\partial \varphi}$ and $\frac{\partial}{\partial \theta}$ vary. When r and θ are fixed and φ varies, then the corresponding point (r, θ, φ) is moving along a meridian of radius r with a constant angular speed 1. Hence,

$$||\frac{\partial}{\partial \varphi}|| = r.$$

When r and φ are fixed and θ varies, then the point (r, φ, θ) is moving along a latitude of radius $r \sin \varphi$ with a constant angular speed 1. Hence,

$$||\frac{\partial}{\partial \theta}|| = r \sin \varphi.$$

Note that it is customary to introduce unit vector fields in the direction of $\frac{\partial}{\partial r}$, $\frac{\partial}{\partial \theta}$ and $\frac{\partial}{\partial \varphi}$:

$$\mathbf{e}_r := \frac{\partial}{\partial r}, \ \mathbf{e}_{\varphi} := \frac{1}{r} \frac{\partial}{\partial \varphi}, \ \mathbf{e}_{\theta} = \frac{1}{r \sin \varphi} \frac{\partial}{\partial \theta},$$

which form an orthonormal basis at every point. The vector fields \mathbf{e}_r , \mathbf{e}_θ and \mathbf{e}_φ are not defined at the origin and at the poles $\phi = 0, \pi$.

The chain rule allows us to express the vector fields $\frac{\partial}{\partial u_1}, \dots, \frac{\partial}{\partial u_n}$ through the vector fields $\frac{\partial}{\partial x_1}, \dots, \frac{\partial}{\partial x_n}$. Indeed, for any function $f: U \to \mathbb{R}$ we have

$$\frac{\partial f}{\partial u_i} = \sum_{j=1}^n \frac{\partial f}{\partial x_j} \frac{\partial x_j}{\partial u_i},$$

and, therefore,

$$\frac{\partial}{\partial u_i} = \sum_{j=1}^n \frac{\partial x_j}{\partial u_i} \frac{\partial}{\partial x_j},$$

For instance, spherical coordinates (r, θ, ϕ) are related to the cartesian coordinates (x, y, z) by the formulas (8.5.1), and hence we derive the following expression of the vector fields $\frac{\partial}{\partial r}, \frac{\partial}{\partial \theta}, \frac{\partial}{\partial \varphi}$ through the vector fields $\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}$:

$$\begin{split} \frac{\partial}{\partial r} &= \sin \varphi \cos \theta \frac{\partial}{\partial x} + \sin \varphi \sin \theta \frac{\partial}{\partial y} + \cos \varphi \frac{\partial}{\partial z} \,, \\ \frac{\partial}{\partial \theta} &= -r \sin \varphi \sin \theta \frac{\partial}{\partial x} + r \sin \varphi \cos \theta \frac{\partial}{\partial y} \,, \\ \frac{\partial}{\partial \varphi} &= r \cos \varphi \cos \theta \frac{\partial}{\partial x} + r \cos \varphi \sin \theta \frac{\partial}{\partial y} - r \sin \varphi \frac{\partial}{\partial z} \,. \end{split}$$

8.6 Case n = 3. Summary of isomorphisms

Let U be a domain in the 3-dimensional space V. We will consider 5 spaces associated with U.

 $\Omega^0(U) = C^{\infty}(U)$ —the space of 0-forms, i.e. the space of smooth functions;

 $\Omega^k(U)$ for k = 1, 2, 3—the spaces of differential k-forms on U;

Vect(U)—the space of vector fields on U.

Let us fix a volume form $w \in \Omega^3(U)$ that is any nowhere vanishing differential 3-form. In coordinates w can be written as

$$w = f(x)dx_1 \wedge dx_2 \wedge dx_3$$

where the function $f: U \to \mathbb{R}$ is never equal to 0. The choice of the form w allows us to define the following isomorphisms.

1.
$$\Lambda_w: C^{\infty}(U) \to \Omega^3(U)$$
, $\Lambda_w(h) = hw$ for any function $h \in C^{\infty}(U)$.

Sometimes we will omit the subscript w an write just Λ and \bot .

Our third isomorphism depends on a choice of a scalar product <,> in V. Let us fix a scalar product. This enables us to define an isomorphism

$$\mathcal{D} = \mathcal{D}_{<,>} : \mathrm{Vect}(U) \to \Omega^1(U)$$

which associates with a vector field v on U a differential 1-form $\mathcal{D}(v) = \langle v, . \rangle$. Let us write down the coordinate expressions for all these isomorphisms. Fix a cartesian coordinate system (x_1, x_2, x_3) in V so that the scalar product $\langle x, y \rangle$ in these coordinates equals $x_1y_1 + x_2y_2 + x_3y_3$. Suppose also that $w = dx_1 \wedge dx_2 \wedge dx_3$. Then $\Lambda(h) = hdx_1 \wedge dx_2 \wedge dx_3$.

$$\exists (v) = v_1 dx_2 \wedge dx_3 + v_2 dx_3 \wedge dx_1 + v_3 dx_1 \wedge dx_2$$

where v_1, v_2, v_3 are coordinate functions of the vector field v.

$$\mathcal{D}(v) = v_1 dx_1 + v_2 dx_2 + v_3 dx_3.$$

If V is an *oriented* Euclidean space then one also has isomorphisms

$$\star: \Omega^k(V) \to \Omega^{3-k}(V), \ k = 0, 1, 2, 3.$$

If w is the volume form on V for which the unit cube has volume 1 and which define the given orientation of V (equivalently, if $w = x_1 \wedge x_2 \wedge x_3$ for any Cartesian positive coordinate system on V), then

$$\exists_w(v) = \star \mathcal{D}(v), \text{ and } \Lambda_w = \star : \Omega^0(V) \to \Omega^3(V).$$

8.7 Gradient, curl and divergence of a vector field

The above isomorphism, combined with the operation of exterior differentiation, allows us to define the following operations on the vector fields. First recall that for a function $f \in C^{\infty}(U)$,

$$\operatorname{grad} f = \mathcal{D}^{-1}(df).$$

Now let $v \in \text{Vect }(U)$ be a vector field. Then its divergence div v is the function defined by the formula

$$\operatorname{div} v = \Lambda^{-1}(d(\exists v))$$

In other words, we take the 2-form $v \rfloor w$ (w is the volume form) and compute its exterior differential $d(v \rfloor w)$. The result is a 3-form, and, therefore is proportional to the volume form w, i.e. $d(v \rfloor w) = hw$.

This proportionality coefficient (which is a function; it varies from point to point) is simply the divergence: div v = h.

Given a vector field v, its curl is as another vector field curl v defined by the formula

$$\operatorname{curl} v := \operatorname{J}^{-1} d(\mathcal{D}v) = \mathcal{D}^{-1} * d(\mathcal{D}v).$$

If one fixes a cartesian coordinate system in V such that $w = dx_1 \wedge dx_2 \wedge dx_3$ and $\langle x, y \rangle = x_1y_1 + x_2y_2 + x_3y_3$ then we get the following formulas

$$\operatorname{grad} f = \left(\frac{\partial f}{\partial x_1}, \frac{\partial f}{\partial x_2}, \frac{\partial f}{\partial x_3}\right)$$

$$\operatorname{div} v = \frac{\partial v_1}{\partial x_1} + \frac{\partial v_2}{\partial x_2} + \frac{\partial v_3}{\partial x_3}$$

$$\operatorname{curl} v = \left(\frac{\partial v_3}{\partial x_2} - \frac{\partial v_2}{\partial x_3}, \frac{\partial v_1}{\partial x_3} - \frac{\partial v_3}{\partial x_1}, \frac{\partial v_2}{\partial x_1} - \frac{\partial v_1}{\partial x_2}\right)$$

where $v = (v_1, v_2, v_3)$.

We will discuss the geometric meaning of these operations later in Section ??.

8.8 Example: expressing vector analysis operations in spherical coordinates

Given spherical coordinates (r, θ, ϕ) in \mathbb{R}^3 , consider orthonormal vector fields

$$\mathbf{e}_r = \frac{\partial}{\partial r}, \ \mathbf{e}_{\varphi} = \frac{1}{r} \frac{\partial}{\partial \varphi} \ \text{ and } \ \mathbf{e}_{\theta} = \frac{1}{r \sin \varphi} \frac{\partial}{\partial \theta}$$

tangent to the coordinate lines and which form orthonormal basis at every point.

Operator \mathcal{D} .

Given a vector field $\mathbf{v} = a_1 \mathbf{e}_r + a_2 \mathbf{e}_{\varphi} + a_3 \mathbf{e}_{\theta}$ we compute $\mathcal{D} \mathbf{v}$. For any vector field $h = h_1 \mathbf{e}_r + h_2 \mathbf{e}_{\theta} + h_3 \mathbf{e}_{\varphi}$ we have $\langle \mathbf{v}, h \rangle = \mathcal{D} \mathbf{v}(h)$. Let us write $\mathcal{D} \mathbf{v} = c_1 dr + c_2 d\varphi + c_3 d\theta$. Then

$$\langle \mathbf{v}, h \rangle = a_1 h_1 + a_2 h_2 + a_3 h_3$$

and

$$\mathcal{D}\mathbf{v}(h) = (c_1 dr + c_2 d\varphi + c_3 d\theta)(h_1 \mathbf{e}_r + h_2 \mathbf{e}_\varphi + h_3 \mathbf{e}_\theta)$$
$$= c_1 h_1 dr(\mathbf{e}_r) + c_2 h_2 d\varphi(\mathbf{e}_\varphi) + c_3 h_3 d\theta(\mathbf{e}_\theta) = c_1 h_1 + \frac{c_2}{r} h_2 + \frac{c_3}{r \sin \varphi} h_3.$$

Hence, $c_1 = a_1$, $c_2 = ra_2$, $c_3 = ra_3 \sin \varphi$, i.e.

$$\mathcal{D}\mathbf{v} = a_1 dr + r a_2 d\varphi + r a_3 \sin \varphi d\theta. \tag{8.8.1}$$

Operator *. First we express the volume form $\Omega = dx \wedge dy \wedge dz$ in spherical coordinates. One way to do that is just to plug into the form the expression of x, y, z through spherical coordinates. But we can also argue as follows. We have $\Omega = cdr \wedge d\varphi \wedge d\theta$. Let us evaluate both sides of this equality on vectors $\mathbf{e}_r, \mathbf{e}_\varphi, \mathbf{e}_\theta$. Then $\Omega(\mathbf{e}_r, \mathbf{e}_\varphi, \mathbf{e}_\theta) = 1$ because these vectors fields are at every point orthonormal and define the standard orientation of the space. On the other hand,

$$(dr \wedge d\varphi \wedge d\theta)(\mathbf{e}_r, \mathbf{e}_{\varphi}, \mathbf{e}_{\theta})$$

$$= \frac{1}{r^2 \sin \varphi} (dr \wedge d\varphi \wedge d\theta) (\frac{\partial}{\partial r}, \frac{\partial}{\partial \varphi}, \frac{\partial}{\partial \theta}) = \frac{1}{r^2 \sin \varphi},$$

and therefore $c = r^2 \sin \varphi$, i.e.

$$\Omega = r^2 \sin \varphi dr \wedge d\varphi \wedge d\theta.$$

In particular, given a 3-form $\eta = adr \wedge d\varphi \wedge d\theta$ we get $*\eta = \frac{a}{r^2 \sin \varphi}$.

Let us now compute $*dr, *d\varphi$ and $*d\theta$. We argue that

$$*dr = Ad\varphi \wedge d\theta, *d\theta = Bdr \wedge d\varphi, *d\varphi = Cd\theta \wedge dr.$$

Indeed, suppose that $*dr = Ad\varphi \wedge d\theta + A'd\theta \wedge dr + A''dr \wedge d\varphi$ and compute the value of both sides on vectors $\mathbf{e}_{\theta}, \mathbf{e}_{r}$:

$$*dr(\mathbf{e}_{\theta}, \mathbf{e}_r) = dr(\mathbf{e}_{\phi}) = 0,$$

because the 2-dimensional volume (i.e the area) of the 2-dimensional parallelepiped $P(\mathbf{e}_{\theta}, \mathbf{e}_{r})$ and the 1-dimensional volume (i.e. the length) of the 1-dimensional parallelepiped $P(\mathbf{e}\varphi)$ are both equal to 1. On the other hand,

$$(Ad\varphi \wedge d\theta + A'd\theta \wedge dr + A''dr \wedge d\varphi)(\mathbf{e}_{\theta}, \mathbf{e}_{r}) = \frac{1}{r \sin \varphi} A',$$

and therefore A'=0. Similarly, A''=0. The argument for $*d\theta$ and $*d\varphi$ is also similar.

It remains to compute the coefficients A, B, C. We have

$$* dr(\mathbf{e}\varphi, \mathbf{e}_{\theta}) = dr(\mathbf{e}_{r}) = 1 = Ad\varphi \wedge d\theta(\mathbf{e}\varphi, \mathbf{e}_{\theta}) = \frac{A}{r^{2}\sin\varphi},$$

$$* d\theta(\mathbf{e}_{r}, \mathbf{e}_{\varphi}) = d\theta(\mathbf{e}_{\theta}) = \frac{1}{r\sin\varphi} = Bdr \wedge d\varphi(\mathbf{e}_{r}, \mathbf{e}_{\varphi}) = \frac{B}{r},$$

$$* d\varphi(\mathbf{e}_{\theta}, \mathbf{e}_{r}) = d\varphi(\mathbf{e}_{\varphi}) = \frac{1}{r} = Cd\varphi \wedge dr(\mathbf{e}_{\theta}, \mathbf{e}_{r}) = \frac{C}{r\sin\varphi}.$$

Thus, i.e.

$$A = r^2 \sin \varphi, \ B = \frac{1}{\sin \varphi}, \ C = \sin \varphi,$$

$$*dr = r^2 \sin \varphi \, d\varphi \wedge d\theta, \ *d\theta = \frac{1}{\sin \varphi} \, dr \wedge d\varphi, \ *d\varphi = \sin \varphi \, d\theta \wedge dr.$$
 (8.8.2)

Hence we also have

$$*d\varphi \wedge d\theta = \frac{1}{r^2} dr,$$

$$*dr \wedge d\varphi = \sin \varphi d\theta,$$

$$*d\theta \wedge dr = \frac{1}{\sin \varphi} d\varphi.$$

$$(8.8.3)$$

Now we are ready to express the vector analysis operations in spherical coordinates.

Exercise 8.6. (Gradient) Given a function $f(r, \theta, \varphi)$ expressed in cylindrical coordinates, compute ∇f .

Solution. We have

$$\nabla f = D^{-1}(f_r dr + f_{\varphi} d\varphi + f_{\theta} d\theta) = f_r \mathbf{e}_r + \frac{f_{\varphi}}{r} \mathbf{e}_{\varphi} + \frac{f_{\varphi}}{r \sin \varphi} \mathbf{e}_{\theta}.$$

Here we denote by f_r, f_{φ} and f_{θ} the respective partial derivatives of the function f.

Exercise 8.7. Given a vector field $\mathbf{v} = a\mathbf{e}_r + b\mathbf{e}_{\varphi} + c\mathbf{e}_{\theta}$ compute div \mathbf{v} and curl \mathbf{v} .

Solution.

a) **Divergence**. We have

$$\mathbf{v} \, \exists \, \Omega = (a\mathbf{e}_r + b\mathbf{e}_\varphi + c\mathbf{e}_\theta) \, \exists \, r^2 \sin\varphi dr \wedge d\varphi \wedge d\theta$$
$$= ar^2 \sin\varphi \, d\varphi \wedge d\theta + cr \, dr \wedge d\varphi + br \sin\varphi \, d\theta \wedge dr$$

and

$$d(\mathbf{v} \, \exists \, \Omega) = \left((r^2 a_r + 2ra) \sin \varphi + rc_\theta + rb_\varphi \sin \varphi + rb \cos \varphi \right) dr \wedge d\varphi \wedge d\theta$$
$$= \left(a_r + \frac{2a}{r} + \frac{c_\theta}{r \sin \varphi} + \frac{b_\varphi}{r} + \frac{b}{r} \cot \varphi \right) \Omega.$$

Hence,

$$\operatorname{div} \mathbf{v} = a_r + \frac{2a}{r} + \frac{c_\theta}{r\sin\varphi} + \frac{b_\varphi}{r} + \frac{b}{r}\cot\varphi. \tag{8.8.4}$$

b) Curl. We have

$$\operatorname{curl} \mathbf{v} = \mathcal{D}^{-1} * d\mathcal{D} \mathbf{v} = \mathcal{D}^{-1} * d(adr + rbd\varphi + rc\sin\varphi d\theta)$$

$$= \mathcal{D}^{-1} * \left((a_{\varphi} - b - rb_{r})dr \wedge d\varphi + (c\sin\varphi + rc_{r}\sin\varphi - a_{\theta})d\theta \wedge dr \right)$$

$$+ (rb_{\theta} - rc_{\varphi}\sin\varphi - rc\cos\varphi)d\varphi \wedge d\theta)$$

$$= \mathcal{D}^{-1} \left(\frac{rb_{\theta} - rc_{\varphi}\sin\varphi - rc\cos\varphi}{r^{2}}dr + \frac{c\sin\varphi + rc_{r}\sin\varphi - a_{\theta}}{\sin\varphi}d\varphi \right)$$

$$+ (a_{\varphi} - b - rb_{r})\sin\varphi d\theta)$$

$$= \frac{rb_{\theta} - rc_{\varphi}\sin\varphi - rc\cos\varphi}{r^{2}}\mathbf{e}_{r} + \frac{c\sin\varphi + rc_{r}\sin\varphi - a_{\theta}}{r\sin\varphi}\mathbf{e}_{\varphi}$$

$$+ \frac{a_{\varphi} - b - rb_{r}}{r}\mathbf{e}_{\theta}.$$

$$(8.8.5)$$

8.9 Complex-valued differential k-forms

One can consider complex-valued differential k-forms. A \mathbb{C} -valued differential 1-form is a field of \mathbb{C} -valued k-forms, or simply it is an expression $\alpha + i\beta$, where α, β are usual real-valued k-forms. All operations on complex valued k-forms (exterior multiplication, pull-back and exterior differential) are defined in a natural way:

$$(\alpha_1 + i\beta_1) \wedge (\alpha_2 + i\beta_2) = \alpha_1 \wedge \alpha_2 - \beta_1 \wedge \beta_2 + i(\alpha_1 \wedge \beta_2 + \beta_1 \wedge \alpha_2),$$

$$f^*(\alpha + i\beta) = f^*\alpha + if^*\beta, \ d(\alpha + i\beta) = d\alpha + id\beta.$$

We will be in particular interested in complex valued on \mathbb{C} . Note that a complex-valued function (or 0-form) is on a domain $U \subset \mathbb{C}$ is just a map $f = u + iv : U \to \mathbb{C}$. Its differential df is the same as the differential of this map, but it also can be viewed as a \mathbb{C} -valued differential 1-form df = du + idv.

Example 8.8.

$$dz = dx + idy, d\bar{z} = dx - idy, zdz = (x + iy)(dx + idy) = xdx - ydy + i(xdy + ydx),$$
$$dz \wedge d\bar{z} = (dx + idy) \wedge (dx - idy) = -2idx \wedge dy.$$

Exercise 8.9. Prove that $d(z^n) = nz^{n-1}dz$ for any integer $n \neq 0$.

Solution. Let us do the computation in polar coordinates. Then $z^n=r^ne^{in\phi}$ and assuming that $n\neq 0$ we have

$$d(z^n) = nr^{n-1}e^{in\phi}dr + inr^ne^{in\phi}d\phi = nr^{n-1}e^{in\phi}(dr + id\phi).$$

On the other hand,

$$nz^{n-1}dz = nr^{n-1}e^{i(n-1)\phi}d(re^{i\phi}) = nr^{n-1}e^{i(n-1)\phi}\left(e^{i\phi}dr + ie^{i\phi}d\phi\right) = nr^{n-1}e^{in\phi}(dr + id\phi).$$

Comparing the two expressions we conclude that $d(z^n) = nz^{n-1}dz$.

It follows that the 1-form $\frac{dz}{z^n}$ is exact on $\mathbb{C} \setminus 0$ for n > 1. Indeed,

$$\frac{dz}{z^n} = d\left(\frac{1}{(1-n)z^{n-1}}\right).$$

On the other hand the form $\frac{dz}{z}$ is closed on $\mathbb{C} \setminus 0$ but not exact. Indeed,

$$\frac{dz}{z} = \frac{d(re^{i\phi})}{re^{i\phi}} = \frac{e^{i\phi}dr + ire^{i\phi}d\phi}{re^{i\phi}} = \frac{dr}{r} + id\phi,$$

and hence $d\left(\frac{dz}{z}\right) = 0$.

On the other hand, we already had seen that the form $d\phi$, and hence $\frac{dz}{z}$, is not exact.

Chapter 9

Integration of differential forms and functions

9.1 Useful technical tools: partition of unity and cut-off functions

Let us recall that the support of a function θ is the closure of the set of points where it is not equal to 0. We denote the support by $\operatorname{Supp}(\theta)$. We say that θ is supported in an open set U if $\operatorname{Supp}(\theta) \subset U$.

Lemma 9.1. There exists a C^{∞} function $\rho: \mathbb{R} \to [0, \infty)$ with the following properties:

- $\rho(x) \equiv 0, |x| \ge 1;$
- $\bullet \ \rho(x) = \rho(-x);$
- $\rho(x) > 0$ for |x| < 1.

Proof. There are a lot of functions with this property. For instance, one can be constructed as follows. Take the function

$$h(x) = \begin{cases} e^{-\frac{1}{x^2}} &, x > 0\\ 0 &, x \le 0. \end{cases}$$
 (9.1.1)

The function $e^{-\frac{1}{x^2}}$ has the property that all its derivatives at 0 are equal to 0, and hence the function h is C^{∞} -smooth. Then the function $\rho(x) := h(1+x)h(1-x)$ has the required properties.

Lemma 9.2. Existence of cut-off functions Let $C \subset V$ be compact set and $U \supset C$ its open neighborhood. Then there exists a C^{∞} -smooth function $\sigma_{C,U}: V \to [0,\infty)$ with its support in U which is equal to 1 on C

Proof. Let us fix a Euclidean structure in V and a Cartesian coordinate system. Thus we can identify V with \mathbb{R}^n with the standard dot-product. Given a point $a \in V$ and $\delta > 0$ let us denote by $\psi_{a,\delta}$ the *bump* function on V defined by

$$\psi_{a,\delta}(x) := \rho\left(\frac{||x-a||^2}{\delta^2}\right),\tag{9.1.2}$$

where $\rho : \mathbb{R}^{\to}[0,\infty)$ is the function constructed in Lemma 9.1. Note that $\psi_{a,\delta}(x)$ is a C^{∞} -function with Supp $(\psi_{a,\delta}) = D_{\delta} := \overline{B_{\delta}(a)}$ and such that $\psi_{a,\delta}(x) > 0$ for $x \in B_{\delta}(a)$.

Let us denote by $U_{\epsilon}(C)$ the ϵ -neighborhood of C, i.e.

$$U_{\epsilon}(C) = \{x \in V; \exists y \in C, ||y - x|| < \epsilon.\}$$

There exists $\epsilon > 0$ such that $U_{\epsilon}(C) \subset U$. Using compactness of C we can find finitely many points $z_1, \ldots, z_N \in C$ such that the balls $B_{\epsilon}(z_1), \ldots, B_{\epsilon}(z_N) \subset U$ cover $\overline{U_{\frac{\epsilon}{2}}(C)}$, i.e. $\overline{U_{\frac{\epsilon}{2}}(C)} \subset \bigcup_{1}^{N} B_{\epsilon}(z_j)$. Consider a function

$$\sigma_1 := \sum_{1}^{N} \psi_{z_i, \frac{\epsilon}{2}} : V \to \mathbb{R}.$$

The function ψ_1 is positive on $\overline{U}_{\frac{\epsilon}{2}}(C)$ and has $\operatorname{Supp}(\psi_1) \subset U$.

The complement $E = V \setminus U_{\frac{\epsilon}{2}}(C)$ is a closed but unbounded set. Take a large R > 0 such that $B_R(0) \supset \overline{U}$. Then $E_R = D_R(0) \setminus U_{\frac{\epsilon}{2}}(C)$ is compact. Choose finitely many points $x_1, \ldots, x_M \in E_R$ such that $\bigcup_{i=1}^{M} B_{\frac{\epsilon}{4}}(x_i) \supset E_R$. Notice that $\bigcup_{i=1}^{M} B_{\frac{\epsilon}{4}}(x_i) \cap C = \emptyset$. Denote

$$\sigma_2 := \sum_1^M \psi_{x_i, \frac{\epsilon}{4}}.$$

Then the function σ_2 is positive on V_R and vanishes on C. Note that the function $\sigma_1 + \sigma_2$ is positive on $B_R(0)$ and it coincides with σ_1 on C. Finally, define the function $\sigma_{C,U}$ by the formula

$$\sigma_{C,U} := \frac{\sigma_1}{\sigma_1 + \sigma_2}$$

on $B_(R)(0)$ and extend it to the whole space V as equal to 0 outside the ball $B_R(0)$. Then $\sigma_{C,U} = 1$ on C and $\text{Supp}(\sigma_{C,U}) \subset U$, as required.

Let $C \subset V$ be a compact set. Consider its finite covering by open sets U_1, \ldots, U_N , i.e.

$$\bigcup_{1}^{N} U_{j} \supset C.$$

We say that a finite sequence $\theta_1, \ldots, \theta_K$ of C^{∞} -functions defined on some open neighborhood U of C in V forms a partition of unity over C subordinated to the covering $\{U_j\}_{j=1,\ldots,N}$ if

- $\sum_{1}^{K} \theta_j(x) = 1$ for all $x \in C$;
- Each function θ_j , j = 1, ..., K is supported in one of the sets U_i , i = 1, ..., K.

Lemma 9.3. For any compact set C and its open covering $\{U_j\}_{j=1,...,N}$ there exists a partition of unity over C subordinated to this covering.

Proof. In view of compactness of there exists $\epsilon > 0$ and finitely many balls $B_{\epsilon}(z_j)$ centered at points $z_j \in C$, j = 1, ..., K, such that $\bigcup_{1}^{K} B_{\epsilon}(z_j) \supset C$ and each of these balls is contained in one of the open sets U_j , j = 1, ..., N. Consider the functions $\psi_{z_j,\epsilon}$ defined in (9.1.2). We have $\sum_{1}^{K} \psi_{z_j,\epsilon} > 0$ on some neighborhood $U \supset C$. Let $\sigma_{C,U}$ be the cut-off function constructed in Lemma 9.2. For j = 1, ..., K we define

$$\theta_j(x) = \begin{cases} \frac{\psi_{z_j,\epsilon}(x)\sigma_{C,U}(x)}{\frac{K}{1}}, & \text{if } x \in U, \\ \sum_{1}^{K} \psi_{z_j,\epsilon}(x) & & \\ 0, & \text{otherwise} \end{cases}.$$

Each of the functions is supported in one of the open sets U_j , j = 1, ..., N, and we have for every $x \in C$

$$\sum_{1}^{K} \theta_{j}(x) = \frac{\sum_{1}^{K} \psi_{z_{j},\epsilon}(x)}{\sum_{1}^{K} \psi_{z_{j},\epsilon}(x)} = 1.$$

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Figure 9.1: Bernhard Riemann (1826-1866)

9.2 One-dimensional Riemann integral for functions and differential 1-forms

A partition \mathcal{P} of an interval [a,b] is a finite sequence $a=t_0 < t_1 < \cdots < t_N = b$. We will denote by $T_j, j=0,\ldots,N$ the vector $t_{j+1}-t_j \in \mathbb{R}_{t_j}$ and by Δ_j the interval $[t_j,t_{j+1}]$. The length $t_{j+1}-t_j=||T_j||$ of the interval Δ_j will be denoted by δ_j . The number $\max_{j=1,\ldots,N}\delta_j$ is called the fineness or the size of the partition \mathcal{P} and will be denoted by $\delta(\mathcal{P})$. Let us first recall the definition of (Riemann) integral of a function of one variable. Given a function $f:[a,b]\to\mathbb{R}$ we will form a lower and upper integral sums corresponding to the partition \mathcal{P} :

$$L(f; \mathcal{P}) = \sum_{0}^{N-1} (\inf_{[t_j, t_{j+1}]} f)(t_{j+1} - t_j),$$

$$U(f; \mathcal{P}) = \sum_{0}^{N-1} (\sup_{[t_j, t_{j+1}]} f)(t_{j+1} - t_j),$$
(9.2.1)

The function is called *Riemann integrable* if

$$\sup_{\mathcal{P}} L(f; \mathcal{P}) = \inf_{\mathcal{P}} U(f; \mathcal{P}),$$

and in this case this number is called the (Riemann) integral of the function f over the interval [a, b]. The integrability of f can be equivalently reformulated as follows. Let us choose a set C =

 $\{c_1,\ldots,c_{N-1}\},\,c_j\in\Delta_j,\,\text{and consider an integral sum}$

$$I(f; \mathcal{P}, C) = \sum_{j=0}^{N-1} f(c_j)(t_{j+1} - t_j), c_j \in \Delta_j.$$
(9.2.2)

Then the function f is integrable if there exists a limit $\lim_{\delta(\mathcal{P})\to 0} I(f,\mathcal{P},C)$. In this case this limit is equal to the integral of f over the interval [a, b]. Let us emphasize that if we already know that the function is integrable, then to compute the integral one can choose any sequence of integral sum. provided that their fineness goes to 0. In particular, sometimes it is convenient to choose $c_j = t_j$, and in this case we will write $I(f; \mathcal{P})$ instead of $I(f; \mathcal{P}, C)$.

The integral has different notations. It can be denoted sometimes by $\int_{[a,b]} f$, but the most common

notation for this integral is $\int_{a}^{b} f(x)dx$. This notation hints that we are integrating here the differential form f(x)dx rather than a function f. Indeed, given a differential form $\alpha = f(x)dx$ we have $f(c_j)(t_{j+1}-t_j)=\alpha_{c_j}(T_j),^1$ and hence

$$I(\alpha; \mathcal{P}, C) = I(f; \mathcal{P}, C) = \sum_{j=0}^{N-1} \alpha_{c_j}(T_j), c_j \in \Delta_j.$$

$$(9.2.3)$$

We say that a differential 1-form α is integrable if there exists a limit $\lim_{\delta(\mathcal{P})\to 0} I(\alpha, \mathcal{P}, C)$, which is called in this case the integral of the differential 1-form α over the oriented interval [a,b] and will be denoted by $\int_{[a,b]}^{} \alpha$, or simply $\int_{a}^{b} \alpha$. By definition, we say that $\int_{[a,b]}^{} \alpha = -\int_{[a,b]}^{} \alpha$. This agrees with the definition $\int_{[a,b]}^{} \alpha = \lim_{\delta(\mathcal{P})\to 0} \sum_{1}^{N-1} \alpha_{c_{j}}(-T_{j})$, and with the standard calculus rule $\int_{b}^{a} f(x)dx = -\int_{a}^{b} f(x)dx$.

Let us recall that a map $\phi:[a,b]\to[c,d]$ is called a diffeomorphism if it is smooth and has a smooth inverse map $\phi^{-1}:[c,d]\to[a,b]$. This is equivalent to one of the following:

- $\phi(a) = c$; $\phi(b) = d$ and $\phi' > 0$ everywhere on [a, b]. In this case we say that ϕ preserves orientation.
- $\phi(a)=d; \phi(b)=c$ and $\phi'<0$ everywhere on [a,b]. In this case we say that ϕ reverses orientation.

¹Here we parallel transported the vector T_j from the point t_j to the point $c_j \in [t_j, t_{j+1}]$.

Theorem 9.4. Let $\phi: [a,b] \to [c,d]$ be a diffeomorphism. Then if a 1-form $\alpha = f(x)dx$ is integrable over [c,d] then its pull-back $f^*\alpha$ is integrable over [a,b], and we have

$$\int_{\stackrel{\rightarrow}{[a,b]}} \phi^* \alpha = \int_{\stackrel{\rightarrow}{[c,d]}} \alpha, \tag{9.2.4}$$

if ϕ preserves the orientation and

$$\int_{\stackrel{\rightarrow}{[a,b]}} \phi^* \alpha = \int_{\stackrel{\leftarrow}{[c,d]}} \alpha = -\int_{\stackrel{\rightarrow}{[c,d]}} \alpha,$$

if ϕ reverses the orientation.

Remark 9.5. We will show later a stronger result:

$$\int_{\stackrel{\rightarrow}{[a,b]}} \phi^* \alpha = \int_{\stackrel{\rightarrow}{[c,d]}} \alpha$$

for any $\phi:[a,b]\to[c,d]$ with $\phi(a)=c,\phi(b)=d$, which is not necessarily a diffeomorphism.

Proof. We consider only the orientation preserving case, and leave the orientation reversing one to the reader. Choose any partition $\mathcal{P} = \{a = t_0 < \dots < t_{N-1} < t_N = b\}$ of the interval [a, b] and choose any set $C = \{c_0, \dots, c_{N-1}\}$ such that $c_j \in \Delta_j$. Then the points $\widetilde{t}_j = \phi(t_j) \in [c, d], j = 0, \dots N$ form a partition of [c, d]. Denote this partition by $\widetilde{\mathcal{P}}$, and denote $\widetilde{\Delta}_j := [\widetilde{t}_j, \widetilde{t}_{j+1}] \subset [c, d], \widetilde{c}_j = \phi(c_j),$ $\widetilde{C} = \phi(C) = \{\widetilde{c}_0, \dots, \widetilde{c}_{N-1}\}$. Then we have

$$I(\phi^*\alpha, \mathcal{P}, C) = \sum_{0}^{N-1} \phi^*\alpha_{c_j}(T_j) = \sum_{0}^{N-1} \alpha_{\widetilde{c}_j}(d\phi(T_j)) = \sum_{0}^{N-1} \alpha_{\widetilde{c}_j}(\phi'(c_j)\delta_j).$$

$$(9.2.5)$$

Recall that according to the mean value theorem there exists a point $d_j \in \Delta_j$, such that

$$\widetilde{T}_i = \widetilde{t}_{i+1} - \widetilde{t}_i = \phi(t_{i+1}) - \phi(t_i) = \phi'(d_i)(t_{i+1} - t_i).$$

Note also that the function ϕ' is uniformly continuous, i.e. for any $\epsilon > 0$ there exists $\delta > 0$ such that for any $t, t' \in [a, b]$ such that $|t - t'| < \delta$ we have $|\phi'(t) - \phi'(t')| < \epsilon$. Besides, the function ϕ'

is bounded above and below by some positive constants: $m < \phi' < M$. Hence $m\delta_j < \widetilde{\delta}_j < M\delta_j$ for all j = 1, ..., N - 1. Hence, if $\delta(\mathcal{P}) < \delta$ then we have

$$\left| I(\phi^* \alpha, \mathcal{P}, C) - I(\alpha; \widetilde{\mathcal{P}}, \widetilde{C}) \right| = \left| \sum_{j=0}^{N-1} \alpha_{\widetilde{c}_j} \left((\phi'(c_j) - \phi'(d_j)) \delta_j \right) \right| =$$

$$\leq \frac{\epsilon}{m} \left| \sum_{j=0}^{N-1} f(\widetilde{c}_j) \widetilde{\delta}_j \right| = \frac{\epsilon}{m} \left| I(\alpha; \widetilde{\mathcal{P}}, \widetilde{C}) \right|.$$

$$(9.2.6)$$

When $\delta(\mathcal{P}) \to 0$ we have $\widetilde{\delta}(\mathcal{P}) = 0$, and hence by assumption $I(\alpha; \widetilde{\mathcal{P}}, \widetilde{C}) \to \int_{c}^{d} \alpha$, but this implies that $I(\phi^*\alpha, \mathcal{P}, C) - I(\alpha; \widetilde{\mathcal{P}}, \widetilde{C}) \to 0$, and thus $\phi^*\alpha$ is integrable over [a, b] and

$$\int_{a}^{b} \phi^* \alpha = \lim_{\delta(\mathcal{P}) \to 0} I(\phi^* \alpha, \mathcal{P}, C) = \lim_{\delta(\widetilde{\mathcal{P}}) \to 0} I(\alpha, \widetilde{\mathcal{P}}, \widetilde{C}) = \int_{c}^{d} \alpha.$$

If we write $\alpha = f(x)dx$, then $\phi^*\alpha = f(\phi(t))\phi'(t)dt$ and the formula (9.4) takes a familiar form of the change of variables formula from the 1-variable calculus:

$$\int_{c}^{d} f(x)dx = \int_{a}^{b} f(\phi(t))\phi'(t)dt.$$

9.3 Integration of differential 1-forms along curves

Curves as paths

A path, or parametrically given curve in a domain U in a vector space V is a map $\gamma:[a,b]\to U$. We will assume in what follows that all considered paths are differentiable. Given a differential 1-form α in U we define the integral of α over γ by the formula

$$\int_{\gamma} \alpha = \int_{[a,b]} \gamma^* \alpha.$$

Example 9.6. Consider the form $\alpha = dz - ydx + xdy$ on \mathbb{R}^3 . Let $\gamma : [0, 2\pi] \to \mathbb{R}^3$ be a *helix* given by parametric equations $x = R\cos t, y = R\sin t, z = Ct$. Then

$$\int_{\gamma} \alpha = \int_{0}^{2\pi} (Cdt + R^{2}(\sin^{2}tdt + \cos^{2}tdt)) = \int_{0}^{2\pi} (C + R^{2})dt = 2\pi(C + R^{2}).$$

Note that $\int_{\gamma} \alpha = 0$ when $C = -R^2$. One can observe that in this case the curve γ is tangent to the plane field ξ given by the Pfaffian equation $\alpha = 0$.

Proposition 9.7. Let a path $\widetilde{\gamma}$ be obtained from $\gamma:[a,b]\to U$ by a reparameterization, i.e. $\widetilde{\gamma}=\gamma\circ\phi$, where $\phi:[c,d]\to[a,b]$ is an orientation preserving diffeomorphism. Then $\int\limits_{\widetilde{\gamma}}\alpha=\int\limits_{\gamma}\alpha$.

Indeed, applying Theorem 9.4 we get

$$\int_{\widetilde{\gamma}} \alpha = \int_{c}^{d} \widetilde{\gamma}^* \alpha = \int_{c}^{d} \phi^* (\gamma^* \alpha) \int_{a}^{b} \gamma^* \alpha = \int_{\gamma} \alpha.$$

A vector $\gamma'(t) \in V_{\gamma(t)}$ is called the *velocity* vector of the path γ .

Curves as 1-dimensional submanifolds

A subset $\Gamma \subset U$ is called a 1-dimensional submanifold of U if for any point $x \in \Gamma$ there is a neighborhood $U_x \subset U$ and a diffeomorphism $\Phi_x : U_x \to \Omega_x \subset \mathbb{R}^n$, such that $\Phi_x(x) = 0 \in \mathbb{R}^n$ and $\Phi_x(\Gamma \cap U_x)$ either coincides with $\{x_2 = \dots x_n = 0\} \cap \Omega_x$, or with $\{x_2 = \dots x_n = 0, x_1 \geq 0\} \cap \Omega_x$. In the latter case the point x is called a boundary point of Γ . In the former case it is called an interior point of Γ .

A 1-dimensional submanifold is called *closed* if it is compact and has no boundary. An example of a closed 1-dimensional manifold is the circle $S^1 = \{x^2 + y^2 = 1\} \subset \mathbb{R}^2$.

WARNING. The word *closed* is used here in a different sense than when one speaks about closed *subsets*. For instance, a circle in \mathbb{R}^2 is both, a closed subset and a closed 1-dimensional submanifold, while a closed interval is a closed subset but not a closed submanifold: it has 2 boundary points. An open interval in \mathbb{R} (or any \mathbb{R}^n) is a submanifold without boundary but it is not closed because it is not compact. A line in a vector space is a 1-dimensional submanifold which is a closed subset of the ambient vector space. However, it is not compact, and hence not a closed submanifold.

Proposition 9.8. 1. Suppose that a path $\gamma:[a,b]\to U$ is an embedding. This means that $\gamma'(t)\neq 0$ for all $t\in [a,b]$ and $\gamma(t)\neq \gamma(t')$ if $t\neq t'$. Then $\Gamma=\gamma([a,b])$ is 1-dimensional compact submanifold with boundary.

²If only the former property is satisfied that γ is called an *immersion*.

- 2. Suppose $\Gamma \subset U$ is given by equations $F_1 = 0, \ldots, F_{n-1} = 0$ where $F_1, \ldots, F_{n-1} : U \to \mathbb{R}$ are smooth functions such that for each point $x \in \Gamma$ the differential $d_x F_1, \ldots, d_x F_{n-1}$ are linearly independent. Then Γ is a 1-dimensional submanifold of U.
- Any compact connected 1-dimensional submanifold Γ ⊂ U can be parameterized either by an embedding γ : [a, b] → Γ ↔ U if it has non-empty boundary, or by an embedding γ : S¹ → Γ ↔ U if it is closed.

Proof. 1. Take a point $c \in [a, b]$. By assumption $\gamma'(c) \neq 0$. Let us choose an affine coordinate system (y_1, \ldots, y_n) in V centered at the point $C = \gamma(c)$ such that the vector $\gamma'(c) \in V_C$ coincide with the first basic vector. In these coordinates the map gamma can be written as $(\gamma_1, \ldots, \gamma_n)$ where $\gamma'_1(c) = 1$, $\gamma'_j(c) = 0$ for j > 1 and $\gamma_j(c) = 0$ for all $j = 1, \ldots, n$. By the inverse function theorem the function γ_1 is a diffeomorphism of a neighborhood of c onto a neighborhood of c in \mathbb{R} (if c is one of the end points of the interval, then it is a diffeomorphism onto the corresponding one-sided neighborhood of c0). Let c0 be the inverse function defined on the interval c1 equal to c2 equal to c3, c4, c5, c6, c7, c8, and c8, c9, c9, and c9, respectively, depending on whether c1 is an interior point, c2 and c3 equal to the equations:

$$y_2 = \widetilde{\gamma}_2(u) := \gamma_2(\sigma(y_1)), \ldots, y_n = \widetilde{\gamma}_n(u) := \gamma_n(\sigma(y_1)); y_1 \in \Delta.$$

Let us denote

$$\theta = \theta(\delta) := \max_{j=2,\dots,n} \max_{u \in \Delta} |\widetilde{\gamma}_j(u)|.$$

Denote

$$P_{\delta} := \{ |y_1| \leq \delta, |y_j| \leq \theta(\delta) \}.$$

We have $\gamma(\Delta) \subset P_{\delta}$.

We will show now that for a sufficiently small δ we have $\gamma([a,b]) \cap P_{\delta} = \gamma(\Delta)$. For every point $t \in [a,b] \setminus \operatorname{Int} \Delta$ denote $d(t) = ||\gamma(t) - \gamma(c)||$. Recall now the condition that $\gamma(t) \neq \gamma(t')$ for $t \neq t'$. Hence d(t) > 0 for all $t \in [a,b] \setminus \operatorname{Int} \Delta$. The function d(t) is continuous and hence achieve the minimum value on the compact set $[a,b] \setminus \operatorname{Int} \Delta$. Denote $d := \min_{t \in [a,b] \setminus \operatorname{Int} \Delta} d(t) > 0$. Chose $\delta' < \min(d,\delta)$ and such that $\theta(\delta') = \max_{j=2,\dots,n} \max_{|u-c| < \delta'} |\widetilde{\gamma}_j(u)| < d$. Let $\Delta' = \Delta \cap \{|u| \le \delta'\}|$. Then

$$\gamma([a,b]) \cap P_{\delta'} = \{y_2 = \widetilde{\gamma}_2(u), \ldots, y_n = \widetilde{\gamma}_n(u); y_1 \in \Delta'\}.$$

2. Take a point $c \in \Gamma$. The linear independent 1-forms $d_cF_1, \ldots, d_cF_{n-1} \in V_c^*$ can be completed by a 1-form $l \in V_c^*$ to a basis of V_c^* . We can choose an affine coordinate system in V with c as its origin and such that the function x_n coincides with l. Then the Jacobian matrix of the functions $F_1, \ldots, F_{n-1}, x_n$ is non-degenerate at c = 0, and hence by the inverse function theorem the map $F = (F_1, \ldots, F_{n-1}, x_n) : V \to \mathbb{R}^n$ is invertible in the neighborhood of c = 0, and hence these functions can be chosen as new curvilinear coordinates $y_1 = F_1, \ldots, y_{n-1} = F_{n-1}, y_n = x_n$ near the point c = 0. In these coordinates the curve Γ is given near c by the equations $y_1 = \cdots = y_{n-1} = 0$.

3. See Exercise ??.

In the case where Γ is closed we will usually parameterize it by a path $\gamma:[a,b]\to\Gamma\subset U$ with $\gamma(a)=\gamma(b)$. For instance, we parameterize the circle $S^1=\{x^2+y^2=1\}\subset\mathbb{R}^2$ by a path $[0,2\pi]\mapsto(\cos t,\sin t)$. Such γ , of course, cannot be an embedding, but we will require that $\gamma'(t)\neq 0$ and that for $t\neq t'$ we have $\gamma(t)\neq \gamma(t')$ unless one of these points is a and the other one is b. We will refer to 1-dimensional submanifolds simply as curves, respectively closed, with boundary etc.

Given a curve Γ its tangent line at a point $x \in \Gamma$ is a subspace of V_x generated by the velocity vector $\gamma'(t)$ for any local parameterization $\gamma:[a,b] \to \Gamma$ with $\gamma(t)=x$. If Γ is given implicitly, as in 9.8.2, then the tangent line is defined in V_x by the system of linear equations $d_xF_1=0,\ldots,d_xF_{n-1}=0$.

Orientation of a curve Γ is the continuously depending on points orientation of all its tangent lines. If the curve is given as a path $\gamma:[a,b]\to\Gamma\subset U$ such that $\gamma'(t)\neq 0$ for all $t\in[a,b]$ than it is canonically oriented. Indeed, the orientation of its tangent line l_x at a point $x=\gamma(t)\in\Gamma$ is defined by the velocity vector $\gamma'(t)\in l_x$.

It turns out that one can define an integral of a differential form α over an oriented compact curve directly without referring to its parameterization. For simplicity we will restrict our discussion to the case when the form α is continuous.

Let Γ be a compact connected oriented curve. A partition of Γ is a sequence of points $\mathcal{P} = \{z_0, z_1, \ldots, z_N\}$ ordered according to the orientation of the curve and such that the boundary points of the curve (if they exist) are included into this sequence. If Γ is closed we assume that $z_N = z_0$. The fineness $\delta(\mathcal{P})$ of \mathcal{P} is by definition is $\max_{j=0,\ldots,N-1} \operatorname{dist}(z_j, z_{j+1})$ (we assume here that V a Euclidean space).

Definition 9.9. Let α be a differential 1-form and Γ a compact connected oriented curve. Let $\mathcal{P} = \{z_0, \ldots, z_N\}$ be its partition. Then we define

$$\int_{\Gamma} \alpha = \lim_{\delta(\mathcal{P}) \to 0} I(\alpha, \mathcal{P}),$$

where
$$I(\alpha, \mathcal{P}) = \sum_{j=0}^{N-1} \alpha_{z_j}(Z_j), \ \ Z_j = z_{j+1} - z_j \in V_{z_j}.$$

When Γ is a closed submanifold then one sometimes uses the notation $\oint_{\Gamma} \alpha$ instead of $\int_{\Gamma} \alpha$.

Proposition 9.10. If one chooses a parameterization $\gamma:[a,b]\to \Gamma$ which respects the given orientation of Γ then

$$\int\limits_{\gamma} \alpha = \int\limits_{a}^{b} \gamma^* \alpha = \int\limits_{\Gamma} \alpha.$$

Proof. Indeed, let $\widetilde{\mathcal{P}} = \{t_0, \dots, t_N\}$ be a partition of [a, b] such that $\gamma(t_j) = z_j, j = 0, \dots, N$.

$$I(\gamma^* \alpha, \widetilde{\mathcal{P}}) = \sum_{1}^{N-1} \gamma^* \alpha_{t_j}(T_j) = \sum_{1}^{N-1} \alpha_{z_j}(U_j),$$

where $U_j = d_{t_j} \gamma(T_j) \in V_{z_j}$ is a tangent vector to Γ at the point z_j . Let us evaluate the difference $U_j - Z_j$. Choosing some Cartesian coordinates in V we denote by $\gamma_1, \ldots, \gamma_n$ the coordinate functions of the path γ . Then using the mean value theorem for each of the coordinate functions we get $\gamma_i(t_{j+1}) - \gamma_i(t_j) = \gamma_i'(c_j^i)\delta_j$ for some $c_j^i \in \Delta_j$, $i = 1, \ldots, n$; $j = 0, \ldots, N-1$. Thus

$$Z_j = \gamma(t_{j+1}) - \gamma(t_j) = (\gamma'_1(c_j^1), \dots, \gamma'_i(c_j^n))\delta_j.$$

On the other hand, $U_j = d_{t_j} \gamma(T_j) \in V_{z_j} = \gamma'(t_j) \delta_j$. Hence,

$$||Z_j - U_j|| = \delta_j \sqrt{\sum_{1}^{n} (\gamma_i'(c_j^i) - \gamma_i'(t_j))^2}.$$

Note that if $\delta(\mathcal{P}) \to 0$ then we also have $\delta(\widetilde{\mathcal{P}}) \to 0$, and hence using smoothness of the path γ we conclude that for any $\epsilon > 0$ there exists $\delta > 0$ such that $||Z_j - U_j|| < \epsilon \delta_j$ for all $j = 1, \ldots, N$. Thus

$$\sum_{1}^{N-1} \alpha_{z_j}(\widetilde{T}_j) - \sum_{1}^{N-1} \alpha_{z_j}(Z_j) \underset{\delta(\mathcal{P})}{\rightarrow} 0,$$

and therefore

$$\int_{a}^{b} \gamma^* \alpha = \lim_{\delta(\widetilde{\mathcal{P}}) \to 0} I(\gamma^* \alpha, \widetilde{\mathcal{P}}) = \lim_{\delta(\mathcal{P}) \to 0} I(\alpha, \mathcal{P}) = \int_{\Gamma} \alpha.$$

9.4 Integrals of closed and exact differential 1-forms

Theorem 9.11. Let $\alpha = df$ be an exact 1-form in a domain $U \subset V$. Then for any path $\gamma : [a,b] \to U$ which connects points $A = \gamma(a)$ and $B = \gamma(b)$ we have

$$\int_{\gamma} \alpha = f(B) - f(A).$$

In particular, if γ is a loop then $\oint_{\gamma} \alpha = 0$.

Similarly for an oriented curve $\Gamma \subset U$ with boundary $\partial \Gamma = B - A$ we have

$$\int_{\Gamma} \alpha = f(B) - f(A).$$

Proof. We have
$$\int_{\gamma} df = \int_{a}^{b} \gamma^* df = \int_{a}^{b} d(f \circ \gamma) = f(\gamma(b)) - f(\gamma(a)) = f(B) - f(A)$$
.

It turns out that closed forms are locally exact. A domain $U \subset V$ is called star-shaped with respect to a point $a \in V$ if with any point $x \in U$ it contains the whole interval $I_{a,x}$ connecting a and x, i.e. $I_{a,x} = \{a + t(x - a); t \in [0,1]\}$. In particular, any convex domain is star-shaped.

Proposition 9.12. Let α be a closed 1-form in a star-shaped domain $U \subset V$. Then it is exact.

Proof. Define a function $F: U \to \mathbb{R}$ by the formula

$$F(x) = \int_{\overrightarrow{I_{a,x}}} \alpha, \quad x \in U,$$

where the intervals $I_{a,x}$ are oriented from 0 to x.

We claim that $dF = \alpha$. Let us identify V with the \mathbb{R}^n choosing a as the origin a = 0. Then α can be written as $\alpha = \sum_{k=1}^{n} P_k(x) dx_k$, and $I_{0,x}$ can be parameterized by

$$t \mapsto tx, \ t \in [0,1].$$

Hence,

$$F(x) = \int_{\stackrel{\longrightarrow}{I_0}} \alpha = \int_0^1 \sum_{1}^n P_k(tx) x_k dt.$$
 (9.4.1)

Differentiating the integral over x_j as parameters, we get

$$\frac{\partial F}{\partial x_j} = \int_0^1 \sum_{k=1}^n tx_k \frac{\partial P_k}{\partial x_j}(tx)dt + \int_0^1 P_j(tx)dt.$$

But $d\alpha = 0$ implies that $\frac{\partial P_k}{\partial x_j} = \frac{\partial P_j}{\partial x_k}$, and using this we can further write

$$\frac{\partial F}{\partial x_{j}} = \int_{0}^{1} \sum_{k=1}^{n} tx_{k} \frac{\partial P_{j}}{\partial x_{k}}(tx)dt + \int_{0}^{1} P_{j}(tx)dt = \int_{0}^{1} t \frac{dP_{j}(tx)}{dt}dt + \int_{0}^{1} P_{j}(tx)dt$$
$$= (tP_{j}(tx))|_{0}^{1} - \int_{0}^{1} P_{j}(tx)dt + \int_{0}^{1} P_{j}(tx)dt = P_{j}(tx)$$

Thus

$$dF = \sum_{j=1}^{n} \frac{\partial F}{\partial x_j} dx_j = \sum_{j=1}^{n} P_j(x) dx = \alpha$$

9.5 Integration of functions over domains in high-dimensional spaces

Riemann integral over a domain in \mathbb{R}^n .

In this section we will discuss integration of bounded functions over bounded sets in a vector space V. We will fix a basis e_1, \ldots, e_n and the corresponding coordinate system x_1, \ldots, x_n in the space and thus will identify V with \mathbb{R}^n . Let η denote the volume form $x_1 \wedge \ldots x_n$. As it will be clear below, the definition of an integral will not depend on the choice of a coordinate system but only on the background volume form, or rather its absolute value because the orientation of V will be irrelevant.

We will need a special class of parallelepipeds in V, namely those which are generated by vectors proportional to basic vectors, or in other words, parallelepipeds with edges parallel to the coordinate axes. We will also allow these parallelepipeds to be parallel transported anywhere in the space. Let us denote

$$P(a_1, b_1; a_2, b_2; \dots; a_n, b_n) := \{a_i \le x_i \le b_i; i = 1, \dots, n\} \subset \mathbb{R}^n.$$

We will refer to $P(a_1, b_1; a_2, b_2; ...; a_n, b_n)$ as a special parallelepiped, or rectangle.

Let us fix one rectangle $P := P(a_1, b_1; a_2, b_2; \ldots; a_n, b_n)$. Following the same scheme as we used in the 1-dimensional case, we define a partition \mathcal{P} of P as a product of partitions $a_1 = t_0^1 < \cdots < t_{N_1}^1 = b_1, \ldots, a_n = t_0^n < \cdots < t_{N_n}^n = b_n$, of intervals $[a_1, b_1], \ldots, [a_n, b_n]$. For simplicity of notation we will always assume that each of the coordinate intervals is partitioned into the same number of intervals, i.e. $N_1 = \cdots = N_n = N$. This defines a partition of P into N^n smaller rectangles $P_{\mathbf{j}} = \{t_{j_1}^1 \leq x_1 \leq t_{j_1+1}^1, \ldots, t_{j_n}^n \leq x_n \leq t_{j_n+1}^n\}$, where $\mathbf{j} = (j_1, \ldots, j_n)$ and each index j_k takes values between 0 and N-1. Let us define

$$Vol(P_{\mathbf{j}}) := \prod_{k=1}^{n} (t_{j_k+1}^k - t_{j_k}^k).$$
(9.5.1)

This agrees with the definition of the volume of a parallelepiped which we introduced earlier (see formula (3.3.1) in Section 3.3). We will also denote $\delta_{\mathbf{j}} := \max_{k=1,\dots,n} (t_{j_k+1}^k - t_{j_k}^k)$ and $\delta(\mathcal{P}) := \max_{\mathbf{j}} (\delta_{\mathbf{j}})$. Let us fix a point $c_{\mathbf{j}} \in P_{\mathbf{j}}$ and denote by C the set of all such $c_{\mathbf{j}}$. Given a function $f: P \to \mathbb{R}$ we form an integral sum

$$I(f; \mathcal{P}, C) = \sum_{\mathbf{j}} f(c_{\mathbf{j}}) \operatorname{Vol}(P_{\mathbf{j}})$$
(9.5.2)

where the sum is taken over all elements of the partition. If there exists a limit $\lim_{\sigma(\mathcal{P})\to 0} I(f;\mathcal{P},C)$ then the function $f:P\to\mathbb{R}$ is called integrable (in the sense of Riemann) over P, and this limit is called the integral of f over P. There exist several different notations for this integral: $\int_P f, \int_P f dV$, $\int_P f dV$ ol, etc. In the particular case of n=2 one often uses notation $\int_P f dA$, or $\int_P f dA$. Sometime, the functions we integrate may depend on a parameter, and in these cases it is important to indicate with respect to which variable we integrate. Hence, one also uses the notation like $\int_P f(x,y) dx^n$, where the index n refers to the dimension of the space over which we integrate. One also use the notation $\int_P f(x_1, \dots, x_n) dx_1 \dots dx_n$, which is reminiscent both of the integral $\int_P f(x_1, \dots, x_n) dx_1 \wedge \dots \wedge \int_P f(x_1, \dots, x_n) dx_1 \dots dx_n$, which is reminiscent both of the integral $\int_P f(x_1, \dots, x_n) dx_1 \wedge \dots \wedge \int_P f(x_n, \dots, x_n) dx_1 \dots dx_n$

 dx_n which will be defined later in Section 9.7 and the notation $\int_{a_n}^{b_n} \dots \int_{a_1}^{b_1} f(x_1, \dots x_n) dx_1 \dots dx_n$ for n interated integral which will be discussed in Section 9.6.

Alternatively and equivalently the integrability can be defined via upper and lower integral sum,

similar to the 1-dimensional case. Namely, we define

$$U(f; \mathcal{P}) = \sum_{\mathbf{j}} M_{\mathbf{j}}(f) \operatorname{Vol}(P_{\mathbf{j}}), \ L(f; \mathcal{P}) = \sum_{\mathbf{j}} m_{\mathbf{j}}(f) \operatorname{Vol}(P_{\mathbf{j}}),$$

where $M_{\mathbf{j}}(f) = \sup_{P_{\mathbf{j}}} f, m_{\mathbf{j}}(f) = \inf_{P_{\mathbf{j}}} f$, and say that the function f is integrable over P if $\inf_{\mathcal{P}} U(f; \mathcal{P}) = \sup_{\mathcal{P}} L(f, \mathcal{P})$.

Note that $\inf_{\mathcal{P}} U(f; \mathcal{P})$ and $\sup_{\mathcal{P}} L(f, \mathcal{P})$ are sometimes called upper and lower integrals, respectively, and denoted by $\overline{\int_P} f$ and $\int_{\overline{P}} f$. Thus a function $f: P \to \mathbb{R}$ is integrable iff $\overline{\int_P} f = \int_{\overline{P}} f$. Let us list some properties of Riemann integrable functions and integrals.

Proposition 9.13. Let $f, g : P \to \mathbb{R}$ be integrable functions. Then

- 1. af + bg, where $a, b \in \mathbb{R}$, is integrable and $\int_{P} af + bg = a \int_{P} f + b \int_{P} g$;
- 2. If $f \leq g$ then $\int_{P} f \leq \int_{P} g$;
- 3. $h = \max(f, g)$ is integrable; in particular the functions $f_+ := \max(f, 0)$ and $f_- := \max(-f, 0)$ and $|f| = f_+ + f_-$ are integrable;
- 4. fg is integrable.

Proof. Parts 1 and 2 are straightforward and we leave them to the reader as an exercise. Let us check properties 3 and 4.

3. Take any partition \mathcal{P} of P. Note that

$$M_{\mathbf{j}}(h) - m_{\mathbf{j}}(h) \le \max(M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f), M_{\mathbf{j}}(g) - m_{\mathbf{j}}(g)).$$
 (9.5.3)

Indeed, we have $M_{\mathbf{j}}(h) = \max(M_{\mathbf{j}}(f), M_{\mathbf{j}}(g))$ and $m_{\mathbf{j}}(h) \ge \max(m_{\mathbf{j}}(f), m_{\mathbf{j}}(g))$. Suppose for determinacy that $\max(M_{\mathbf{j}}(f), M_{\mathbf{j}}(g)) = M_{\mathbf{j}}(f)$. We also have $m_{\mathbf{j}}(h) \ge m_{\mathbf{j}}(f)$. Thus

$$M_{\mathbf{j}}(h) - m_{\mathbf{j}}(h) \leq M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f) \leq \max(M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f), M_{\mathbf{j}}(g) - m_{\mathbf{j}}(g)).$$

Then using (9.5.3) we have

$$U(h; \mathcal{P}) - L(h; \mathcal{P}) = \sum_{\mathbf{j}} (M_{\mathbf{j}}(h) - m_{\mathbf{j}}(h)) \operatorname{Vol}(P_{\mathbf{j}}) \le$$

$$\sum_{\mathbf{j}} \max (M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f), M_{\mathbf{j}}(g) - m_{\mathbf{j}}(g)) \operatorname{Vol}(P_{\mathbf{j}}) =$$

$$\max (U(f; \mathcal{P}) - L(f; \mathcal{P}), U(f; \mathcal{P}) - L(f; \mathcal{P})).$$

By assumption the right-hand side can be made arbitrarily small for an appropriate choice of the partition \mathcal{P} , and hence h is integrable.

4. We have $f = f_+ - f_-$, $g = g_+ - g_-$ and $fg = f_+g_+ + f_-g_- - f_+g_- - f_-g_+$. Hence, using 1 and 3 we can assume that the functions f, g are non-negative. Let us recall that the functions f, g are by assumption bounded, i.e. there exists a constant C > 0 such that $f, g \leq C$. We also have $M_{\mathbf{j}}(fg) \leq M_{\mathbf{j}}(f)M_{\mathbf{j}}(g)$ and $m_{\mathbf{j}}(fg) \geq m_{\mathbf{j}}(f)m_{\mathbf{j}}(g)$. Hence

$$U(fg; \mathcal{P}) - L(fg; \mathcal{P}) = \sum_{\mathbf{j}} (M_{\mathbf{j}}(fg) - m_{\mathbf{j}}(fg)) \operatorname{Vol}(P_{\mathbf{j}}) \leq$$

$$\sum_{\mathbf{j}} (M_{\mathbf{j}}(f)M_{\mathbf{j}}(g) - m_{\mathbf{j}}(f)m_{\mathbf{j}}(g)) \operatorname{Vol}(P_{\mathbf{j}}) =$$

$$\sum_{\mathbf{j}} (M_{\mathbf{j}}(f)M_{\mathbf{j}}(g) - m_{\mathbf{j}}(f)M_{\mathbf{j}}(g) + m_{\mathbf{j}}(f)M_{\mathbf{j}}(g) - m_{\mathbf{j}}(f)m_{\mathbf{j}}(g)) \operatorname{Vol}(P_{\mathbf{j}}) \leq$$

$$\sum_{\mathbf{j}} ((M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f))M_{\mathbf{j}}(g) + m_{\mathbf{j}}(f)(M_{\mathbf{j}}(g) - m_{\mathbf{j}}(g))) \operatorname{Vol}(P_{\mathbf{j}}) \leq$$

$$C(U(f; \mathcal{P}) - L(f; \mathcal{P}) + U(g; \mathcal{P}) - L(g; \mathcal{P})).$$

By assumption the right-hand side can be made arbitrarily small for an appropriate choice of the partition \mathcal{P} , and hence fg is integrable.

Consider now a bounded subset $K \subset \mathbb{R}^n$ and choose a rectangle $P \supset K$. Given any function $f: K \to \mathbb{R}$ one can always extend it to P as equal to 0. A function $f: K \to \mathbb{R}$ is called *integrable* over K if this trivial extension \overline{f} is integrable over P, and we define $\int_K f dV := \int \overline{f} dV$. When this will not be confusing we will usually keep the notation f for the above extension.

Volume

We further define the volume

$$Vol(K) = \int_{K} 1 dV = \int_{P} \chi_{K} dV,$$

provided that this integral exists. In this case we call the set K measurable in the sense of Riemann, or just measurable.³ Here χ_K is the characteristic or indicator function of K, i.e. the function which

³There exists a more general and more common notion of measurability in the sense of Lebesgue. Any Riemann measurable set is also measurable in the sense of Lebesgue, but not the other way around. Historically an attribution

is equal to 1 on K and 0 elsewhere. In the 2-dimensional case the volume is called the area, and in the 1-dimensional case the length.

Remark 9.14. For any bounded set A there is defined a *lower* and *upper* volumes,

$$\underline{\operatorname{Vol}}(A) = \int \chi_A dV \le \overline{\operatorname{Vol}}(A) = \overline{\int} \chi_A dV.$$

The set is measurable iff $\underline{\mathrm{Vol}}(A) = \overline{\mathrm{Vol}}(A)$. If $\overline{\mathrm{Vol}}(A) = 0$ then $\underline{\mathrm{Vol}}(A) = 0$, and hence A is measurable and $\mathrm{Vol}(A) = 0$.

Exercise 9.15. Prove that for the rectangles this definition of the volume coincides with the one given by the formula (9.5.1).

The next proposition lists some properties of the volume.

Proposition 9.16. 1. Volume is monotone, i.e. if $A, B \subset P$ are measurable and $A \subset B$ then $\operatorname{Vol}(A) \leq \operatorname{Vol}(B)$.

2. If sets $A, B \subset P$ are measurable then $A \cap B$, $A \setminus B$ and $A \cup B$ are measurable as well and we have

$$Vol(A \cup B) = Vol(A) + Vol(B) - Vol(A \cap B).$$

- 3. If A can be covered by a measurable set of arbitrarily small total volume then Vol(A) = 0. Conversely, if Vol(A) = 0 then for any ε > 0 there exists a δ > 0 such that for any partition P with δ(P) < δ the elements of the partition which intersect A have arbitrarily small total volume.</p>
- 4. A is measurable iff $Vol(\partial A) = 0$.

Proof. The first statement is obvious. To prove the second one, we observe that $\chi_{A \cup B} = \max(\chi_A, \chi_B)$, $\chi_{A \cap B} = \chi_A \chi_B$, $\max(\chi_A, \chi_B) = \chi_A + \chi_B - \chi_A \chi_B$, $\chi_{A \setminus B} = \chi_A - \chi_{A \cap B}$ and then apply Proposition 9.13. To prove 9.16.3 we first observe that if a set B is measurable and $\operatorname{Vol} B < \epsilon$ then then for a sufficiently fine partition \mathcal{P} we have $U(\chi_B; \mathcal{P}) < \operatorname{Vol} B + \epsilon < 2\epsilon$. Since $A \subset B$ then $\chi_A \leq \chi_B$, and of this notion to Riemann is incorrect. It was defined by Camille Jordan and Giuseppe Peano before Riemann integral was introduced. What we call in these notes *volume* is also known by the name *Jordan content*.

therefore $U(\chi_A, \mathcal{P}) \leq U(\chi_B, \mathcal{P}) < 2\epsilon$. Thus, $\inf_{\mathcal{P}} U(\chi_A, \mathcal{P}) = 0$ and therefore A is measurable and $\operatorname{Vol}(A) = 0$. Conversely, if $\operatorname{Vol}(A) = 0$ then for any $\epsilon > 0$ for a sufficiently fine partition \mathcal{P} we have $U(\chi_A; \mathcal{P}) < \epsilon$. But $U(\chi_A; \mathcal{P})$ is equal to the sum of volumes of elements of the partition which have non-empty intersection with A.

Finally, let us prove 9.16.4. Consider any partition \mathcal{P} of P and form lower and upper integral sums for χ_A . Denote $M_{\mathbf{j}} := M_{\mathbf{j}}(\chi_A)$ and $m_{\mathbf{j}} = m_{\mathbf{j}}(\chi_A)$. Then all numbers $M_{\mathbf{j}}, m_{\mathbf{j}}$ are equal to either 0 or 1. We have $M_{\mathbf{j}} = m_{\mathbf{j}} = 1$ if $P_{\mathbf{j}} \subset A$; $M_{\mathbf{j}} = m_{\mathbf{j}} = 0$ if $P_{\mathbf{j}} \cap A = \emptyset$ and $M_{\mathbf{j}} = 1$, $m_{\mathbf{j}} = 0$ if $P_{\mathbf{j}}$ has non-empty intersection with both A and $P \setminus A$. In particular,

$$B(\mathcal{P}) := \bigcup_{\mathbf{j}; M_{\mathbf{j}} - m_{\mathbf{i}} = 1} P_{\mathbf{j}} \supset \partial A.$$

Hence, we have

$$U(\chi_A; \mathcal{P}) - L(\chi_A; \mathcal{P}) = \sum_{\mathbf{j}} (M_{\mathbf{j}} - m_{\mathbf{j}}) \operatorname{Vol}(P_{\mathbf{j}}) = \operatorname{Vol}B(\mathcal{P}).$$

Suppose that A is measurable. Then there exists a partition such that $U(\chi_A; \mathcal{P}) - L(\chi_A; \mathcal{P}) < \epsilon$, and hence ∂A is can be covered by the set $B(\mathcal{P})$ of volume $< \epsilon$. Thus applying part 3 we conclude that $\operatorname{Vol}(\partial A) = 0$. Conversely, we had seen below that if $\operatorname{Vol}(\partial A) = 0$ then there exists a partition such that the total volume of the elements intersecting ∂A is $< \epsilon$. Hence, for this partition we have $L(\chi_A; \mathcal{P}) \le U(\chi_A; \mathcal{P}) < \epsilon$, which implies the integrability of χ_A , and hence measurability of A.

Corollary 9.17. If a bounded set $A \subset V$ is measurable then its interior \overline{A} and its closure \overline{A} are also measurable and we have in this case

$$Vol A = Vol Int A = Vol \overline{A}$$
.

Proof. 1. We have $\partial \overline{A} \subset \partial A$ and $\partial(\operatorname{Int} A) \subset \partial A$. Therefore, $\operatorname{Vol} \partial \overline{A} = \operatorname{Vol} \partial \operatorname{Int} A = 0$, and therefore the sets \overline{A} and $\operatorname{Int} A$ are measurable. Also $\operatorname{Int} A \cup \partial A = \overline{A}$ and $\operatorname{Int} A \cap \partial A = \emptyset$. Hence, the additivity of the volume implies that $\operatorname{Vol} \overline{A} = \operatorname{Vol} \partial \operatorname{Int} A + \operatorname{Vol} \partial A = \operatorname{Vol} \partial \operatorname{Int} A$. On the other hand, $\operatorname{Int} A \subset A \subset \overline{A}$, and hence the monotonicity of the volume implies that $\operatorname{Vol} A \leq \operatorname{Vol} \overline{A}$. Hence, $\operatorname{Vol} A = \operatorname{Vol} \operatorname{Int} A = \operatorname{Vol} \overline{A}$.

Exercise 9.18. If Int A or \overline{A} are measurable then this does not imply that A is measurable. For instance, if A is the set of rational points in interval $I = [0,1] \subset \mathbb{R}$ then Int $A = \emptyset$ and $\overline{A} = I$. However, show that A is not Riemann measurable.

2. A set A is called *nowhere dense* if Int $A = \emptyset$. Prove that if A is nowhere dense then either VolA = 0, or A is not measurable in the sense of Riemann. Find an example of a non-measurable nowhere dense set.

Lipshitz maps

Let V, W be two Euclidean spaces. Recall the definition of the norm of a linear operator $\mathcal{A}: V \to W$:

$$||\mathcal{A}|| = \max_{||x||=1} \frac{||\mathcal{A}(x)||}{||x||}.$$

Thus we have $\mathcal{A}(B_R(0)) \subset B_{aR}(0) \subset W$, where we denoted $a := ||\mathcal{A}||$.

Given a subset $A \subset V$ a map $f: A \to W$ is called Lipshitz if there exists a constant C > 0 such that for any $x, y \in A$ we have

$$||f(y) - f(x)|| \le C||y - x||.$$

Lemma 9.19. Let $A \subset V$ be a compact set. Then any C^1 -smooth map $A \to W$ is Lipshitz.

Let us recall that given a compact set $C \subset V$, we say that a map $f: C \to W$ is smooth if it extends to a smooth map defined on an open neighborhood $U \supset C$.

Proof. Let $K: A \to \mathbb{R}$ be the function defined by $K(x) = ||d_x f||, x \in A$. The function K is continuous because f is C^1 -smooth. Hence it is bounded: there exists a constant E > 0 such that $K(x) \leq E$ for all $x \in A$.

Let us first consider the case when A is a convex set, i.e. with any two points $x, y \in A$ the interval connecting them is also contained in A. Given two points $x, y \in A$ at a distance d = ||x - y|| > 0 consider a path $\phi(s) = x + \frac{s}{d}(y - x)$, $s \in [0, d]$ which connects them. Note that the velocity vector $\phi'(s) = \frac{y-x}{d}$ has the unit length. Denote $\tilde{f}(s) = f(\phi(s))$, $s \in [0, d]$. Note that $\tilde{f}'(s) = df_{\phi(s)}(\phi'(s))$ by the chain rule. In particular,

$$||\widetilde{f}'(s)|| \le ||df_{\phi(s)}||||\phi'(s)|| \le E.$$

We also have $f(y) - f(x) = \widetilde{f}(d) - \widetilde{f}(0)$. Let $\widetilde{f} = (\widetilde{f}_1, \dots, \widetilde{f}_m)$ be the coordinate functions of \widetilde{f} in some Cartesian coordinates in W. By the intermediate value theorem, for each $k = 1, \dots, n$ we

have $\widetilde{f}_k(d) - \widetilde{f}_k(0) = \widetilde{f}'_k(c_k)d$ for some $c_k \in [0, d]$. Hence, $|\widetilde{f}_k(d) - \widetilde{f}_k(0)| \leq Cd$ and therefore

$$||f(y) - f(x)|| = ||\widetilde{f}(d) - \widetilde{f}(0)|| = \sqrt{\sum_{1}^{n} (\widetilde{f}_{k}(d) - \widetilde{f}_{k}(0))^{2}} \le E\sqrt{n}d = \widetilde{E}||y - x||, \tag{9.5.4}$$

where we denoted $\widetilde{E} := E\sqrt{n}$.

For a general bounded set A choose an open neighborhood $U \supset A$ to which the map f extends C^1 -smoothly. Let $U' \subset U$ be a smaller open neighborhood of A such that $\overline{U}' \subset U$. We will also assume that \overline{U}' is bounded, and hence compact.

For every point $x \in A$ there is $\epsilon(x) > 0$ such that the closed ball $\overline{B_{\epsilon(x)}(x)}$ is contained in U'. We have $\bigcup_{x \in A} B_{\frac{\epsilon(x)}{2}}(x) \supset A$, and hence by compactness of A there are finitely many balls $B_i := B_{\frac{\epsilon(x_i)}{2}}(x_i)$, $i = 1, \ldots, N$, such that $\bigcup_{1}^{N} B_i \supset A$. Denote $\epsilon := \min_{i=1,\ldots,N} \frac{\epsilon(x_i)}{2}$. Then for any two points $x, y \in A$ with $||y - x|| \le \epsilon$ belong to one of the balls $\widehat{B}_i := B_{\epsilon(x_i)}(x_i)$ which is convex, and hence according to (9.5.4) we have $||f(y) - f(x)|| \le \widetilde{E}||y - x||$.

Denote by D the diameter of the compact set f(A), i.e. $D:=\max_{x,y\in A}||f(y)-f(x)||$. Then if for the distance between two points $x,y\in A$ is $\geq \epsilon$ we have $||f(y)-f(x)||\leq \frac{D}{\epsilon}||y-x||$. Finally, if we denote $C:=\max(\widetilde{E},\frac{D}{\epsilon})$ we get

$$||f(y) - f(x)|| \le C||y - x||$$
 for any $x, y \in A$.

Volume and smooth maps

Lemma 9.20. Let $A \subset V$ be a compact set of volume 0 and $f: V \to W$ a Lipshitz map, where $\dim W \ge \dim V$. Then $\operatorname{Vol} f(A) = 0$.

Proof. According to Lemma 9.19 there is a constant C > 0 such that $||f(y) - f(x)|| \le C||y - x||$ for any $x, y \in A$. In particular, the image f(P) of a cube P of size δ in V_x , $x \in A$, centered at x, is contained in a cube of size $K\delta$ in $W_{f(x)}$ centered at f(x).

The volume 0 assumption implies that for any $\epsilon > 0$ there exists a partition of some size $\delta > 0$ of a larger cube containing A such that the total volume $N\delta^n$ of cubes P_1, \ldots, P_N intersecting A

is $<\epsilon$. But then $\bigcup_{1}^{N} f(P_j) \supset f(A)$ while $\overline{\mathrm{Vol}}_m(f(P_j)) \leq K^M \delta^m$, and hence

$$\overline{\mathrm{Vol}}_m f(A) \leq N \delta^n K^m \delta^{m-n} \leq \epsilon K^m \delta^{m-n} \underset{\epsilon \to 0}{to} \to 0.$$

Corollary 9.21. Let $A \subset V$ be any compact set and $f : A \to W$ a C^1 -smooth map. Suppose that $n = \dim V < m = \dim W$. Then $\operatorname{Vol}(f(A)) = 0$.

Indeed, f can be extended to a smooth map defined on a neighborhood of $A \times 0$ in $V \times \mathbb{R}$ (e.g. as independent of the new coordinate $t \in \mathbb{R}$). But $\operatorname{Vol}_{n+1}(A \times 0) = 0$ and $m \geq n+1$. Hence, the required statement follows from Lemma 9.20.

Remark 9.22. The statement of Corollary 9.21 is wrong for continuous maps. For instance, there exists a continuous map $h:[0,1]\to\mathbb{R}^2$ such that h([0,1]) is the square $\{0\leq x_1,x_1\leq 1\}$. (This is a famous Peano curve passing through every point of the square.)

Corollary 9.21 is a simplest special case of *Sard's theorem* which asserts that the set of critical values of a sufficiently smooth map has volume 0. More precisely,

Proposition 9.23. (A. SARD, 1942) Given a C^k -smooth map $f: A \to W$ (where A is a compact subset of V, $\dim V = n$, $\dim W = m$) let us denote by

$$\Sigma(f): \{x \in A; \text{ rank } d_x f < m\}.$$

Then if $k \ge \max(n - m + 1, 1)$ then $\operatorname{Vol}_m(f(\Sigma(f))) = 0$.

If m > n then $\Sigma(f) = A$, and hence the statement is equivalent to Corollary 9.21.

Proof. We prove the proposition only for the case m = n.

To clarify the main idea we first consider the case when n=1 and A=[0,1], i.e. f is a function $f:A\to\mathbb{R}$ with a continuous derivative f'. According to Cantor's theorem f' is uniformly continuous and hence for any ϵ there exists $\delta>0$ such that

$$|f'(x) - f'(y)| < \epsilon \text{ when } |x - y| < \delta. \tag{9.5.5}$$

Let us take a partition of the interval of the size $< \delta$. Let I_1, \ldots, I_N be the interval of the partition which contain critical points, i.e. points where the derivative is 0. Then for any point c in on

of these intervals $|f'(c)| < \epsilon$, and hence by the intermidiate value theorem for any two points $x, y \in I_j$, j = 1, ..., N we have $|f(x) - f(y)| = |f'(c)||x - y| < \epsilon \delta$, i.e. the image $f(I_j)$ is contained in an interval of length $\epsilon \delta$. But total length of the intrvals I_j is ≤ 1 , and thus $f(\Sigma)(f)$ is covered by the union of intervals of the total length $< \epsilon$. Hence $\operatorname{Vol}_1(f(\Sigma(f))) = 0$.

Consider now the case of a general n. Again the C^1 -smoothness of f implies that $d_x f$ is uniformly continuous, i.e. for every $\epsilon > 0$ there exists $\delta > 0$ such that

$$||d_x f - d_y f|| < \epsilon \text{ when } ||x - y|| < \delta. \tag{9.5.6}$$

The inequality $||d_x f - d_y f|| < \epsilon$ means that for any unit vector h in V

$$||d_x f(h) - d_y f(h)|| < \epsilon, \tag{9.5.7}$$

where we parallel transport the vector h to points x and y.

Consider a partition of a cube Q containing A by cubes of size $<\frac{\delta}{2\sqrt{n}}$, so that the ball surrounding each of the cubes and centered at any point of the cube has radius $<\delta$.

Let Q_1, \ldots, Q_N be the cubes of the partition intersecting $\Sigma(f)$. The total volume $2^n n^{-\frac{n}{2}} N \delta^n$ of these cubes is bounded by VolP, where P is a fixed cube containing A.

Choose a point $c_j \in Q_j \cap \Sigma(f)$ for each j = 1, ..., N Let $B_1, ..., B_N$ be balls of radius δ centered at c_j . As we already pointed out, $B_j \supset Q_j$ for each j = 1, ..., N.

The differential $d_{c_j}f$ is degenerate, and hence the image $d_{c_j}f(V_{c_j}) \subset W_{f(c_j)}$ is contained in a codimension 1 subspace $L_j \subset W_{f(c_j)}$. Let us choose a Cartesian coordinate system (y_1, \ldots, y_n) in $W_{f(c_j)}$ such that $L_j = \{y_n = 0\}$ we can view y_j as coordinates in V with the origin shifted to $f(c_j)$. Let (f_1, \ldots, f_n) be the coordinate functions of f with respect to these coordinates. Then $d_{c_j}f_n = 0$, i.e. the directional derivatives of the function f at c_j at every direction are equal to 0. But then according to inequality (9.5.7) the (absolute value of the) directional derivatives of f_n at any point of B_j are $< \epsilon$. Hence, using our above 1-dimensional argument along each radius of B_j , we conclude that $|f_n(x)| < \epsilon \delta$, i.e. the image $f(B_j)$ is contained in an $\epsilon \delta$ -neighborhood $U_{\epsilon \delta}$ of the hyperplane L_j viewed as an affine hyperplane in V. We also recall that the map f is Lipshitz, and hence the image $f(B_j)$ is contained in a ball $B_{K\delta}(f(c_j)) \subset W$ of radius $K\delta$ centered at $f(c_j)$ for some constant K > 0. Hence, $f(B_j) \subset U_{\epsilon \delta} \cap B_{K\delta}(f(c_j))$, so $f(B_j)$ is contained in a rectangular P_j with all sides equal to $2K\delta$ and one side of size $2\epsilon \delta$. In particular, $Vol(P_j) = K^{n-1}2^n\delta^n\epsilon$.

Therefore, $f(\Sigma(f))$ can be covered by N such rectangular of total volume

$$K^{n-1}2^nN\delta^n\epsilon = (K^{n-1}n^{\frac{n}{2}}\mathrm{Vol}P)\epsilon \underset{\epsilon \to 0}{\to} 0.$$

Hence, $Vol f(\Sigma(f)) = 0$.

Corollary 9.24. Let $A \subset \mathbb{R}^n$ be a measurable set and $f: A \to \mathbb{R}^q$, where $q \geq n$, be a C^1 -smooth map. Then the image $f(A) \subset \mathbb{R}^q$ is also measurable.

Proof. If q > n then $\operatorname{Vol}_q f(A) = 0$ according to Corollary 9.21, and hence f(A) is measurable. Suppose that q = n. Then for any interior point $a \in A$ such that $\operatorname{rank} d_a f = n$ the image f(a) is an interior point of f(A) according to the implicit function theorem. Hence, the boundary $\partial(f(A)) \subset f(\partial A) \cup f(\Sigma(f))$, where $\Sigma(f)$ is the set critical points of f (i.e. points $a \in A$ where $\operatorname{rank} d_a f < n$). But $\operatorname{Vol}\partial A = 0$ and hence, according to Lemma 9.20 $\operatorname{Vol}(f(\partial A)) = 0$. On the other hand, Sard's theorem 9.23 implies that $\operatorname{Vol}f(\Sigma(f)) = 0$, and therefore $\operatorname{Vol}(\partial f(A)) = 0$, which means that f(A) is measurable.

Properties which hold almost everywhere

We say that some property holds almost everywhere (we will abbreviate a.e.) if it holds in the complement of a set of volume 0. For instance, we say that a bounded function $f: P \to \mathbb{R}$ is almost everywhere continuous (or a.e. continuous) if it is continuous in the complement of a set $A \subset P$ of volume 0. For instance, a characteristic function of any measurable set is a.e. continuous. Indeed, it is constant away from the set ∂A which according to Proposition 9.16.4 has volume 0.

Proposition 9.25. Suppose that the bounded functions $f, g : P \to \mathbb{R}$ coincide a.e. Then if f is integrable, then so is g and we have $\int_{\mathcal{D}} f = \int_{\mathcal{D}} g$.

Proof. Denote $A = \{x \in P : f(x) \neq g(x)\}$. By our assumption, $\operatorname{Vol} A = 0$. Hence, for any ϵ there exists a $\delta > 0$ such that for every partition \mathcal{P} with $\delta(\mathcal{P}) \leq \delta$ the union B_{δ} of all rectangles of the partition which have non-empty intersection with A has volume $< \epsilon$. The functions f, g are bounded, i.e. there exists C > 0 such $-C \leq |f(x)|, |g(x)| \leq C$ for all $x \in P$. Due to integrability of f we can choose δ small enough so that $|U(f,\mathcal{P}) - L(f,\mathcal{P})| \leq \epsilon$ when $\delta(\mathcal{P}) \leq \delta$. Then we have

$$|U(g,\mathcal{P}) - U(f,\mathcal{P})| = \left| \sum_{J: P_J \subset B_\delta} \sup_{P_J} g - \sup_{P_J} f \right| \le 2C \mathrm{Vol} B_\delta \le 2C \epsilon.$$

Similarly, $|L(g, \mathcal{P}) - L(f, \mathcal{P})| \leq 2C\epsilon$, and hence

$$|U(g,\mathcal{P}) - L(g,\mathcal{P})| \le |U(g,\mathcal{P}) - U(f,\mathcal{P})| + |U(f,\mathcal{P}) - L(f,\mathcal{P})| + |L(f,\mathcal{P}) - L(g,\mathcal{P})|$$
$$\le \epsilon + 4C\epsilon \underset{\delta \to 0}{\longrightarrow} 0,$$

and hence g is integrable and

$$\int\limits_{P} g = \lim_{\delta(\mathcal{P}) \to 0} U(g, \mathcal{P}) = \lim_{\delta(\mathcal{P}) \to 0} U(f, \mathcal{P}) = \int\limits_{P} f.$$

Proposition 9.26. 1. Suppose that a function $f: P \to \mathbb{R}$ is a.e. continuous. Then f is integrable.

2. Let $A \subset V$ be compact and measurable, $f: U \to W$ a C^1 -smooth map defined on a neighborhood $U \supset A$. Suppose that $\dim W = \dim V$. Then f(A) is measurable.

Proof. 1. Let us begin with a

Warning. One could think that in view of Proposition 9.25 it is sufficient to consider only the case when the function f is continuous. However, this is not the case, because for a given a.e. continuous function one cannot, in general, find a continuous function g which coincides with f a.e.

Let us proceed with the proof. Given a partition \mathcal{P} we denote by J_A the set of multi-indices \mathbf{j} such that Int $P_{\mathbf{j}} \cap A \neq \emptyset$, and by \overline{J}_A the complementary set of multi-indices, i.e. for each $\mathbf{j} \in \overline{J}_A$ we have $P_{\mathbf{j}} \cap A = \emptyset$. Let us denote $C := \bigcup_{\mathbf{j} \in J_A} P_{\mathbf{j}}$. According to Proposition 9.16.3 for any $\epsilon > 0$ there exists a partition \mathcal{P} such that $\operatorname{Vol}(C) = \sum_{\mathbf{j} \in J_A} \operatorname{Vol}(P_{\mathbf{j}}) < \epsilon$. By assumption the function f is continuous over a compact set $B = \bigcup_{\mathbf{j} \in \overline{J}_A} P_{\mathbf{j}}$, and hence it is uniformly continuous over it. Thus there exists $\delta > 0$ such that $|f(x) - f(x')| < \epsilon$ provided that $x, x' \in B$ and $||x - x'|| < \delta$. Thus we can further subdivide our partition, so that for the new finer partition \mathcal{P}' we have $\delta(\mathcal{P}') < \delta$. By assumption the function f is bounded, i.e. there exists a constant K > 0 such that $M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f) < K$ for all

indices **j**. Then we have

$$\begin{split} U(f;\mathcal{P}') - L(f,\mathcal{P}') &= \sum_{\mathbf{j}} (M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f)) \mathrm{Vol}(P_{\mathbf{j}}) = \\ &\sum_{\mathbf{j}; P_{\mathbf{j}} \subset B} (M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f)) \mathrm{Vol}(P_{\mathbf{j}}) + \sum_{\mathbf{j}; P_{\mathbf{j}} \subset C} (M_{\mathbf{j}}(f) - m_{\mathbf{j}}(f)) \mathrm{Vol}(P_{\mathbf{j}}) < \\ &\epsilon \mathrm{Vol}B + K \mathrm{Vol}C < \epsilon (\mathrm{Vol}P + K). \end{split}$$

Hence $\inf_{\mathcal{P}} U(f; \mathcal{P}) = \sup_{\mathcal{P}} L(f; \mathcal{P})$, i.e. the function f is integrable.

2. If x is an interior point of A and $det Df(x) \neq 0$ then the inverse function theorem implies that $f(x) \in \text{Int } f(A)$. Denote $C = \{x \in A; \det Df(x) = 0\}$. Hence, $\partial f(A) \subset f(\partial A) \cup f(C)$. But $\text{Vol}(\partial A) = 0$ because A is measurable and Vol f(C) = 0 by Sard's theorem 9.23. Therefore, $\text{Vol} \partial f(A) = 0$ and thus f(A) is measurable.

Orthogonal invariance of the volume and volume of a parallelepiped

The following lemma provides a way of computing the volume via packing by balls rather then cubes. An admissible set of balls in A is any finite set of disjoint balls $B_1, \ldots, B_K \subset A$

Lemma 9.27. Let A be a measurable set. Then VolA is the supremum of the total volume of admissible sets of balls in A. Here the supremum is taken over all admissible sets of balls in A.

Proof. Let us denote this supremum by β . The monotonicity of volume implies that $\beta \leq \text{Vol}A$. Suppose that $\beta < VolA$. Let us denote by μ_n the volume of an n-dimensional ball of radius 1 (we will compute this number later on). This ball is contained in a cube of volume 2^n . It follows then that the ratio of the volume of any ball to the volume of the cube to which it is inscribed is equal to $\frac{\mu_n}{2^n}$. Choose an $\epsilon < \frac{\mu_n}{2^n}(\text{Vol}A - \beta)$. Then there exists a finite set of disjoint balls $B_1, \ldots, B_K \subset A$ such that $\text{Vol}(\bigcup_{1}^{K} B_j) > \beta - \epsilon$. The volume of the complement $C = A \setminus \bigcup_{1}^{K} B_j$ satisfies

$$\operatorname{Vol} C = \operatorname{Vol} A - \operatorname{Vol} \left(\bigcup_{1}^{K} B_{j} \right) > \operatorname{Vol} A - \beta.$$

Hence there exists a partition \mathcal{P} of P by cubes such that the total volume of cubes Q_1, \ldots, Q_L contained in C is $> \text{Vol}A - \beta$. Let us inscribe in each of the cubes Q_j a ball \widetilde{B}_j . Then $B_1, \ldots, B_K, \widetilde{B}_1, \ldots, \widetilde{B}_L$

is an admissible set of balls in A. Indeed, all these balls are disjoint and contained in A. The total volume of this admissible set is equal to

$$\sum_{1}^{K} \operatorname{Vol}B_{j} + \sum_{1}^{L} \operatorname{Vol}\widetilde{B}_{i} \ge \beta - \epsilon + \frac{\mu_{n}}{2^{n}} (\operatorname{Vol}A - \beta) > \beta,$$

in view of our choice of ϵ , but this contradicts to our assumption $\beta < \text{Vol}A$. Hence, we have $\beta = \text{Vol}A$.

Lemma 9.28. Let $A \subset V$ be any measurable set in a Euclidean space V. Then for any linear orthogonal transformation $F: V \to V$ the set F(A) is also measurable and we have Vol(F(A)) = Vol(A).

Proof. First note that if VolA = 0 then the claim follows from Lemma 9.20. Indeed, an orthogonal transformation is, of course a smooth map.

Let now A be an arbitrary measurable set. Note that $\partial F(A) = F(\partial A)$. Measurability of A implies $\operatorname{Vol}(\partial A) = 0$. Hence, as we just have explained, $\operatorname{Vol}(\partial F(A)) = \operatorname{Vol}(F(\partial A)) = 0$, and hence F(A) is measurable. According to Lemma 9.27 the volume of a measurable set can be computed as a supremum of the total volume of disjoint inscribed balls. But the orthogonal transformation F moves disjoint balls to disjoint balls of the same size, and hence $\operatorname{Vol} A = \operatorname{Vol} F(A)$.

Next proposition shows that the volume of a parallelepiped can be computed by formula (3.3.1) from Section 3.3.

Proposition 9.29. Let $v_1, \ldots, v_n \in V$ be linearly independent vectors. Then

$$\operatorname{Vol} P(v_1, \dots, v_n) = |x_1 \wedge \dots \wedge x_n(v_1, \dots, v_n)|. \tag{9.5.8}$$

Proof. The formula (9.5.8) holds for rectangles, i.e. when $v_j = c_j e_j$ for some non-zero numbers c_j , j = 1, ...n. Using Lemma 9.28 we conclude that it also holds for any orthogonal basis. Indeed, any such basis can be moved by an orthogonal transformation to a basis of the above form $c_j e_j$, j = 1, ...n. Lemma 9.28 ensures that the volume does not change under the orthogonal transformation, while Proposition 2.17 implies the same about $|x_1 \wedge \cdots \wedge x_n(v_1, ..., v_n)|$.

The Gram-Schmidt orthogonalization process shows that one can pass from any basis to an orthogonal basis by a sequence of following elementary operations:

- - reordering of basic vectors, and
- shears, i.e. an addition to the last vector a linear combination of the other ones:

$$v_1, \ldots, v_{n-1}, v_n \mapsto v_1, \ldots, v_{n-1}, v_n + \sum_{j=1}^{n-1} \lambda_j v_j.$$

Note that the reordering of vectors v_1, \ldots, v_n changes neither $VolP(v_1, \ldots, v_n)$, nor the absolute value

 $|x_1 \wedge \cdots \wedge x_n(v_1, \ldots, v_n)|$. On the other hand, a shear does not change

$$x_1 \wedge \cdots \wedge x_n(v_1 \ldots, v_n).$$

It remains to be shown that a shear does not change the volume of a parallelepiped. We will consider here only the case n=2 and will leave to the reader the extension of the argument to the general case.

Let v_1, v_2 be two orthogonal vectors in \mathbb{R}^2 . We can assume that $v_1 = (a, 0), v_2 = (0, b)$ for a, b > 0, because we already proved the invariance of volume under orthogonal transformations. Let $v_2' = v_2 + \lambda v_1 = (a', b)$, where $a' = a + \lambda b$. Let us partition the rectangle $P = P(v_1, v_2)$ into N^2 smaller rectangles $P_{i,j}, i, j = 0, \ldots, N-1$, of equal size. We number the rectangles in such a way that the first index corresponds to the first coordinate, so that the rectangles $P_{00}, \ldots, P_{N-1,0}$ form the lower layer, $P_{01}, \ldots, P_{N-1,1}$ the second layer, etc. Let us now shift the rectangles in k-th layer horizontally by the vector $(\frac{k\lambda b}{N}, 0)$. Then the total volume of the rectangles, denoted \tilde{P}_{ij} remains the same, while when $N \to \infty$ the volume of part of the parallelogram $P(v_1, v_2')$ that is not covered by rectangles $\tilde{P}_{i,j}, i, j = 0, \ldots, N-1$ converges to 0.

9.6 Fubini's Theorem

Let us consider \mathbb{R}^n as a direct product of \mathbb{R}^k and \mathbb{R}^{n-k} for some $k=1,\ldots,n-1$. We will denote coordinates in \mathbb{R}^k by $x=(x_1,\ldots,x_k)$ and coordinates in \mathbb{R}^{n-k} by $y=(y_1,\ldots,y_{n-k})$, so the coordinates in \mathbb{R}^n are denoted by $(x_1,\ldots,x_k,y_1,\ldots,y_{n-k})$. Given rectangles $P_1\subset\mathbb{R}^k$ and $P_2\subset\mathbb{R}^{n-k}$ their product $P=P_1\times P_2$ is a rectangle in \mathbb{R}^n .



Figure 9.2: Guido Fubini (1879-1943)

The following theorem provides us with a basic tool for computing multiple integrals.

Theorem 9.30 (Guido Fubini). Suppose that a function $f: P \to \mathbb{R}$ is integrable over P. Given a point $x \in P_1$ let us define a function $f_x: P_2 \to \mathbb{R}$ by the formula $f_x(y) = f(x, y), y \in P_2$. Then

$$\int\limits_P f dV_n = \int\limits_{P_1} \left(\int\limits_{\overline{P_2}} f_x dV_{n-k} \right) dV_k = \int\limits_{P_1} \left(\int\limits_{\overline{P_2}} f_x dV_{n-k} \right) dV_k.$$

In particular, if the function f_x is integrable for all (or almost all) $x \in P_1$ then one has

$$\int_{P} f dV_n = \int_{P_1} \left(\int_{P_2} f_x dV_{n-k} \right) dV_k.$$

Here by writing dV_k , dV_{n-k} and dV_n we emphasize the integration with respect to the k-, (n-k)and n-dimensional volumes, respectively.

Proof. Choose any partition \mathcal{P}_1 of P_1 and \mathcal{P}_2 of P_2 . We will denote elements of the partition \mathcal{P}_1 by $P_1^{\mathbf{j}}$ and elements of the partition \mathcal{P}_2 by $P_2^{\mathbf{i}}$. Then products of $P^{\mathbf{j},\mathbf{i}} = P_1^{\mathbf{j}} \times P_2^{\mathbf{i}}$ form a partition \mathcal{P} of $P = P_1 \times P_2$. Let us denote

$$\overline{I}(x) := \overline{\int_{P_2}} f_x, \ \underline{I}(x) := \int_{\overline{P_2}} f_x, \ x \in P_1.$$

Let us show that

$$L(f, \mathcal{P}) \le L(\underline{I}, \mathcal{P}_1) \le U(\overline{I}, \mathcal{P}_1) \le U(f, \mathcal{P}).$$
 (9.6.1)

Indeed, we have

$$L(f, \mathcal{P}) = \sum_{\mathbf{i}} \sum_{\mathbf{i}} m_{\mathbf{j}, \mathbf{i}}(f) \operatorname{Vol}_n P^{\mathbf{j}, \mathbf{i}}.$$

Here the first sum is taken over all multi-indices \mathbf{j} of the partition \mathcal{P}_1 , and the second sum is taken over all multi-indices \mathbf{i} of the partition \mathcal{P}_2 . On the other hand,

$$L(\underline{I}, \mathcal{P}_1) = \sum_{\mathbf{j}} \inf_{x \in P_1^{\mathbf{j}}} \left(\int_{\underline{P}_2} f_x dV_{n-k} \right) \operatorname{Vol}_k P_1^{\mathbf{j}}.$$

Note that for every $x \in P_1^{\mathbf{j}}$ we have

$$\int_{\overline{P_2}} f_x dV_{n-k} \ge L(f_x; \mathcal{P}_2) = \sum_{\mathbf{i}} m_{\mathbf{i}}(f_x) \operatorname{Vol}_{n-k}(P_2^{\mathbf{i}}) \ge \sum_{\mathbf{i}} m_{\mathbf{i}, \mathbf{j}}(f) \operatorname{Vol}_{n-k}(P_2^{\mathbf{i}}),$$

and hence

$$\inf_{x \in P_1^{\mathbf{j}}} \int_{P_2} f_x dV_{n-k} \ge \sum_{\mathbf{i}} m_{\mathbf{i},\mathbf{j}}(f) \operatorname{Vol}_{n-k}(P_2^{\mathbf{i}}).$$

Therefore,

$$L(\underline{I},\mathcal{P}_1) \geq \sum_{\mathbf{i}} \sum_{\mathbf{i}} m_{\mathbf{i},\mathbf{j}}(f) \operatorname{Vol}_{n-k}(P_2^{\mathbf{i}}) \operatorname{Vol}_k(P_1^{\mathbf{j}}) = \sum_{\mathbf{i}} \sum_{\mathbf{i}} m_{\mathbf{j},\mathbf{i}}(f) \operatorname{Vol}_n(P^{\mathbf{j},\mathbf{i}}) = L(f,\mathcal{P}).$$

Similarly, one can check that $U(\overline{I}, \mathcal{P}_1) \leq U(f, \mathcal{P})$. Together with an obvious inequality $L(\underline{I}, \mathcal{P}_1) \leq U(\overline{I}, \mathcal{P}_1)$ this completes the proof of (9.6.1). Thus we have

$$\max(U(\overline{I}, \mathcal{P}_1) - L(\overline{I}, \mathcal{P}_1), U(\underline{I}, \mathcal{P}_1) - L(\underline{I}, \mathcal{P}_1)) \leq U(\overline{I}, \mathcal{P}_1) - L(\underline{I}, \mathcal{P}_1) \leq U(f, \mathcal{P}) - L(f, \mathcal{P}).$$

By assumption for appropriate choices of partitions, the right-hand side can be made $< \epsilon$ for any a priori given $\epsilon > 0$. This implies the integrability of the function $\underline{I}(x)$ and $\overline{I}(x)$ over P_1 . But then we can write

$$\int_{P_1} \underline{I}(x)dV_{n-k} = \lim_{\delta(\mathcal{P}_1) \to 0} L(\underline{I}; \mathcal{P}_1)$$

and

$$\int_{P_1} \overline{I}(x)dV_{n-k} = \lim_{\delta(\mathcal{P}_1) \to 0} U(\overline{I}; \mathcal{P}_1).$$

We also have

$$\lim_{\delta(\mathcal{P})\to 0} L(f;\mathcal{P}) = \lim_{\delta(\mathcal{P})\to 0} U(f;\mathcal{P}) = \int_{\mathcal{P}} f dV_n.$$

Hence, the inequality (9.6.1) implies that

$$\int\limits_P f dV_n = \int\limits_{P_1} \left(\int\limits_{\overline{P_2}} f_x dV_{n-k} \right) dV_k = \int\limits_{P_1} \left(\int\limits_{\overline{P_2}} f_x dV_{n-k} \right) dV_k.$$

Corollary 9.31. Suppose $f: P \to \mathbb{R}$ is a continuous function. Then

$$\int\limits_P f = \int\limits_{P_1} \int\limits_{P_2} f_x = \int\limits_{P_2} \int\limits_{P_1} f_y.$$

Thus if we switch back to the notation x_1, \ldots, x_n for coordinates in \mathbb{R}^n , and if $P = \{a_1 \leq x_1 \leq b_1, \ldots, a_n \leq x_n \leq b_n\}$ then we can write

$$\int_{P} f = \int_{a_n}^{b_n} \left(\dots \left(\int_{a_1}^{b_1} f(x_1, \dots, x_n) dx_1 \right) \dots \right) dx_n.$$
 (9.6.2)

The integral in the right-hand side of (9.6.2) is called an *iterated integral*. Note that the order of integration is irrelevant there. In particular, for continuous functions one can change the order of integration in the iterated integrals.

9.7 Integration of *n*-forms over domains in *n*-dimensional space

Differential forms are much better suited to be integrated than functions. For integrating a function, one needs a measure. To integrate a differential form, one needs nothing except an orientation of the domain of integration.

Let us start with the integration of a n-form over a domain in a n-dimensional space. Let ω be a n-form on a domain $U \subset V$, dim V = n.

Let us fix now an orientation of the space V. Pick any coordinate system $(x_1 \dots x_n)$ that agrees with the chosen orientation.

We proceed similar to the way we defined an integral of a function. Let us fix a rectangle $P = P(a_1, b_1; a_2, b_2; \dots; a_n, b_n) = \{a_i \leq x_i \leq b_i; i = 1, \dots, n\}$. Choose its partition \mathcal{P} by N^n smaller rectangles $P_{\mathbf{j}} = \{t_{j_n}^1 \leq x_1 \leq t_{j_n+1}^1, \dots, t_{j_1}^n \leq x_1 \leq t_{j_1+1}^n\}$, where $\mathbf{j} = (j_1, \dots, j_n)$ and each index j_k takes values between 0 and N-1. Let us fix a point $c_{\mathbf{j}} \in P_{\mathbf{j}}$ and denote by C the set of all such $c_{\mathbf{j}}$. We also denote by $t_{\mathbf{j}}$ the point with coordinates $t_{j_1}^1, \dots, t_{j_n}^n$ and by $T_{\mathbf{j},m} \in V_{c_{\mathbf{j}}}, m = 1, \dots, n$ the vector $t_{\mathbf{j}+1_m} - t_{\mathbf{j}}$, parallel-transported to the point $c_{\mathbf{j}}$. Here we use the notation $\mathbf{j} + 1_m$ for the multi-index $j_1, \dots, j_{m-1}, j_m + 1, j_{m+1}, \dots, j_n$. Thus the vector $T_{\mathbf{j},m}$ is parallel to the m-th basic vector and has the length $|t_{j_m+1} - t_{j_m}|$.

Given a differential n-form α on P we form an integral sum

$$I(\alpha; \mathcal{P}, C) = \sum_{\mathbf{j}} \alpha(T_1^{\mathbf{j}}, T_2^{\mathbf{j}}, \dots, T_{\mathbf{j}, n}), \tag{9.7.1}$$

where the sum is taken over all elements of the partition. We call an n-form α integrable if there exists a limit $\lim_{\delta(\mathcal{P})\to 0} I(\alpha;\mathcal{P},C)$ which we denote by $\int_{\mathcal{P}} \alpha$ and call the integral of α over P. Note that if $\alpha = f(x)dx_1 \wedge \cdots \wedge dx_n$ then the integral sum $I(\alpha,\mathcal{P},C)$ from (9.7.1) coincides with the integral sum $I(f;\mathcal{P},C)$ from (9.5.2) for the function f. Thus the integrability of α is the same as integrability of f and we have

$$\int_{P} f(x)dx_1 \wedge \dots \wedge dx_n = \int_{P} fdV. \tag{9.7.2}$$

Note, however, that the equality (9.7.2) holds only if the coordinate system (x_1, \ldots, x_n) defines the given orientation of the space V. The integral $\int_P f(x)dx_1 \wedge \cdots \wedge dx_n$ changes its sign with a change of the orientation while the integral $\int_P f dV$ is not sensitive to the orientation of the space V.

It is not clear from the above definition whether the integral of a differential form depends on our choice of the coordinate system. It turns out that it does not, as the following theorem, which is the main result of this section, shows. Moreover, we will see that one even can use arbitrarty curvilinear coordinates.

In what follows we use the convention introduced at the end of Section 6.1. Namely by a diffeomorphism between two closed subsets of vector spaces we mean a diffeomorphism between their neighborhoods.

Theorem 9.32. Let $A, B \subset \mathbb{R}^n$ be two measurable compact subsets. Let $f : A \to B$ be an orientation preserving diffeomorphism. Let η be a differential n-form defined on B. Then if η is integrable over B then $f^*\alpha$ is integrable over A and we have

$$\int_{A} f^* \eta = \int_{B} \eta. \tag{9.7.3}$$

For an orientation reversing diffeomorphism f we have $\int_A f^* \eta = -\int_B \eta$.

Let $\alpha = g(x)dx_1 \wedge \cdots \wedge dx_n$. Then $f^*\alpha = g \circ f \det Df dx_1 \wedge \cdots \wedge dx_n$, and hence the formula (9.7.3) can be rewritten as

$$\int_{P} g(x_1, \dots, x_n) dx_1 \wedge \dots \wedge dx_n = \int_{P} g \circ f \det Df dx_1 \wedge \dots \wedge dx_n.$$

Here

$$\det Df = \begin{vmatrix} \frac{\partial f_1}{\partial x_1} & \cdots & \frac{\partial f_1}{\partial x_n} \\ \vdots & \ddots & \vdots \\ \frac{\partial f_n}{\partial x_1} & \cdots & \frac{\partial f_n}{\partial x_n} \end{vmatrix}$$

is the determinant of the Jacobian matrix of $f = (f_1, \ldots, f_n)$.

Hence, in view of formula (9.7.2) we get the following *change of variables* formula for multiple integrals of functions.

Corollary 9.33. [Change of variables in a multiple integral] Let $g: B \to \mathbb{R}$ be an integrable function and $f: A \to B$ a diffeomorphism. Then the function $g \circ f$ is also integrable and

$$\int_{B} g dV = \int_{A} g \circ f |\det Df| dV. \qquad (9.7.4)$$

We begin the proof with the following special case of Theorem 9.32.

Proposition 9.34. The statement of 9.32 holds when $\eta = dx_1 \wedge \cdots \wedge dx_n$ and the set A is the unit cube $I = I^n$. In other words,

$$\operatorname{Vol} f(I) = \left| \int_A f^* \eta \right|.$$

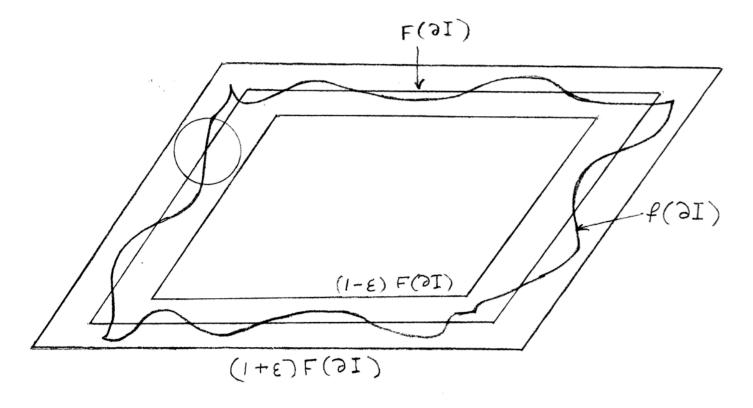


Figure 9.3: Image of a cube under a diffeomorphism and its linearization

We will use below the following notation.

For any set $A \subset V$ and any positive number $\lambda > 0$ we denote by λA the set $\{\lambda x, x \in A\}$. For any linear operator $F: V \to W$ between two Euclidean spaces V and W we define its norm ||F|| by the formula

$$||F|| = \max_{||v||=1} ||F(v)|| = \max_{v \in V, v \neq 0} \frac{||F(v)||}{||v||}.$$

Equivalently, we can define ||F|| as follows. The linear map F maps the unit sphere in the space V onto an ellipsoid in the space W. Then ||F|| is the biggest semi-axis of this ellipsoid.

Let us begin by observing the following geometric fact:

Lemma 9.35. Let $I = \{|x_i| \leq \frac{1}{2}, i = 1, \dots, n\} \subset \mathbb{R}^n$ be the unit cube centered at 0 and $F : \mathbb{R}^n \to \mathbb{R}^n$ a non-degenerate linear map. Take any $\epsilon \in (0,1)$ and set $\sigma = \frac{\epsilon}{\|F^{-1}\|}$. Then for any boundary point $z \in \partial I$ we have

$$B_{\sigma}(F(z)) \subset (1+\epsilon)F(I) \setminus (1-\epsilon)F(I),$$
 (9.7.5)

see Fig. 9.7

Proof. Inclusion (9.7.5) can be rewritten as

$$F^{-1}(B_{\sigma}(F(z)) \subset (1+\epsilon)I \setminus (1-\epsilon)I.$$

But the set $F^{-1}(B_{\sigma}(F(z)))$ is an ellipsoid centered at z whose greatest semi-axis is equal to $\sigma||F^{-1}||$. Hence, if $\sigma||F^{-1}|| \leq \epsilon$ then $F^{-1}(B_{\sigma}(F(z))) \subset (1+\epsilon)I \setminus (1-\epsilon)I$.

Recall that we denote by δI the cube I scaled with the coefficient δ , i.e. $\delta I = \{|x_i| \leq \frac{\delta}{2}, i = 1, \ldots, n\} \subset \mathbb{R}^n$. We will also need

Lemma 9.36. Let $U \subset \mathbb{R}^n$ is an open set, $f: U \to \mathbb{R}^n$ such that f(0) = 0. Suppose that f is differentiable at 0 and its differential $F = d_0 f: \mathbb{R}^n \to \mathbb{R}^n$ at 0 is non-degenerate. Then for any $\epsilon \in (0,1)$ there exists $\delta > 0$ such that

$$(1 - \epsilon)F(\delta I) \subset f(\delta I) \subset (1 + \epsilon)F(\delta I), \tag{9.7.6}$$

see Fig. 9.7.

Proof. First, we note that inclusion (9.7.5) implies, using linearity of F, that for any $\delta > 0$ we have

$$B_{\delta\sigma}(F(z)) \subset (1+\epsilon)F(\delta I) \setminus (1-\epsilon)F(\delta I),$$
 (9.7.7)

where $z \in \partial(\delta I)$ and, as in Lemma 9.35, we assume that $\epsilon = \sigma ||F^{-1}||$.

According to the definition of differentiability we have

$$f(h) = F(h) + o(||h||).$$

Denote $\widetilde{\sigma} := \frac{\sigma}{\sqrt{n}} = \frac{\epsilon}{\sqrt{n}||F^{-1}||}$. There exists $\rho > 0$ such that if $||h|| \le \rho$ then

$$||f(h) - F(h)|| \le \widetilde{\sigma}||h|| \le \widetilde{\sigma}\rho.$$

Denote $\delta := \frac{\rho}{\sqrt{n}}$. Then $\delta I \subset B_{\rho}(0)$, and hence $||f(z) - F(z)|| \leq \tilde{\sigma}\rho$ for any $z \in \delta I$. In particular, for any point $z \in \partial(\delta I)$ we have

$$f(z) \in B_{\widetilde{\sigma}\rho}(F(z)) = B_{\sqrt{n}\widetilde{\sigma}\delta}(F(z)) = B_{\sigma\delta}(F(z)),$$

and therefore in view of (9.7.7)

$$f(\partial(\delta I)) \subset (1+\epsilon)F(\delta I) \setminus (1-\epsilon)F(\delta I).$$

But this is equivalent to inclusion (9.7.6).

Lemma 9.37. Let $F: \mathbb{R}^n \to \mathbb{R}^n$ be a non-degenerate orientation preserving linear map, $P = P(v_1, \ldots, v_n)$ a parallelepiped, and $\eta = dx_1 \wedge \cdots \wedge dx_n$. Then $\int\limits_{F(P)} \eta = \int\limits_{P} F^* \eta$. Here we assume that the orientation of P and F(P) are given by the orientation of \mathbb{R}^n .

Proof. We have $\int\limits_{F(P)} \eta = \int\limits_{F(P)} dx_1 \wedge \dots dx_n = \operatorname{Vol} F(P) = (\det F) \operatorname{Vol} P$. On the other hand, $F^* \eta = \det F \eta$, and hence $\int\limits_{P} F^* \eta = \det F \int\limits_{P} \eta = (\det F) \operatorname{Vol} P$.

Proof of Proposition 9.34. We have $f^*\eta = (\det Df)dx_1 \wedge \cdots \wedge dx_n$, and hence the form $f^*\eta$ is integrable because f is C^1 -smooth, and hence $\det Df$ is continuous.

Choose a partition \mathcal{P} of the cube I by N^n small cubes I_K , $K = 1, \ldots, N^n$, of the same size $\frac{1}{N}$. Let $c_K \in I_K$ be the center of the cube I_K . Then

$$\int\limits_I f^*\eta = \sum_{K=1}^{N^n} \int\limits_{I_K} f^*\eta.$$

Note that in view of the uniform continuity of the function $\det Df$, for any $\epsilon > 0$ the number N can be chosen so large that $|\det Df(x) - \det Df(x')| < \epsilon$ for any two points $x, x' \in I_K$ and any $K = 1, \ldots, N^n$. Let η_K be the form $\det Df(c_k)dx_1 \wedge \cdots \wedge dx_n$ on I_K . Then

$$\left| \int_{I_K} f^* \eta - \int_{I_K} \eta_K \right| \le \int_{I_K} |\det Df(x) - \det Df(c_K)| \, dx_1 \wedge \dots \wedge dx_n \le \epsilon \operatorname{Vol}(I_K) = \frac{\epsilon}{N^n}.$$

Thus

$$\left| \int_{I} f^* \eta - \sum_{K=1}^{N^n} \int_{I_K} \eta_K \right| \le \epsilon. \tag{9.7.8}$$

Next, let us analyze the integral $\int_{I_K} \eta_K$. Denote $F_K := d_{c_K}(f)$. We can assume without loss of generality that $c_K = 0$, and $f(c_K) = 0$, and hence F_K can be viewed just as a linear map $\mathbb{R}^n \to \mathbb{R}^n$.

Using Lemma 9.36 we have for a sufficiently large N

$$(1-\epsilon)F_K(I_K) \subset f(I_K) \to (1+\epsilon)F_K(I_K).$$

Again in view of compactness of I the number ϵ can be chosen the same for all cubes I_K . Hence

$$(1 - \epsilon)^n \operatorname{Vol}(F_K(I_K)) \le \operatorname{Vol}f(I_K) \le (1 + \epsilon)^n \operatorname{Vol}(F_K(I_K)). \tag{9.7.9}$$

Note that $\operatorname{Vol} f(I_K) = \int_{f(I_K)} \eta$, and hence summing up inequality (9.7.9) over K we get

$$(1 - \epsilon)^n \sum_{K=1}^{N^n} \text{Vol}(F_K(I_K)) \le \sum_{K=1}^{N^n} \text{Vol}f(I_K) = \sum_{K=1}^{N^n} \int_{f(I_K)} \eta = \int_{f(I)} \eta \le (1 + \epsilon)^n \sum_{K=1}^{N^n} \text{Vol}(F_K(I_K)).$$
(9.7.10)

Note that $\eta_K = F_K^* \eta$ and by Lemma 9.37 we have $\int_{I_K} \eta_K = \int_{I_K} F_K^* \eta = \int_{F_K(I_K)} \eta = \text{Vol}(F_K(I_K))$. Hence, it follows from (9.7.10) that

$$(1 - \epsilon)^n \sum_{K=1}^{N^n} \int_{I_K} \eta_K \le \int_{f(I)} \eta \le (1 + \epsilon)^n \sum_{K=1}^{N^n} \int_{I_K} \eta_K.$$
 (9.7.11)

Recall that from (9.7.8) we have

$$\int\limits_{I} f^* \eta - \epsilon \le \sum_{K=1}^{N^n} \int\limits_{I_K} \le \int\limits_{I} f^* \eta + \epsilon.$$

Combining with (9.7.11) we get

$$(1 - \epsilon)^n \left(\int_I f^* \eta - \epsilon \right) \le \int_{f(I)} \eta \le (1 + \epsilon)^n \left(\int_I f^* \eta + \epsilon \right). \tag{9.7.12}$$

Passing to the limit when $\epsilon \to 0$ we get

$$\int_{I} f^* \eta \le \int_{f(I)} \eta \le \int_{I} f^* \eta, \tag{9.7.13}$$

i.e.
$$\int\limits_I f^* \eta = \int\limits_{f(I)} \eta$$
.

Corollary 9.38. The statement of Theorem 9.32 holds for $\eta = dx_1 \wedge \cdots \wedge dx_n$ and an arbitrary measurable A.

Proof. Suppose for determinacy that f preserves the orientation. By assumption the diffeomorphism A extends to an open neighborhood $U \supset A$. Consider a cube containing $P \supset A$. Choose

a $\delta > 0$ and consider a partition \mathcal{P} of P by small cube of size $< \delta$. Denote by A_{δ}^+ the union of elements of the partition which intersect A, and by A_{δ}^- the union of elements which are completely inside A. If δ is small enough then $A_{\delta}^+ \subset U$. We have

$$\int_{A^-\delta} f^* \eta \le \int_A f^* \eta \le \int_A A_\delta^+ f^* \eta$$

and

$$\int A_{\delta}^{+} f^{*} \eta - \int_{A-\delta} f^{*} \eta \underset{\delta \to 0}{\longrightarrow} 0.$$

On the other hand, each of the sets A_{δ}^{+} is a union of cubes. Hence, Proposition 9.34 implies that

$$\int A_{\delta}^{\pm} f^* \eta = \int_{f(A_{\delta}^{\pm})} \delta = \operatorname{Vol}(A_{\delta}^{\pm}).$$

But $f(A_{\delta}^-) \subset f(A) \subset f(A_{\delta}^+)$, and hence $\operatorname{Vol}(f(A_{\delta}^-)) \leq \operatorname{Vol}(f(A) \leq \operatorname{Vol}(f(A_{\delta}^+)))$ which implies when $\delta \to 0$ that $\operatorname{Vol} f(A) = \int\limits_{f(A)} \eta = \int\limits_{A} f^* \eta$.

Proof of Theorem 9.32. Let us recall that the diffeomorphism f is defined as a diffeomorphism between open neighborhoods $U \supset A$ and $U' \supset B$. We also assume that the form η is extended to U' as equal to 0 outside B. The form η can be written as $hdx_1 \land \cdots \land dx_n$. Let us take a partition \mathcal{P} of a rectangular containg U' by cubes $I_{\mathbf{j}}$ of the same size δ . Consider forms $\eta_{\mathbf{j}}^+ := M_{\mathbf{j}}(h)dx_1 \land \cdots \land dx_n$ and $\eta_{\mathbf{j}}^- := m_{\mathbf{j}}(h)dx_1 \land \cdots \land dx_n$ on $I_{\mathbf{j}}$, where $m_{\mathbf{j}}(h) = \inf_{I_{\mathbf{j}}} h$, $M_{\mathbf{j}}(h) = \sup_{I_{\mathbf{j}}} (h)$. Let η^{\pm} be the form on U' equal to $\eta_{\mathbf{j}}^{\pm}$ on each cube $I_{\mathbf{j}}$. The assumption of integrability of η over B guarantees that for any $\epsilon > 0$ if δ is chosen small enough we have $\int_B \eta^+ - \int_B \eta \le \epsilon$. The forms $f^*\eta^+$ are a.e. continuous, and hence integrable over A and we have $\int_A f^*\eta^- \le \int_A f^*\eta^+$. Hence, if we prove that $\int_A f^*\eta^\pm = \int_B \eta^\pm$ then this will imply that η is integrable and $\int_A \eta = \int_B \eta$.

On the other hand, $\int_{B} \eta^{\pm} = \sum_{\mathbf{j}} \int_{I_{\mathbf{j}}} \eta_{\mathbf{j}}^{\pm}$ and $\int_{A} f^{*} \eta^{\pm} = \sum_{\mathbf{j}} \int_{B_{\mathbf{j}}} f^{*} \eta_{\mathbf{j}}^{\pm}$, where $B_{\mathbf{j}} = f^{-1}(I_{\mathbf{j}})$. But according to Corollary ?? we have $\int_{B_{\mathbf{j}}} f^{*} \eta_{\mathbf{j}}^{\pm} = \int_{I_{\mathbf{j}}} \eta_{\mathbf{j}}^{\pm}$, and hence $\int_{A} f^{*} \eta^{\pm} = \int_{B} \eta^{\pm}$.

9.8 Manifolds and submanifolds

9.8.1 Manifolds

Manifolds of dimension n are spaces which are locally look like open subsets of \mathbb{R}^n but globally could be much more complicated. We give a precise definition below.

Let $U, U' \subset \mathbb{R}^n$ be open sets. A map $f: U \to U'$ is called a *homeomorpism* if it is continuous one-to-one map which has a continuous inverse $f^{-1}: U' \to U$.

A map $f: U \to U'$ is called a C^k -diffeomorpism, $k = 1, ..., \infty$, if it is C^k -smooth, one-to-one map which has a C^k -smooth inverse $f^{-1}: U' \to U$. Usually we will omit the reference to the class of smoothness, and just call f a diffeomorphism, unless it will be important to emphasize the class of smoothness.

A set M is called an n-dimensional C^k -smooth (resp. topological) manifold if there exist subsets $U_{\lambda} \subset X$, $\lambda \in \Lambda$, where Λ is a finite or countable set of indices, and for every $\lambda \in \Lambda$ a map $\Phi_{\lambda}: U_{\lambda} \to \mathbb{R}^n$ such that

M1.
$$M = \bigcup_{\lambda \in \Lambda} U_{\lambda}$$
.

M2. The image $G_{\lambda} = \Phi_{\lambda}(U_{\lambda})$ is an open set in \mathbb{R}^n .

M3. The map Φ_{λ} viewed as a map $U_{\lambda} \to G_{\lambda}$ is one-to-one.

M4. For any two sets $U_{\lambda}, U_{\mu}, \lambda, \mu \in \Lambda$ the images $\Phi_{\lambda}(U_{\lambda} \cap U_{\mu}), \Psi_{\mu}(U_{\lambda} \cap U_{\mu}) \subset \mathbb{R}^{n}$ are open and the map

$$h_{\lambda\mu} := \Phi_{\mu} \circ \Phi_{\lambda}^{-1} : \Phi_{\lambda}(U_{\lambda} \cap U_{\mu}) \to \Phi_{\mu}(U_{\lambda} \cap U_{\mu}) \subset \mathbb{R}^{n}$$

is a C^k -diffeomorphism (resp. homeomorphism).

Sets U_{λ} are called *coordinate neighborhoods* and maps $\Phi_{\lambda}: U_{\lambda} \to \mathbb{R}^n$ are called *coordinate maps*. The pairs $(U_{\lambda}, \Phi_{\lambda})$ are also called *local coordinate charts*. The maps $h_{\lambda\mu}$ are called *transiton maps* between different coordinate charts. The inverse maps $\Psi_{\lambda} = \Phi_{\lambda}^{-1}: G_{\lambda} \to U_{\lambda}$ are called *(local) parameterization maps*. An *atlas* is a collection $\mathfrak{A} = \{U_{\lambda}, \Phi_{\lambda}\}_{\lambda \in \Lambda}$ of all coordinate charts.

One says that two atlases $\mathfrak{A} = \{U_{\lambda}, \Phi_{\lambda}\}_{{\lambda} \in {\Lambda}}$ and $\mathfrak{A}' = \{U'_{\gamma}, \Phi'_{\gamma}\}_{{\gamma} \in {\Gamma}}$ on the same manifold X are equivalent, or that they define the same smooth structure on X if their union $\mathfrak{A} \cup \mathfrak{A}' = \{U_{\lambda}, \Phi_{\lambda}\}_{{\lambda} \in {\Lambda}}$

 $\{(U_{\lambda}, \Phi_{\lambda}), (U'_{\gamma}, \Phi'_{\gamma})\}_{\lambda \in \Lambda, \gamma \in \Gamma}$ is again an atlas on X. In other words, two atlases define the same smooth structure if transition maps from local coordinates in one of the atlases to the local coordinates in the other one are given by smooth functions.

A subset $G \subset M$ is called *open* if for every $\lambda \in \Lambda$ the image $\Phi_{\lambda}(G \cap U_{\lambda}) \subset \mathbb{R}^n$ is open. In particular, coordinate charts U_{λ} themselves are open, and we can equivalently say that a set G is open if its intersection with every coordinate chart is open. By a *neighborhood* of a point $a \in M$ we will mean any open subset $U \subset M$ such that $a \in U$.

Given two smooth manifolds M and \widetilde{M} of dimension m and n then a map $f: M \to \widetilde{M}$ is called *continuous* if if for every point $a \in M$ there exist local coordinate charts $(U_{\lambda}, \Phi_{\lambda})$ in M and $(\widetilde{U}_{\lambda}, \widetilde{\Phi}_{\lambda})$ in \widetilde{M}_{λ} , such that $a \in U_{\lambda}, f(U_{\lambda}) \subset \widetilde{U}_{\lambda}$ and the composition map

$$G_{\lambda} = \Phi_{\lambda}(U_{\lambda}) \stackrel{\Psi_{\lambda}}{\to} U_{\lambda} \stackrel{f}{\to} \widetilde{U}_{\lambda} \stackrel{\widetilde{\Phi}_{\lambda}}{\to} \mathbb{R}^{n}$$

is continuous.

Similarly, for $k=1,\ldots,\infty$ a map $f:M\to \widetilde{M}$ is called C^k -smooth if for every point $a\in M$ there exist local coordinate charts (U_λ,Φ_λ) in M and $(\widetilde{U}_\lambda,\widetilde{\Phi}_\lambda)$ in \widetilde{M}_λ , such that $a\in U_\lambda, f(U_\lambda)\subset \widetilde{U}_\lambda$ and the composition map

$$G_{\lambda} = \Phi_{\lambda}(U_{\lambda}) \stackrel{\Psi_{\lambda}}{\to} U_{\lambda} \stackrel{f}{\to} \widetilde{U}_{\lambda} \stackrel{\widetilde{\Phi}_{\lambda}}{\to} \mathbb{R}^{n}$$

is C^k -smooth. In other words, a map is continuous or smooth, if it is continuous or smooth when expressed in local coordinates.

A map $f: M \to N$ is called a *diffeomorphism* if it is smooth, one-to-one, and the inverse map is also smooth. One-to-one continuous maps with continuous inverses are called *homeomorphisms*.

Note that in view of the chain rule the C^k -smoothness is independent of the choice of local coordinate charts $(U_{\lambda}, \Phi_{\lambda})$ and $(\widetilde{U}_{\lambda}, \widetilde{\Phi}_{\lambda})$. Note that for C^k -smooth manifolds one can talk only about C^l -smooth maps for $l \leq k$. For topological manifolds one can talk only about continuous maps.

If one replaces condition M2 in the definition of a manifold by

 $M2^b$. The image $G_{\lambda} = \Psi_{\lambda}(U_{\lambda})$ is either an open set in \mathbb{R}^n or an intersection of an open set in \mathbb{R}^n with $\mathbb{R}^n_+ = \{x_1 \geq 0\}$

then one gets a definition of a manifold with boundary.

A slightly awkward nuance in the above definition is that a manifold with boundary is not a manifold! It would be, probably, less confusing to write this as a 1 word manifold-with-boundary, but of course nobody does that.

The points of a manifold M with boundary which are mapped by coordinate maps Ψ_{λ} to points in $\mathbb{R}^{n-1} = \partial \mathbb{R}^n_+$ are called the *boundary points* of M. The set of boundary points is called *the* boundary of M and denoted by ∂M . It is itself a manifold of dimension n-1.

Note that any (interior) point a of an n-dimensional manifold M has a neighborhood B diffeomorphic to an open ball $B_1(0) \subset \mathbb{R}^n$, while any boundary point has a neighborhood diffeomorphic to a semi-ball $B_1(0) \cap \{x_1 \geq 0\} \subset \mathbb{R}^n$.

Exercise 9.39. Prove that a boundary point *does not have* a neighborhood diffeomorphic to an open ball. In other words, the notion of boundary and interior point of a manifold with boundary are well defined.

Next we want to introduce a notion of compactness for subsets in a manifold. Let us recall that for subsets in a Euclidean vector space we introduced three equivalent definition of compactness, see Section 6.1. The first definition, COMP1 is unapplicable because we cannot talk about bounded sets in a manifold. However, definitions COMP2 and COMP3 make perfect sense in an arbitrary manifold. For instance, we can say that a subset $A \subset M$ is compact if from any infinite sequence of points in A one can choose a subsequence converging to a point in A.

A compact manifold (without boundary) is called *closed*. Note that the word closed is used here in a different sense than a closed set. For instance, a closed interval is not a closed manifold because it has a boundary. An open interval or a real line \mathbb{R} is not a closed manifold because it is not compact. On the other hand, a circle, or a sphere S^n of any dimension n is a closed manifold.

The notions of connected and path connected subsets of a manifold are defined in the same way as in an Euclidean space.

9.8.2 Gluing construction

The construction which is described in this section is called *gluing* or *quotient* construction. It provides a rich source of examples of manifolds. We discuss here only very special cases of this

construction.

a) Let M be a manifold and U, U' its two open disjoint subsets. Let us moreover assume that each point $x \in M$ has a neighborhood which does not intersect at least one of the sets U and U'.⁴ Consider a diffeomorphism $f: U \to U'$.

Let us denote by $M/\{f(x) \sim x\}$ the set obtained from M by identifying each point $x \in U$ with its image $f(x) \in U'$. In other words, a point of $M/\{f(x) \sim x\}$ is either a point from $x \in M \setminus (U \cup U')$, or a pair of points (x, f(x)), where $x \in U$. Note that there exists a canonical projection $\pi : M \to M/\{f(x) \sim x\}$. Namely $\pi(x) = x$ if $x \notin U \cup U'$, $\pi(x) = (x, f(x))$ if $x \in U$ and $\pi(x) = (f^{-1}(x), x)$ if $x \in U'$. By our assumption each point $x \in M$ has a coordinate neighborhood $G_x \ni x$ such that $f(G_x \cap U) \cap G_x = \emptyset$. In particular, the projection $\pi|_{G_x} : G_x \to \widetilde{G}_x = \pi(G_x)$ is one-to-one. We will declare by definition that \widetilde{G}_x is a coordinate neighborhood of $\pi(x) \in M/\{f(x) \sim x\}$ and define a coordinate map $\widetilde{\Phi} : \widetilde{G}_x \to \mathbb{R}^n$ by the formula $\widetilde{\Phi} = \Phi \circ \pi^{-1}$. It is not difficult to check that this construction define a structure of an n-dimensional manifold on the set $M/\{f(x) \sim x\}$. We will call the resulted manifold the quotient manifold of M, or say that $M/\{f(x) \sim x\}$ is obtained from M by gluing U with U' with the diffeomorphism f.

Though the described above gluing construction always produce a manifold, the result could be quite pathological, if no additional care is taken. Here is an example of such pathology.

Example 9.40. Let $M = I \cup I'$ be the union of two disjoint open intervals I = (0, 2) and I' = (3, 5). Then M is a 1-dimensional manifold. Denote $U := (0, 1) \subset I, U' := (3, 4) \subset I'$. Consider a diffeomorphism $f: U \to U'$ given by the formula f(t) = t + 3, $t \in U$. Let $\widetilde{M} = M/\{f(x) \sim x\}$ be the corresponding quotient manifold. In other words, \widetilde{M} is the result of gluing the intervals I and I' along their open sub-intervals U and U'. Note that the points $1 \in I$ and $1 \in I'$ are not identified, but $1 - \epsilon, 4 - \epsilon$ are identified for an arbitrary small $1 \in I'$ and $1 \in I'$ are not identified and any neighborhood of $1 \in I'$ have non-empty intersection.

In order to avoid such pathological examples one usually (but not always) requires that manifolds satisfy an additional axiom, called *Hausdorff property*:

M5. Any two distinct points $x, y \in M$ have non-intersecting neighborhoods $U \ni x, G \ni y$.

⁴Here is an example when this condition is not satisfied: M = (0,2), U = (0,1), U' = (1,2). In this case any neighborhood of the point 1 intersect both sets, U and U'.



Figure 9.4: Felix Hausdorff (1868-1942)

In what follows we always assume that the manifolds satisfy the Hausdorff property M5.

Let us make the following general remark about diffeomorphisms $f:(a,b)\to(c,d)$ between two open intervals. Such diffeomorphism is simply a differentiable function whose derivative never vanishes and whose range is equal to the interval (c,d). If derivative is positive then the diffeomorphism is orientation preserving, and it is orientation reversing otherwise. The function f always extends to a continuous (but necessarily differentiable function $\overline{f}:[a,b]\to[c,d]$ such that $\overline{f}(a)=c$ and $\overline{f}(b)=d$ in the orientation preserving case, and $\overline{f}(a)=d$ $\overline{f}(b)=c$ in the orientation reversing case.

Lemma 9.41. Given $a, b, a'b' \in (0,1)$ such that a < b and a' < b' consider an orientation preserving diffeomorphisms $f: (0,a) \to (0,a')$ and $(b,1) \to (b'1)$. Then for any $\widetilde{a} \in (0,a)$ and $\widetilde{b} \in (b,1)$ there exists a diffeomorphism $F: (0,1) \to (0,1)$ which coincides with f on $(0,\widetilde{a})$ and coincides with g on $(\widetilde{b},1)$.

Proof. Choose real numbers $c, \widetilde{c}, \widetilde{d}, d$ such that $a < c < \widetilde{c} < \widetilde{d} < d < b$. Consider a cut-off C^{∞} -function $\theta : (0,1) \to (0,1)$ which is equal to 1 on $(0,\widetilde{a}] \cup [\widetilde{c},\widetilde{d}] \cup [\widetilde{b},1)$ and equal to 0 on $[a,c] \cup [d,b]$. For positive numbers $\epsilon > 0$ and C > 0 (which we will choose later) consider a function $h_{\epsilon,C}$ on

(0,1) defined by the formula

$$h_{\epsilon,C}(x) = \begin{cases} \theta(x)f'(x) + (1 - \theta(x))\epsilon, & x \in (0, a); \\ \epsilon, & x \in [a, c] \cup [d, b]; \\ C\theta(x) + (1 - \theta(x))\epsilon, & x \in (c, d); \\ \theta(x)g'(x) + (1 - \theta(x))\epsilon, & x \in (b, 1). \end{cases}$$

Note that $h_{\epsilon,C}(x) = f'(x)$ on $(0, \tilde{a}]$, $h_{\epsilon,C}(x) = g'(x)$ on $[\tilde{b}, 1)$ and equal to C on $[\tilde{c}, \tilde{d}]$. Define the function $F_{\epsilon,C}: (0,1) \to (0,1)$ by the formula

$$F_{\epsilon,C}(x) = \int_{0}^{x} h(u)du.$$

Note that the derivative $F'_{\epsilon,C}$ is positive, and hence the function $F_{\epsilon,C}$ is strictly increasing. It coincides with f on $(0, \tilde{a}]$ and coincides up to a constant with g on $(\tilde{b}, 1)$. Note that when ϵ and C are small we have $F_{\epsilon,C}(\tilde{b}) < b' < g(\tilde{b})$, and $\lim_{C \to \infty} F_{\epsilon,C}(\tilde{b}) = \infty$. Hence, by continuity one can choose $\epsilon, C > 0$ in such a way that $F_{\epsilon,C}(\tilde{b}) = g(\tilde{b})$. Then the function $F = F_{\epsilon,C}$ is a diffeomorphism $(0,1) \to (0,1)$ with the required properties.

Lemma 9.42. Suppose that a 1-dimensional manifold M (which satisfies the Hausdorff axiom M5) is covered by two coordinate charts, $M = U \cup U'$, with coordinate maps $\Phi : U \to (0,1), \Phi' : U' \to (0,1)$ such that $\Phi(U \cap U') = (0,a) \cup (b,1), \Phi'(U \cap U') = (0,a') \cup (b'1)$ for some $a,a',b,b' \in (0,1)$ with a < b,a' < b'. Then M is diffeomorphic to the circle S^1 .

Proof. Denote by Ψ and Ψ' the parameterization maps Φ^{-1} and $(\Phi')^{-1}$, and set $G := \Phi(U \cap U')$ and $G' := \Phi'(U \cap U')$. Let $h = \Phi' \circ \Psi : G \to G'$ be the transition diffeomorphism. There could be two cases: h((0,a)) = (0,a'), h((b,1)) = (b',1) and h((0,a)) = (b',1), h((b,1)) = (0,a). We will analyze the first case. The second one is similar.

Let \overline{h} be the continuous extension of h to $[0,a] \cup [b,1]$. We claim that $\overline{h}(0) = a', \overline{h}(a) = 0$, $\overline{h}(b) = 1$ and $\overline{h}(1) = b'$. Indeed, assuming otherwise we come to a contradiction with the Hausdorff property M5. Indeed, suppose $\overline{h}(a) = a'$. Note the points $A := \Psi(a), A' := \Psi'(a') \in M$ are disjoint. On the other hand, for any neighborhood $\Omega \ni A$ its image $\Phi(\Omega) \subset I$ contains an interval $(a - \epsilon, a)$,

and similarly for any neighborhood $\Omega' \ni A'$ its image $\Phi'(\Omega) \subset I$ contains an interval $(a' - \epsilon, a')$. for a sufficiently small ϵ . But $h((a - \epsilon, a)) = (a' - \epsilon', a')$ for some $\epsilon' > 0$ and hence

$$\Omega \supset \Psi'((a', a' - \epsilon')) = \Psi' \circ h((a, a - \epsilon)) = \Psi' \circ \Phi' \circ \Phi((a, a - \epsilon)) = \Phi((a, a - \epsilon)) \subset \Omega,$$

i.e. $\Omega \cap \Omega' \neq \emptyset$. In other words, any neighborhoods of the distinct points $A, A' \in M$ intersect, which violates axiom M5. Similarly we can check that h(b) = b'.

Now take the unit circle $S^1 \subset \mathbb{R}^2$ and consider the polar coordinate ϕ on S^1 . Let us define a map $g': U' \to S^1$ by the formula $\phi = -\pi \Phi'(x)$. Thus g' is a diffeomorphism of U' onto an arc of S^1 given in polar coordinates by $-\pi < \phi < 0$. The points $A' = \Psi'(a')$ and $B' = \Psi'(b')$ are mapped to points with polar coordinates $\phi = -\pi a'$ and $\phi = -\pi b'$. On the intersection $U \cap U'$ we can describe the map g' in terms of the coordinate in U. Thus we get a map $f := g' \circ \Psi : (0,a) \cup (b,1) \to S^1$. We have $\overline{f}(0) = g'(A'), f(a) = g'(0), \overline{f}(1) = g'(B'), f(b) = g'(1)$. Here we denoted by \overline{f} the continuous extension of f to $[0,a] \cup [b,1]$. Thus $f((0,a)) = \{-\pi a' < \phi < 0\}$ and $f((b,1) = \{\pi < \phi < 3\pi - \pi b'\}$. Note that the diffeomorphism f is orientation preserving assuming that the circle is oriented counter-clockwise. Using Lemma 9.41 we can find a diffeomorphism F from (0,1) to the arc $\{-\pi a' < \phi < 3\pi - \pi b'\} \subset S^1$ which coincides with f on $(0,\widetilde{a}) \cup (\widetilde{b},1)$ for any $\widetilde{a} \in (0,a)$ and $\widetilde{b} \in (b,1)$. Denote $\widetilde{a}' := h(\widetilde{a}), \widetilde{b}' = h(\widetilde{b})$. Notice that the neighborhoods U and $\widetilde{U}' = \Psi'((\widetilde{a},\widetilde{b}))$ cover M. Hence, the required diffeomorphism $\widetilde{F}: M \to S^1$ we can define by the formula

$$\widetilde{F}(x) = \begin{cases} g(x), & x \in \widetilde{U}'; \\ F \circ \Phi(x), & x \in U. \end{cases}$$

Similarly (and even simpler), one can prove

Lemma 9.43. Suppose that a 1-dimensional manifold M (which satisfies the Hausdorff axiom M5) is covered by two coordinate charts, $M = U \cup U'$, with coordinate maps $\Phi : U \to (0,1), \Phi' : U' \to (0,1)$ such that $U \cap U'$ is connected. Then M is diffeomorphic to the open interval (0,1).

Theorem 9.44. Any (Hausdorff) connected closed 1-dimensional manifold is diffeomorphic to the circle S^1 .

Exercise 9.45. Show that the statement of the above theorem is not true without the axiom M5, i.e. the assumption that the manifold has the Hausdorff property.

Proof. Let M be a connected closed 1-dimensional manifold. Each point $x \in M$ has a coordinate neighborhood U_x diffeomorphic to an open interval. All open intervals are diffeomorphic, and hence we can assume that each neighborhood G_x is parameterized by the interval I=(0,1). Let $\Psi_x:I\to G_x$ be the corresponding parameterization map. We have $\bigcup_{x\in M}U_x=M$, and due to compactness of M we can choose finitely many U_{x_1},\ldots,U_{x_k} such that $\bigcup_{i=1}^k U_{x_i}=M$. We can further assume that none of these neighborhoods is completrely contained inside another one. Denote $U_1:=U_{x_1},\Psi_1:=\Psi_{x_1}$. Note that $U_1\cap\bigcup_{2}^k U_{x_k}\neq\varnothing$. Indeed, if this were the case then due to connectedness of M we would have $\bigcup_{2}^k U_{x_i}=\varnothing$ and hence $M=U_1$, but this is impossible because M is compact. Thus, there exists $i=2,\ldots,k$ such that $U_{x_i}\cap U_1\neq\varnothing$. We set $U_2:=U_{x_i},\Psi_2=\Psi_{x_i}$. Consider open sets $G_{1,2}:=\Psi_1^{-1}(U_1\cap U_2), G_{2,1}=\Psi_2^{-1}(U_1\cap U_2)\subset I$. The transition map $h_{1,2}:=\Psi_2^{-1}\circ\Psi_1|_{G_{1,2}}:G_{1,2}\to G_{2,1}$ is a diffeomorphism.

Let us show that the set $G_{1,2}$ (and hence $G_{2,1}$) cannot contain more that two connected components. Indeed, in that case one of the components of $G_{1,2}$ has to be a subinterval $I' = (a,b) \subset I = (0,1)$ where 0 < a < b < 1. Denote $I'' := h_{1,2}(I')$. Then at least of the boundary values of the transition diffeomorphism $h_{1,2}|_{I'}$, say $\overline{h}_{1,2}(a)$, which is one of the end points of I'', has to be an interior point $c \in I = (0,1)$. We will assume for determinacy that $I'' = (c,d) \subset I$. But this contradicts the Hausdorff property M5. The argument repeats a similar argument in the proof of Lemma 9.42.

Indeed, note that $\Psi_1(a) \neq \Psi_2(c)$. Indeed, $\Psi_1(a)$ belongs to $U_1 \setminus U_2$ and $\Psi_2(c)$ is in $U_2 \setminus U_1$. Take any neighborhood $\Omega \ni \Psi_1(a)$ in M. Then $\Psi^{-1}(\Omega)$ is an open subset of I which contains the point a. Hence $\Psi_1((a, a + \epsilon)) \subset \Omega$, and similarly, for any neighborhood $\Omega' \ni \Psi_2(c)$ in M we have $\Psi_2((c, c + \epsilon)) \subset \Omega'$ for a sufficiently small $\epsilon > 0$. But $\Psi_1((a, a + \epsilon)) = \Psi_2(h_{1,2}((a, a_{\epsilon}))) = \Psi_2(c, c + \epsilon')$, where $c + \epsilon' = h_{1,2}(a + \epsilon)$. Hence $\Omega \cap \Omega' \neq \emptyset$, i.e. any two neighborhoods of two distict points $\Psi_1(a)$ and $\Psi_2(c)$ have a non-empty intersection, which violates the Hausdorff axiom M5.

If $G_{1,2} \subset (0,1)$ consists of two components then the above argument shows that each of these components must be adjacent to one of the ends of the interval I, and the same is true about the

components of the set $G_{2,1} \subset I$. Hence, we can apply Lemma 9.42 to conclude that the union $U_1 \cup U_2$ is diffeomorphic to S^1 . We also notice that in this case all the remaining neighborhoods U_{x_j} must contain in $U_1 \cup U_2$. Indeed, each U_{x_i} which intersects the circle $U_1 \cup U_2$ must be completely contained in it, because otherwise we would again get a contradiction with the Hausdorff property. Hence, we can eleiminate all neighborhoods which intersect $U_1 \cup U_2$. But then no other neighborhoods could be left because otherwise we would have $M = (U_1 \cup U_2) \cup \bigcup_{U_{x_j} \cap (U_1 \cup U_2) = \varnothing} U_{x_j}$, i.e. the manifold M could be presented as a union of two disjoint non-empty open sets which is impossible due to connectedness of M. Thus we conclude that in this case $M = U_1 \cup U_2$ is diffeomorphic to S^1 .

Finally in the case when $G_{1,2}$ consists of 1 component, i.e. when it is connected, one can Use Lemma 9.43 to show that $U_1 \cup U_2$ is diffeomorphic to an open interval. Hence, we get a covering of M by k-1 neighborhood diffeomorphic to S^1 . Continuing inductively this process we will either find at some step two neighborhoods which intersect each other along two components, or continue to reduce the number of neighborhoods. However, at some moment the first situation should occur because otherwise we would get that M is diffeomorphic to an interval which is impossible because by assumption M is compact.

- b) Let M be a manifold, $f: M \to M$ be a diffeomorphism. Suppose that f satisfies the following property: There exists a positive integer p such that for any point $x \in M$ we have $f^p(x) = \underbrace{f \circ f \circ \cdots \circ f}_p(x) = x$, but the points $x, f(x), \dots, f^{p-1}(x)$ are all disjoint. The set $\{x, f(x), \dots, f^{p-1}(x)\} \subset M$ is called the trajectory of the point x under the action of f. It is clear that trajectory of two different points either coincide or disjoint. Then one can consider the quotient space X/f, whose points are trajectories of points of M under the action of f. Similarly to how it was done in a) one can define on M/f a structure of an n-dimensional manifold.
- c) Here is a version of the construction in a) for the case when trajectory of points are infinite. Let $f: M \to M$ be a diffeomorphism which satisfies the following property: for each point $x \in M$ there exists a neighborhood $U_x \ni x$ such that all sets

$$\dots, f^{-2}(U_x), f^{-1}(U_x), U_x, f(U_x), f^2(U_x), \dots$$

are mutually disjoint. In this case the trajectory $\{\ldots, f^{-2}(x), f^{-1}(x), x, f(x), f^2(x), \ldots\}$ of each point is infinite. As in the case b), the trajectories of two different points either coincide or disjoint.

The set M/f of all trajectories can again be endowed with a structure of a manifold of the same dimension as M.

9.8.3 Examples of manifolds

1. *n*-dimensional sphere S^n . Consider the unit sphere $S^n = \{||x|| = \sqrt{\sum_{1}^{n+1} x_j^2} = 1\} \subset \mathbb{R}^{n+1}$. Let introduce on S^n the structure of an *n*-dimensional manifold. Let $N = (0, \dots, 1)$ and $S = (0, \dots, -1)$ be the North and South poles of S^n , respectively.

Denote $U_- = S^n \setminus S, U_+ = S^n \setminus N$ and consider the maps $p_{\pm} : U_{\pm} \to \mathbb{R}^n$ given by the formula

$$p_{\pm}(x_1, \dots, x_{n+1}) = \frac{1}{1 \mp x_{n+1}}(x_1, \dots, x_n). \tag{9.8.1}$$

The maps $p_+:U_+\to\mathbb{R}_n$ and $p_-:U_-\to\mathbb{R}^n$ are called *stereographic projections* from the North and South poles, respectively. It is easy to see that stereographic projections are one-to one maps. Note that $U_+\cap U_-=S^n\setminus\{S,N\}$ and both images, $p_+(U_+\cap U_-)$ and $p_-(U_+\cap U_-)$ coincide with $\mathbb{R}^n\setminus 0$. The map $p_-\circ p_+^{-1}:\mathbb{R}^n\setminus 0\to\mathbb{R}^n\setminus 0$ is given by the formula

$$p_{-} \circ p_{+}^{-1}(x) = \frac{x}{||x||^2},\tag{9.8.2}$$

and therefore it is a diffeomorphism $\mathbb{R}^n \setminus 0 \to \mathbb{R}^n \setminus 0$.

Thus, the atlas which consists of two coordinate charts (U_+, p_+) and (U_-, p_-) defines on S^n a structure of an n-dimensional manifold. One can equivalently defines the manifold S^n as follows. Take two disjoint copies of \mathbb{R}^n , let call them \mathbb{R}^n_1 and \mathbb{R}^n_2 . Denote $M = \mathbb{R}^n_1 \cup \mathbb{R}^n_2$, $U = \mathbb{R}^n_1 \setminus 0$ and $U' = \mathbb{R}^n_2 \setminus 0$. Let $f: U \to U'$ be a difeomorphism defined by the formula $f(x) = \frac{x}{||x||^2}$, as in (9.8.2). Then S^n can be equivalently described as the quotient manifold M/f.

Note that the 1-dimensional sphere is the circle S^1 . It can be d as follows. Consider the map $T: \mathbb{R} \to \mathbb{R}$ given by the formula T(x) = x + 1, $x \in \mathbb{R}$. It satisfies the condition from 9.8.2 and hence, one can define the manifold \mathbb{R}/T . This manifold is diffeomorphic to S^1 .

2. **Real projective space.** The real projective space $\mathbb{R}P^n$ is the set of all lines in \mathbb{R}^{n+1} passing through the origin. One introduces on $\mathbb{R}P^n$ a structure of an n-dimensional manifold as follows. For each $j=1,\ldots,n+1$ let us denote by U_j the set of lines which are not parallel to the affine subspace $\Pi_j = \{x_j = 1\}$. Clearly, $\bigcup_{1}^{n+1} U_j = \mathbb{R}P^n$. There is a natural one-to one map $\pi_j : U_j \to \Pi_j$

which associates with each line $\mu \in U_j$ the unique intersection point of μ with Π_j . Furthermore, each Π_j can be identified with \mathbb{R}^n , and hence pairs (U_j, π_j) , $j = 1, \ldots, n+1$ can be chosen as an atlas of coordinate charts. We leave it to the reader to check that this atlas indeed define on $\mathbb{R}P^n$ a structure of a smooth manifold, i.e. that the transition maps between different coordinate charts are smooth.

Exercise 9.46. Let us view S^n as the unit sphere in \mathbb{R}^{n+1} . Consider a map $p: S^n \to \mathbb{R}P^n$ which associates to a point of S^n the line passing through this point and the origin. Prove that this two-to-one map is smooth, and moreover a *local diffeomorphism*, i.e. that the restriction of p to a sufficiently small neighborhood of each point is a diffeomorphism. Use it to show that $\mathbb{R}P^n$ is diffeomorphic to the quotient space S^n/f , where $f: S^n \to S^n$ is the antipodal map f(x) = -x.

3. Products of manifolds and n-dimensional tori. Given two manifolds, M and N of dimension m and n, respectively, one can naturally endow the $direct\ product$

$$M \times N = \{(x, y); x \in M, y \in N\}$$

with a structure of a manifold of dimension m+n. Let $\{(U_{\lambda}, \Phi_{\lambda})\}_{{\lambda} \in \Lambda}$ and $\{(V_{\gamma}, \Psi_{\gamma})\}_{{\gamma} \in \Gamma}$ be at lases for M and N, so that $\Phi_{\lambda}: U_{\lambda} \to U'_{\lambda} \subset \mathbb{R}^m$, $\Psi_{\mu}: V_{\mu} \to V'_{\mu} \subset \mathbb{R}^n$ are diffeomorphisms on open subsets of \mathbb{R}^m and \mathbb{R}^n . Then the smooth structure on $M \times N$ can be defined by an atlas

$$\{(U_{\lambda}\times V_{\gamma},\Phi_{\lambda}\times \Psi_{\gamma})\}_{\lambda\in\Lambda,\gamma\in\Gamma},$$

where we denote by $\Phi_{\lambda} \times \Psi_{\gamma} : U_{\lambda} \times V_{\gamma} \to U'_{\lambda} \times V'_{\gamma} \subset \mathbb{R}^{m} \times \mathbb{R}^{n} = \mathbb{R}^{m+n}$ are diffeomorphisms defined by the formula $(x, y) \mapsto (\Phi_{\lambda}(x)\Psi_{\mu}(y))$ for $x \in U_{\lambda}$ and $y \in V_{\mu}$.

One can similarly define the direct product of any finite number of smooth manifolds. In particular the *n*-dimensional torus T^n is defined as the product of *n* circles: $T^n = \underbrace{S^1 \times \cdots \times S^1}_n$. Let us recall, that the circle S^1 is diffeomorphic to \mathbb{R}/T , i.e. a point of S^1 is a real number up to adding any integer. Hence, the points of the torus T^n can viewed as the points of \mathbb{R}^n up to adding any vector with all integer coordinates.

9.8.4 Submanifolds of an *n*-dimensional vector space

Let V be an n-dimensional vector space. A subset $A \subset V$ is called a k-dimensional submanifold of V, or simply a k-submanifold of V, $0 \le k \le n$, if for any points $a \in A$ there exists a local coordinate chart $(U_a, u = (u_1, \ldots, u_n) \to \mathbb{R}^n)$ such that u(a) = 0 (i.e. the point a is the origin in this coordinate system) and

$$A \cap U_a = \{ u = (u_1, \dots, u_n) \in U_a; \ u_{k+1} = \dots = u_n = 0 \}.$$
 (9.8.3)

We will always assume the local coordinates at least as smooth as necessary for our purposes (but at least C^1 -smooth), but more precisely, one can talk of C^m - submanifolds if the implied coordinate systems are at least C^m -smooth.

Note that in the above we can replace the vector space V by any n-dimensional manifold, and thus will get a notion of a k-dimensional submanifold of an n-dimensional manifold V.

Example 9.47. Suppose a subset $A \subset U \subset V$ is given by equations $F_1 = \cdots = F_{n-k} = 0$ for some C^m -smooth functions F_1, \ldots, F_{n-k} on U. Suppose that for any point $a \in A$ the differentials $d_a F_1, \ldots, d_a F_{n-k} \in V_a^*$ are linearly independent. Then $A \subset U$ is a C^m -smooth submanifold.

Indeed, for each $a \in A$ one can choose a linear functions $l_1, \ldots, l_k \in V_a^*$ such that together with $d_a F_1, \ldots, d_a F_{n-k} \in V_a^*$ they form a basis of V^* . Consider functions $L_1, \ldots, L_k : V \to \mathbb{R}$, defined by $L_j(x) = l_j(x-a)$ so that $d_a(L_j) = l_j$, $j = 1, \ldots, k$. Then the Jacobian det $D_a F$ of the map $F: (L_1, \ldots, L_k, F_1, \ldots, F_{n-k}) : U \to \mathbb{R}^n$ does not vanish at a, and hence the inverse function theorem implies that this map is invertible in a smaller neighborhood $U_a \subset U$ of the point $a \in A$. Hence, the functions $u_1 = L_1, \ldots, u_k = L_k, u_{k+1} = F_1, \ldots, u_n = F_{n-k}$ can be chosen as a local coordinate system in U_a , and thus $A \cap U_a = \{u_{k+1} = \cdots = u_n = 0\}$.

Note that the map $u' = (u_1, \ldots, u_k)$ maps $U_a^A = U_a \cap A$ onto an open neighborhood $\widetilde{U} = u(U_a) \cap \mathbb{R}^k$ of the origin in $\mathbb{R}^k \subset \mathbb{R}^n$, and therefore $u' = (u_1, \ldots u_k)$ defines a local coordinates, so that the pair (U_a^A, u') is a coordinate chart. The restriction $\widetilde{\phi} = \phi|_{\widetilde{U}}$ of the parameterization map $\phi = u^{-1}$ maps \widetilde{U} onto U_a^A . Thus $\widetilde{\phi}$ a parameterization map for the neighborhood U_a^A . The atlas $\{(U_a^A, u')\}a \in A$ defines on a a structure of a k-dimensional manifold. The complementary dimension n - k is called the codimension of the submanifold A. We will denote dimension and codimension of A by dim A and codimA, respectively.

As we already mentioned above in Section 9.3 1-dimensional submanifolds are usually called curves. We will also call 2-dimensional submanifolds surfaces and codimension 1 submanifolds hypersurfaces. Sometimes k-dimensional submanifolds are called k-surfaces. Submanifolds of codimension 0 are open domains in V.

An important class form graphical k-submanifolds. Let us recall that given a map $f: B \to \mathbb{R}^{n-k}$, where B is a subset $B \subset \mathbb{R}^k$, then graph is the set

$$\Gamma_f = \{(x, y) \in \mathbb{R}^k \times \mathbb{R}^{n-k} = \mathbb{R}^n; \ x \in B, y = f(x)\}.$$

A (C^m) -submanifold $A \subset V$ is called *graphical* with respect to a splitting $\Phi : \mathbb{R}^k \times \mathbb{R}^{n-k} \to V$, if there exist an open set $U \subset \mathbb{R}^k$ and a (C^m) -smooth map $f : U \to \mathbb{R}^{n-k}$ such that

$$A = \Phi(\Gamma_f).$$

In other words, A is graphical if there exists a coordinate system in V such that

$$A = \{x = (x_1, \dots, x_n); (x_1, \dots, x_k) \in U, x_j = f_j(x_1, \dots, x_k), \ j = k+1, \dots, n\}.$$

for some open set $U \subset \mathbb{R}^k$ and smooth functions, $f_{k+1}, \ldots, f_n : U \to \mathbb{R}$.

For a graphical submanifold there is a global coordinate system given by the projection of the submanifold to \mathbb{R}^k .

It turns out that that any submanifold locally is graphical.

Proposition 9.48. Let $A \subset V$ be a submanifold. Then for any point $a \in A$ there is a neighborhood $U_a \ni a$ such that $U_a \cap A$ is graphical with respect to a splitting of V. (The splitting may depend on the point $a \in A$).

We leave it to the reader to prove this proposition using the implicit function theorem.

One can generalize the discussion in this section and define submanifolds of any manifold M, and not just the vector space V. In fact, the definition (9.8.3) can be used without any changes to define submanifolds of an arbitrary smooth manifold.

A map $f: M \to Q$ is called an *embedding* of a manifold M into another manifold Q if it is a diffeomorphism of M onto a submanifold $A \subset Q$. In other words, f is an embedding if the image A = f(M) is a submanifold of Q and the map f viewed as a map $M \to A$ is a diffeomorphism.

One can prove that any n-dimensional manifold can be embedded into \mathbb{R}^N with a sufficiently large N (in fact N=2n+1 is always sufficient).

Hence, one can think of manifold as submanifold of some \mathbb{R}^n given up to a diffeomorphism, i.e. ignoring how this submanifold is embedded in the ambient space.

In the exposition below we mostly restrict our discussion to submanifolds of \mathbb{R}^n rather than general abstract manifolds.

9.8.5 Submanifolds with boundary

A slightly different notion is of a submanifold with boundary. A subset $A \subset V$ is called a k-dimensional submanifold with boundary, or simply a k-submanifold of V with boundary, $0 \le k < n$, if for any points $a \in A$ there is a neighborhood $U_a \ni a$ in V and local (curvi-linear) coordinates (u_1, \ldots, u_n) in U_a with the origin at a if one of two conditions is satisfied: condition (9.8.3), or the following condition

$$A \cap U_a = \{ u = (u_1, \dots, u_n) \in U_a; \ u_1 \ge 0, u_{k+1} = \dots = u_n = 0 \}.$$
 (9.8.4)

In the latter case the point a is called a *boundary* point of A, and the set of all boundary points is called the *boundary* of A and is denoted by ∂A .

As in the case of submanifolds without boundary, any submanifold with boundary has a structure of a manifold with boundary.

Exercise 9.49. Prove that if A is k-submanifold with boundary then ∂A is a (k-1)-dimensional submanifold (without boundary).

- **Remark 9.50.** 1. As we already pointed out when we discussed manifolds with boundary, a submanifold with boundary is not a submanifold!
 - 2. As it was already pointed out when we discussed 1-dimensional submanifolds with boundary, the boundary of a k-submanifold with boundary is not the same as its set-theoretic boundary, though traditionally the same notation ∂A is used. Usually this should be clear from the context, what the notation ∂A stands for in each concrete case. We will explicitly point this difference out when it could be confusing.

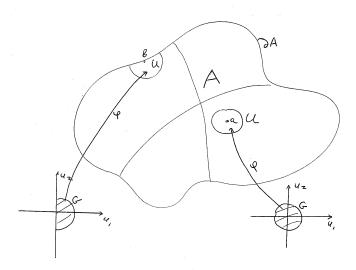


Figure 9.5: The parameterization ϕ introducing local coordinates near an interior point a and a boundary point b.

A compact manifold (without boundary) is called *closed*. The boundary of any compact manifold with boundary is closed, i.e. $\partial(\partial A) = \emptyset$.

Example 9.51. An open ball $B_r^n = B_R^n(0) = \{\sum_1^n x_j^2 < 1\} \subset \mathbb{R}^n$ is a codimension 0 submanifold, A closed ball $D_r^n = D_R^n(0) = \{\sum_1^n x_j^2 \le 1\} \subset \mathbb{R}^n$ ia codimension 0 submanifold with boundary. Its boundary ∂D_R^n is an (n-1)-dimensional sphere $S_R^{n-1} = \{\sum_1^n x_j^2 = 1\} \subset \mathbb{R}^n$. It is a closed hypersurface. For $k = 0, 1 \dots n-1$ let us denote by L^k the subspace $L_k = \{x_{k+1} = \dots = x_n = 0\} \subset \mathbb{R}^n$. Then the intersections

$$B_R^k = B_R^n \cap L^k, D_R^k = D_R^n \cap L^k, \text{ and } S_R^{k-1} = S_R^{n-1} \cap L^k \subset \mathbb{R}^n$$

are, respectively a k-dimensional submanifold, a k-dimensional submanifold with boundary and a closed (k-1)-dimensional submanifold of \mathbb{R}^n . Among all above examples there is only one (which one?) for which the manifold boundary is the same as the set-theoretic boundary.

A neighborhood of a boundary point $a \in \partial A$ can be always locally parameterized by the semi-

open upper-half ball

$$B_{+}(0) = \{x = (x_1, \dots, x_k) \in \mathbb{R}^k; \ x_1 \ge 0, \sum_{j=1}^n x_j^2 < 1\}.$$

We will finish this section by defining submanifolds with piece-wise smooth boundary. A subset $A \subset V$ is called a k-dimensional submanifold of V with piecewise smooth boundary or with boundary with corners, $0 \le k < n$, if for any points $a \in A$ there is a neighborhood $U_a \ni a$ in V and local (curvi-linear) coordinates (u_1, \ldots, u_n) in U_a with the origin at a if one of three condittions satisfied: conditions (9.8.3), (9.8.4) or the following condition

$$A \cap U_a = \{ u = (u_1, \dots, u_n) \in U_a; \ l_1(u) \ge 0, \dots, l_m(u) \ge 0, u_{k+1} = \dots = u_n = 0 \},$$
 (9.8.5)

where m > 1 and $l_1, \ldots, l_m \in (\mathbb{R}^k)^*$ are linear functions In the latter case the point a is called a corner point of ∂A .

Note that the system of linear inequalities $l_1(u) \geq 0, \ldots, l_m(u) \geq 0$ defines a convex cone in \mathbb{R}^k . Hence, near a corner point of its boundary the manifold is diffeomorphic to a convex cone. Thus convex polyhedra and their diffeomorphic images are important examples of submanifolds with boundary with corners.

9.9 Tangent spaces and differential

Suppose we are given two local parameterizations $\phi: G \to A$ and $\widetilde{\phi}: \widetilde{G} \to A$. Suppose that $0 \in G \cap \widetilde{G}$ and $\phi(0) = \widetilde{\phi}(0) = a \in A$. Then there exists a neighborhood $U \ni a$ in A such that $U \subset \phi(G) \cap \widetilde{\phi}(\widetilde{G})$.

Denote $G_1 := \phi^{-1}(U)$, $\widetilde{G}_1 := \widetilde{\phi}^{-1}(\widetilde{U})$. Then one has two coordinate charts on $U : u = (u_1, \dots, u_k) = (\phi|_{G_1})^{-1} : U \to G_1$, and $\widetilde{u} = (\widetilde{u}_1, \dots, \widetilde{u}_k) = (\widetilde{\phi}|_{\widetilde{G}_1})^{-1} : U \to \widetilde{G}_1$.

Denote $h := u \circ \widetilde{\phi}|_{\widetilde{G}_1} = \widetilde{G}_1 \to G_1$. We have

$$\widetilde{\phi} = \phi \circ u \circ \widetilde{\phi} = \phi \circ h,$$

and hence the differentials $d\phi_0$ and $d\widetilde{\phi}_0$ of parameterizations ϕ and $\widetilde{\phi}$ at the origin map \mathbb{R}^k_0 isomorphically onto the same k-dimensional linear subspace $T \subset V_a$. Indeed, $d_0\widetilde{\phi} = d_0\phi \circ d_0h$. Thus the space $T = d_0\phi(\mathbb{R}^k_0) \subset V_a$ is independent of the choice of parameterization. It is called the *tangent*

space to the submanifold A at the point $a \in A$ and will be denoted by T_aA . If A is a submanifold with boundary and $a \in \partial A$ then there are defined both the k-dimensional tangent space T_aA and its (k-1)-dimensional subspace $T_a(\partial A) \subset T_aA$ tangent to the boundary.

Example 9.52. 1. Suppose a submanifold $A \subset V$ is globally parameterized by an embedding $\phi: G \to A \hookrightarrow V, G \subset \mathbb{R}^k$. Suppose the coordinates in \mathbb{R}^k are denoted by (u_1, \ldots, u_k) . Then the tangent space T_aA at a point $a = \phi(b), b \in G$ is equal to the span

Span
$$\left(\frac{\partial \phi}{\partial u_1}(a), \dots, \frac{\partial \phi}{\partial u_k}(a)\right)$$
.

2. In particular, suppose a submanifold A is graphical and given by equations

$$x_{k+1} = g_1(x_1, \dots, x_k), \dots, x_n = g_{n-k}(x_1, \dots, x_k), (x_1, \dots, x_k) \in G \subset \mathbb{R}^k.$$

Take points $b=(b_1,\ldots b_k)\in G$ and $a=(b_1,\ldots ,b_k,g_1(b),\ldots ,g_{n-k}(b))\in A$. Then $T_aA=\operatorname{Span}\left(T_1,\ldots T_k\right)$, where

$$T_1 = \left(\underbrace{1,0,\ldots,0}_{k}, \frac{\partial g_1}{\partial x_1}(b),\ldots, \frac{\partial g_{n-k}}{\partial x_1}(b)\right),$$

$$T_1 = \left(\underbrace{0, 1, \dots, 0}_{k}, \frac{\partial g_1}{\partial x_2}(b), \dots, \frac{\partial g_{n-k}}{\partial x_2}(b)\right),$$

. .

$$T_1 = \left(\underbrace{0,0,\ldots,1}_{k}, \frac{\partial g_1}{\partial x_k}(b),\ldots, \frac{\partial g_{n-k}}{\partial x_k}(b)\right).$$

3. Suppose a hypersurface $\Sigma \subset \mathbb{R}^n$ is given by an equation $\Sigma = \{F = 0\}$ for a smooth function F defined on an neighborhood of Σ and such that $d_a F \neq 0$ for any $a \in \Sigma$. In other words, the function F has no critical points on Σ . Take a point $a \in \Sigma$. Then $T_a \Sigma \subset \mathbb{R}^n_a$ is given by a linear equation

$$\sum_{1}^{n} \frac{\partial F}{\partial x_{j}}(a)h_{j} = 0, \quad h = (h_{1}, \dots, h_{n}) \in \mathbb{R}_{a}^{n}.$$

Note that sometimes one is interested to define $T_a\Sigma$ as an affine subspace of $\mathbb{R}^n=\mathbb{R}^n_0$ and not as a linear subspace of \mathbb{R}^n_a . We get the required equation by shifting the origin:

$$T_a\Sigma = \{x = (x_1, \dots, x_n) \subset \mathbb{R}^n; \sum_{1}^n \frac{\partial F}{\partial x_j}(a)(x_j - a_j) = 0\}.$$

If for some parameterization $\phi: G \to A$ with $\phi(0) = a$ the composition $f \circ \phi$ is differentiable at 0, and the linear map

$$d_0(f \circ \phi) \circ (d_0\phi)^{-1} : T_aA \to W_{f(a)}$$

is called the differential of f at the point a and denoted, as usual, by $d_a f$. Similarly one can define C^m -smooth maps $A \to W$.

Exercise 9.53. Show that a map $f: A \to W$ is differentiable at a point $a \in A$ iff for some neighborhood U of a in V there exists a map $F: U \to W$ that is differentiable at a and such that $F|_{U\cap A} = f|_{U\cap A}$, and we have $dF|_{T_aA} = d_af$. As it follows from the above discussion the map $dF|_{T_aA}$ is independent of this extension. Similarly, any C^m -smooth map of A locally extends to a C^m -smooth map of a neighborhood of A in V.

Suppose that the image f(A) of a smooth map $f:A\to W$ is contained in a submanifold $B\subset W$. In this case the image $d_af(T_aA)$ is contained in $T_{f(a)}B$. Hence, given a smooth map $f:A\to B$ between two submanifolds $A\subset V$ and $B\subset W$ its differential at a point a can be viewed as a linear map $d_af:T_aA\to T_{f(a)}B$.

Let us recall, that given two submanifolds $A \subset V$ and $B \subset W$ (with or without boundary), a smooth map $f: A \to B$ is called a diffeomorphism if there exists a smooth inverse map $: B \to A$, i.e. $f \circ g: B \to B$ and $g \circ f: A \to A$ are both identity map. The submanifolds A and B are called in this case diffeomorphic.

Exercise 9.54. 1. Let A, B be two diffeomorphic submanifolds. Prove that

- (a) if A is path-connected then so is B;
- (b) if A is compact then so is B;
- (c) if $\partial A = \emptyset$ then $\partial B = \emptyset$;
- (d) $\dim A = \dim B$; ⁵
- 2. Give an example of two diffeomorphic submanifolds, such that one is bounded and the other is not.

⁵In fact, we will prove later that even *homeomorphic* manifolds should have the same dimension.

3. Prove that any closed connected 1-dimensional submanifold is diffeomorphic to the unit circle $S^1 = \{x_1^2 + x_2^2 = 1\} \subset \mathbb{R}^2$.

9.10 Vector bundles and their homomorphisms

Let us put the above discussion in a bit more global and general setup.

A collection of all tangent spaces $\{T_aA\}_{a\in A}$ to a submanifold A is called its tangent bundle and denoted by TA or T(A). This is an example of a more general notion of a vector bundle of rank r over a set $A \subset V$. One understands by this a family of r-dimensional vector subspaces $L_a \subset V_a$, parameterized by points of A and continuously (or C^m -smoothly) depending on a. More precisely one requires that each point $a \in A$ has a neighborhood $U \subset A$ such that there exist linear independent vector fields $v_1(a), \ldots, v_r(a) \in L_a$ which continuously (smoothly, etc.) depend on a.

Besides the tangent bundle T(A) over a k-submanifold A an important example of a vector bundle over a submanifold A is its normal bundle NA = N(A), which is a vector bundle of rank n - k formed by orthogonal complements $N_a A = T_a^{\perp} A \subset V_a$ of the tangent spaces $T_a A$ of A. We assume here that V is Euclidean space.

A vector bundle L of rank k over A is called *trivial* if one can find k continuous linearly independent vector fields $v_1(a), \ldots, v_k(a) \in L_a$ defined for all $a \in A$. The set A is called the base of the bundle L.

An important example of a trivial bundle is the bundle $TV = \{V_a\}_{a \in V}$.

Exercise 9.55. Prove that the tangent bundle to the unit circle $S^1 \subset \mathbb{R}^2$ is trivial. Prove that the tangent bundle to $S^2 \subset \mathbb{R}^3$ is not trivial, but the tangent bundle to the unit sphere $S^3 \subset \mathbb{R}^4$ is trivial. (The case of S^1 is easy, of S^3 is a bit more difficult, and of S^2 even more difficult. It turns out that the tangent bundle TS^{n-1} to the unit sphere $S^{n-1} \subset \mathbb{R}^n$ is trivial if and only if n=2,4 and 8. The *only if* part is a very deep topological fact which was proved by F. Adams in 1960.

Suppose we are given two vector bundles, L over A and \widetilde{L} over \widetilde{A} and a continuous (resp. smooth) map $\phi: A \to \widetilde{A}$. By a continuous (resp. smooth) homomorphism $\Phi: L \to \widetilde{L}$ which covers the map $\phi: A \to \widetilde{A}$ we understand a continuous (resp. smooth) family of linear maps $\Phi_a: L_a \to \widetilde{L}_{\phi(a)}$. For instance, a C^m -smooth map $f: A \to B$ defines a C^{m-1} -smooth homomorphism $df: TA \to TB$

which covers $f: A \to B$. Here $df = \{d_a f\}_{a \in A}$ is the family of linear maps $d_a f: T_a A \to T_{f(a)} B$, $a \in A$.

9.11 Orientation

By an orientation of a vector bundle $L = \{L_a\}_{a \in A}$ over A we understand continuously depending on a orientation of all vector spaces L_a . An orientation of a submanifold k is the same as an orientation of its tangent bundle T(A). A co-orientation of a k- submanifold A is an orientation of its normal bundle $N(A) = T^{\perp}A$ in V. Note that not all bundles are orientable, i.e. some bundles admit no orientation. But if L is orientable and the base A is connected, then L admits exactly two orientations. Here is a simplest example of a non-orientable rank 1 bundle of the circle $S^1 \subset \mathbb{R}^2$. Let us identify a point $a \in S^1$ with a complex number $a = e^{i\phi}$, and consider a line $l_a \in \mathbb{R}^2_a$ directed by the vector $e^{i\phi}$. Hence, when the point completes a turn around S^1 the line l_a rotates by the angle π . We leave it to the reader to make a precise argument why this bundle is not orientable. In fact, rank 1 bundles are orientable if and only if they are trivial.

If the ambient space V is oriented then co-orientation and orientation of a submanifold A determine each other according to the following rule. For each point a, let us choose any basis v_1, \ldots, v_k of $T_a(A)$ and any basis w_1, \ldots, w_{n-k} of $N_a(A)$. Then $w_1, \ldots, w_{n-k}, v_1, \ldots, v_k$ is a basis of $V_a = V$. Suppose one of the bundles, say N(A), is oriented. Let us assume that the basis w_1, \ldots, w_{n-k} defines this orientation. Then we orient T_aA by the basis v_1, \ldots, v_k if the basis $w_1, \ldots, w_{n-k}, v_1, \ldots, v_k$ defines the given orientation of V, and we pick the opposite orientation of T_aA otherwise.

Example 9.56. (Induced orientation of the boundary of a submanifold.) Suppose A is an oriented manifold with boundary. Let us co-orient the boundary ∂A by orienting the rank 1 normal bundle to $T(\partial A)$ in T(A) by the unit ourtward normal to $T(\partial A)$ in T(A) vector field. Then the above rule determine an orientation of $T(\partial A)$, and hence of ∂A .

9.12 Integration of differential k-forms over k-dimensional submanifolds

Let α be a differential k-form defined on an open set $U \subset V$.

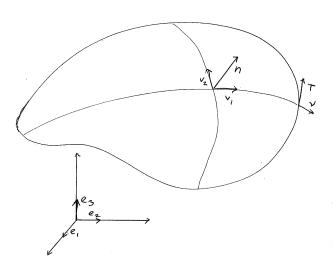


Figure 9.6: The orientation of the surface is induced by its co-orientation by the normal vector \mathbf{n} . The orientation of the boundary us induced by the orientation of the surface.

Consider first a k-dimensional compact submanifold with boundary $A \subset U$ defined parametrically by an embedding $\phi: G \to A \hookrightarrow U$, where $G \subset \mathbb{R}^k$ is possibly with boundary. Suppose that A is oriented by this embedding. Then we define

$$\int_{A} \alpha := \int_{G} \phi^* \alpha.$$

Note that if we define A by a different embedding $\widetilde{\phi}:\widetilde{G}\to A$, then we have $\widetilde{\phi}=\phi\circ h$, where $h=\phi^{-1}\circ\widetilde{\phi}:\widetilde{G}\to G$ is a diffeomorphism. Hence, using Theorem 9.32 we get

$$\int_{\widetilde{G}} \widetilde{\phi}^* \alpha = \int_{\widetilde{G}} h^*(\phi^* \alpha) = \int_{G} \phi^* \alpha,$$

and hence $\int_A \alpha$ is independent of a choice of parameterization, provided that the orientation is preserved.

Let now A be any compact oriented submanifold with boundary. Let us choose a partition of unity $1 = \sum_{1}^{K} \theta_{j}$ in a neighborhood of A such that each function is supported in some coordinate neighborhood of A. Denote $\alpha_{j} = \theta_{j}\alpha$. Then $\alpha = \sum_{1}^{K} \alpha_{j}$, where each form α_{j} is supported in one of coordinate neighborhoods. Hence there exist orientation preserving embeddings $\phi_{j}: G_{j} \to A$ of domains with boundary $G_{j} \subset \mathbb{R}^{k}$, such that $\phi_{j}(G_{j}) \supset \operatorname{Supp}(\alpha_{j}), j = 1, \ldots, K$. Hence, we can define

$$\int_{A} \alpha_{j} := \int_{G_{j}} \phi_{j}^{*} \alpha_{j} \text{ and } \int_{A} \alpha := \sum_{1}^{K} \int_{A} \alpha_{j}.$$

Lemma 9.57. The above definition of $\int_A \alpha$ is independent of a choice of a partition of unity.

Proof. Consider two different partitions of unity $1 = \sum_{1}^{K} \theta_{j}$ and $1 = \sum_{1}^{K} \widetilde{\theta}_{j}$ subordinated to coverings U_{1}, \ldots, U_{K} and $\widetilde{U}_{1}, \ldots, \widetilde{U}_{\widetilde{K}}$, respectively. Taking the product of two partitions we get another partition $1 = \sum_{i=1}^{K} \sum_{j=1}^{\widetilde{K}} \theta_{ij}$, where $\theta_{ij} := \theta_{i}\theta_{j}$, which is subordinated to the covering by intersections $U_{i} \cap U_{j}$, $i = 1, \ldots, K$, $j = 1, \ldots, \widetilde{K}$. Denote $\alpha_{i} := \theta_{i}\alpha$, $\widetilde{\alpha}_{j} := \widetilde{\theta}_{j}\alpha$ and $\alpha_{ij} = \theta_{ij}\alpha$. Then $\sum_{i=1}^{K} \alpha_{ij} = \widetilde{\alpha}_{j}$,

 $\sum_{i=1}^{\widetilde{K}} \alpha_{ij} = \alpha_i \text{ and } \alpha = \sum_{1}^{K} \alpha_i = \sum_{1}^{\widetilde{K}} \widetilde{\alpha}_j. \text{ Then, using the linearity of the integral we get}$

$$\sum_{1}^{K} \int_{A} \alpha_{i} = \sum_{i=1}^{K} \sum_{j=1}^{\widetilde{K}} \int_{A} \alpha_{ij} = \sum_{1}^{\widetilde{K}} \int_{A} \widetilde{\alpha}_{j}.$$

When k = 1 the above definition of the integral coincides with the definition of the integral of a 1-form over an oriented curve which was given above in Section 9.2.

Let us extend the definition of integration of differential forms to an important case of integration of 0-form over oriented 0-dimensional submanifolds. Let us recall a compact oriented 0-dimensional submanifold of V is just a finite set of points $a_1, \ldots, a_m \in V$ with assigned signs to every point. So in view of the additivity of the integral it is sufficient to define integration over 1 point with a sign. On the other hand, a 0-form is just a function $f: V \to \mathbb{R}$. So we define

$$\int_{+a} f := \pm f(a).$$

A partition of unity is a convenient tool for studying integrals, but not so convenient for practical computations. The following proposition provides a more practical method for computations.

Proposition 9.58. Let A be a compact oriented submanifold of V and α a differential k-form given on a neighborhood of A. Suppose that A presented as a union $A = \bigcup_{i=1}^{N} A_j$, where A_j are codimension 0 submanifolds of A with boundary with corners. Suppose that A_i and A_j for any $i \neq j$ intersect only along pieces of their boundaries. Then

$$\int_{A} \alpha = \sum_{1}^{N} \int_{A_{i}} \alpha.$$

In particular, if each A_j is parameterized by an orientation preserving embedding $\phi_j: G_j \to A$, where $G_j \subset \mathbb{R}^k$ is a domain with boundary with corners. Then

$$\int_{A} \alpha = \sum_{1}^{N} \int_{G_{j}} \phi_{j}^{*} \alpha.$$

We leave the proof to the reader as an exercise.

Exercise 9.59. Compute the integral

$$\int_{S} 1/3(x_1 dx_2 \wedge dx_3 + x_2 dx_3 \wedge dx_1 + x_3 dx_1 \wedge dx_2),$$

where S is the sphere

$${x_1^2 + x_2^2 + x_3^2 = 1},$$

cooriented by its exterior normal vector.

Solution. Let us present the sphere as the union of northern and southern hemispheres:

$$S = S_{-} \cup S_{+}$$
, where $S_{-} = S \cap \{x_3 \le 0\}, S_{+} = S \cap \{x_3 \ge 0\}.$

Then $\int\limits_S \omega = \int\limits_{S^+} \omega + \int\limits_{S_-} \omega.$ Let us first compute $\int\limits_{S^+} \omega.$

We can parametrize S_+ by the map $(u, v) \to (u, v, \sqrt{R^2 - u^2 - v^2})$, $(u, v) \in \{u^2 + v^2 \le R^2\} = \mathcal{D}_R$. One can check that this parametrization agrees with the prescribed coordination of S. Thus, we have

$$\int_{S_{+}} \omega = 1/3 \int_{\mathcal{D}_{R}} \left(u dv \wedge d\sqrt{R^{2} - u^{2} - v^{2}} + v d\sqrt{R^{2} - u^{2} - v^{2}} \wedge du + \sqrt{R^{2} - u^{2} - v^{2}} du \wedge dv \right).$$

Passing to polar coordinates (r, φ) in the plane (u, v) we get

$$\int\limits_{S_+} \omega = 1/3 \int\limits_{P} r\cos\varphi d(r\sin\varphi) \wedge d\sqrt{R^2-r^2} + r\sin\varphi d\sqrt{R^2-r^2} \wedge d(r\cos\varphi)$$

$$+\sqrt{R^2-r^2}d(r\cos\varphi)\wedge d(r\sin\varphi),$$

where $P = \{0 \le r \le R, 0 \le \varphi \le 2\pi\}$. Computing this integral we get

$$\int_{S_{+}} \omega = \frac{1}{3} \int_{P} -\frac{r^{3} \cos^{2} \varphi d\varphi \wedge dr}{\sqrt{R^{2} - r^{2}}} + \frac{r^{3} \sin^{2} \varphi dr \wedge d\varphi}{\sqrt{R^{2} - r^{2}}} + \sqrt{R^{2} - r^{2}} dr \wedge d\varphi$$

$$= \frac{1}{3} \int_{P} \left(\frac{r^{3}}{\sqrt{R^{2} - r^{2}}} + r\sqrt{R^{2} - r^{2}} \right) dr \wedge d\varphi$$

$$= \frac{2\pi}{3} \int_{0}^{R} \frac{rR^{2}}{\sqrt{R^{2} - r^{2}}} dr = -\frac{2\pi R^{2}}{3} \sqrt{R^{2} - r^{2}} \int_{0}^{R} = \frac{2\pi R^{3}}{3}$$

Similarly, one can compute that

$$\int_{S} \omega = \frac{2\pi R^3}{3}.$$

Computing this last integral, one should notice the fact that the parametrization

$$(u, v) \mapsto (u, v, -\sqrt{R^2 - u^2 - v^2})$$

defines the wrong orientation of S_{-} . Thus one should use instead the parametrization

$$(u,v) \mapsto (v,u,-\sqrt{R^2-u^2-v^2}),$$

and we get the answer

$$\int_{S} \omega = \frac{4\pi R^3}{3}.$$

This is just the volume of the ball bounded by the sphere. The reason for such an answer will be clear below from Stokes' theorem.

Part III

Stokes' theorem and its applications

Chapter 10

Stokes' theorem

10.1 Statement of Stokes' theorem

Theorem 10.1. Let $A \subset V$ be a compact oriented submanifold with boundary (and possibly with corners). Let ω be a C^2 -smooth differential form defined on a neighborhood $U \supset A$. Then

$$\int_{\partial A} \omega = \int_{A} d\omega.$$

Here $d\omega$ is the exterior differential of the form ω and ∂A is the oriented boundary of A.

We will discuss below what exactly Stokes' theorem means for the case $k \leq 3$ and $n = \dim V \leq 3$.

Let us begin with the case k = 1, n = 2. Thus $V = \mathbb{R}^2$. Let x_1, x_2 be coordinates in \mathbb{R}^2 and U a domain in \mathbb{R}^2 bounded by a smooth curve $\Gamma = \partial U$. Let us co-orient Γ with the outward normal ν to the boundary of U. This defines a counter-clockwise orientation of Γ .

Let $\omega = P_1(x_1, x_2)dx_1 + P_2(x_1, x_2)dx_2$ be a differential 1-form. Then the above Stokes' formula asserts

$$\int_{U} d\omega = \int_{\Gamma} \omega \,,$$

or

$$\int_{U} \left(\frac{\partial P_2}{\partial x_1} - \frac{\partial P_1}{\partial x_2} \right) dx_1 \wedge dx_2 = \int_{\Gamma} P_1 dx_1 + P_2 dx_2.$$



Figure 10.1: George Stokes (1819-1903)

This is called *Green's formula*. In particular, when $d\omega = dx_1 \wedge dx_2$, e.g. $\omega = xdy$ or $\omega = \frac{1}{2}(xdy - ydx)$, the integral $\int_{\Gamma} \omega$ computes the area of the domain U.

Consider now the case n = 3, k = 2. Thus

$$V = \mathbb{R}^3, \ \omega = P_1 dx_2 \wedge dx_3 + P_2 dx_3 \wedge dx_1 + P_3 dx_1 \wedge dx_2.$$

Let $U \subset \mathbb{R}^3$ be a domain bounded by a smooth surface S. We co-orient S with the exterior normal ν . Then

$$d\omega = \left(\frac{\partial P_1}{\partial x_1} + \frac{\partial P_2}{\partial x_2} + \frac{\partial P_3}{\partial x_3}\right) dx_1 \wedge dx_2 \wedge dx_3.$$

Thus, Stokes' formula

$$\int_{S} \omega = \int_{U} d\omega$$

gives in this case

$$\int\limits_{S} P_1 dx_2 \wedge dx_3 + P_2 dx_3 \wedge dx_1 + P_3 dx_1 \wedge dx_2 = \int\limits_{U} \left(\frac{\partial P_1}{\partial x_1} + \frac{\partial P_2}{\partial x_2} + \frac{\partial P_3}{\partial x_3} \right) dx_1 \wedge dx_2 \wedge dx_3$$

This is called the *divergence theorem* or Gauss-Ostrogradski's formula.



Figure 10.2: George Green (1793-1841)

Consider the case k=0, n=1. Thus ω is just a function f on an interval I=[a,b]. The boundary ∂I consists of 2 points: $\partial I=\{a,b\}$. One should orient the point a with the sign - and the point b with the sign +.

Thus, Stokes' formula in this case gives

$$\int_{[a,b]} df = \int_{\{-a,+b\}} f,$$

or

$$\int_{a}^{b} f'(x)dx = f(b) - f(a).$$

This is Newton-Leibnitz' formula. More generally, for a 1-dimensional oriented connected curve

 $\Gamma \subset \mathbb{R}^3$ with boundary $\partial \Gamma = B \cup (-A)$ and any smooth function f we get the formula

$$\int_{\Gamma} df = \int_{B \cup (-A)} f = f(B) - f(A),$$

which we already proved earlier, see Theorem 9.11 .

Consider now the case n = 3, k = 1.

Thus $V = \mathbb{R}^3$ and $\omega = P_1 dx_1 + P_2 dx_2 + P_3 dx_3$. Let $S \subset \mathbb{R}^3$ be an oriented surface with boundary





Figure 10.3: Carl Friedrich Gauss (1777-1855) Mikhail Ostrogradski (1801-1862)

 Γ . We orient Γ in the same way, as in Green's theorem. Then Stokes' formula

$$\int_{S} d\omega = \int_{\Gamma} \omega$$

gives in this case

$$\int_{S} \left(\frac{\partial P_3}{\partial x_2} - \frac{\partial P_2}{\partial x_3} \right) dx_2 \wedge dx_3 + \left(\frac{\partial P_1}{\partial x_3} - \frac{\partial P_3}{\partial x_1} \right) dx_3 \wedge dx_1 + \left(\frac{\partial P_2}{\partial x_1} - \frac{\partial P_1}{\partial x_2} \right) dx_1 \wedge dx_2$$

$$= \int_{S} P_1 dx_1 + P_2 dx_2 + P_3 dx_3.$$

This is the original Stokes' theorem.

Stokes' theorem allows one to clarify the geometric meaning of the exterior differential.

Lemma 10.2. Let β be a differential k-form in a domain $U \subset V$. Take any point $a \in U$ and vectors $X_1, \ldots, X_{k+1} \in V_a$. Given $\epsilon > 0$ let us consider the parallelepiped $P(\epsilon X_1, \ldots, \epsilon X_{k+1})$ as a subset of V with vertices at points $a_{i_1...i_{k+1}} = a + \epsilon \sum_{1}^{k+1} i_j X_j$, where each index i_j takes values 0, 1. Then

$$d\beta_a(X_1, \dots X_{k+1}) = \lim_{\epsilon \to 0} \frac{1}{\epsilon^{k+1}} \int_{\partial P(\epsilon X_1, \dots, \epsilon X_{k+1})} \beta.$$

Proof. First, it follows from the definition of integral of a differential form that

$$d\beta_a(X_1, \dots, X_{k+1}) = \lim_{\epsilon \to 0} \frac{1}{\epsilon^{k+1}} \int_{P(\epsilon X_1, \dots, \epsilon X_{k+1})} d\beta.$$
 (10.1.1)

Then we can continue using Stokes' formula

$$d\beta_a(X_1, \dots, X_{k+1}) = \lim_{\epsilon \to 0} \frac{1}{\epsilon^{k+1}} \int_{P(\epsilon X_1, \dots, \epsilon X_{k+1})} d\beta = \lim_{\epsilon \to 0} \frac{1}{\epsilon^{k+1}} \int_{\partial P(\epsilon X_1, \dots, \epsilon X_{k+1})} \beta.$$
 (10.1.2)

10.2 Proof of Stokes' theorem

We prove in this section Theorem 10.1. We will consider only the case when A is a manifold with boundary without corners and leave the corner case as an exercise to the reader.

Let us cover A by coordinate neighborhoods such that in each neighborhood A is given either by (9.8.3) or (9.8.4). First we observe that it is sufficient to prove the theorem for the case of a form supported in one of these coordinate neighborhoods. Indeed, let us choose finitely many such neighborhoods covering A. Let $1 = \sum_{1}^{N} \theta_{j}$ be a partition of unity subordinated to this covering. We set $\omega_{j} = \theta_{j}\omega$, so that $\omega = \sum_{1}^{N} \omega_{j}$, and each of ω_{j} is supported in one of coordinate neighborhoods. Hence, if formula 10.1 holds for each ω_{j} it also holds for ω .

Let us now assume that ω is supported in one of coordinate neighborhoods. Consider the corresponding parameterization $\phi: G \to U \subset V$, $G \subset \mathbb{R}^n$, introducing coordinates u_1, \ldots, u_n . Then $A \cap U = \phi(G \cap L)$, where L is equal to the subspace $\mathbb{R}^k = \{u_{k+1} = \ldots u_n = 0\}$ in the case (9.8.3) and the upper-half space $\mathbb{R}^k \cap \{u_1 \geq 0\}$. By definition, we have $\int\limits_{A} d\omega = \int\limits_{U} d\omega = \int\limits_{G \cap L} \phi^* d\omega = \int\limits_{G \cap L} d\phi^* \omega$. Though the form $\widetilde{\omega} = \phi^* \omega|_{G \cap L}$ is defined only on $G \cap L$, it is supported in this neighborhood, and hence we can extend it to a smooth form on the whole L by setting it equal to 0 outside the neighborhood. With this extension we have $\int\limits_{G \cap L} d\widetilde{\omega} = \int\limits_{L} d\widetilde{\omega}$. The (k-1)-form $\widetilde{\omega}$ can be written in coordinates u_1, \ldots, u_k as

$$\widetilde{\omega} = \sum_{1}^{j} f_{j}(u) du_{1} \wedge \dots \overset{j}{\underbrace{\cdot}} \dots \wedge du_{k}.$$

¹We assume here that the coordinates u_1, \ldots, u_k define the given orientation of A.

Then

$$\int_{G \cap L} d\widetilde{\omega} = \int_{L} \left(\sum_{1}^{k} \frac{\partial f_{j}}{\partial u_{j}} \right) du_{1} \wedge \dots \wedge du_{k}.$$

Let us choose a sufficiently R > 0 so that the cube $I = \{|u_i| \le R, i = 1, ..., k\}$ contains $\text{Supp}(\widetilde{\omega})$. Thus in the case (9.8.3) we have

$$\int_{G \cap L} d\widetilde{\omega} = \sum_{1}^{k} \int_{\mathbb{R}^{k}} \frac{\partial f_{j}}{\partial u_{j}} dV = \sum_{1}^{k} \int_{-R}^{R} \dots \int_{-R}^{R} \frac{\partial f_{j}}{\partial u_{j}} du_{1} \dots du_{n} =$$

$$\sum_{1}^{k} \int_{-R}^{R} \dots \left(\int_{-R}^{R} \frac{\partial f_{j}}{\partial u_{j}} du_{j} \right) du_{1} \dots du_{j-1} du_{j+1} \dots du_{n} = 0$$
(10.2.1)

because

$$\int_{-R}^{R} \frac{\partial f_j}{\partial u_j} du_j = f_j(u_1, \dots, u_{i-1}, R, u_i, \dots, u_n) - f_j(u_1, \dots, u_{i-1}, -R, u_i, \dots, u_n)) = 0.$$

On the other hand, in this case $\int_{\partial A} \omega = 0$, because the support of ω does not intersect the boundary of A. Hence, Stokes' formula holds in this case. In case (9.8.4) we similarly get

$$\int_{G \cap L} d\widetilde{\omega} = \sum_{1}^{k} \int_{\{u_1 \ge 0\}} \frac{\partial f_j}{\partial u_j} dV =$$

$$\sum_{1}^{k} \int_{0}^{R} \left(\int_{-R}^{R} \dots \int_{-R}^{R} \frac{\partial f_j}{\partial u_j} du_n \dots du_2 \right) du_1 = \int_{-R}^{R} \left(\int_{-R}^{R} \dots \int_{0}^{R} \frac{\partial f_1}{\partial u_1} du_1 \dots du_{n-1} \right) du_n =$$

$$- \int_{-R}^{R} \dots \int_{-R}^{R} f_1(0, u_2, \dots, u_n) du_2 \dots du_n. \tag{10.2.2}$$

because all terms in the sum with j > 1 are equal to 0 by the same argument as in (10.2.1). On the other hand, in this case

$$\int_{\partial A} \omega = \int_{\{u_1=0\}} \phi^* \omega = \int_{\{u_1=0\}} f_1(0, u_2, \dots, u_n) du_2 \wedge \dots \wedge du_n =
- \int_{-R}^{R} \dots \int_{-R}^{R} f_1(0, u_2, \dots, u_n) du_2 \dots du_n.$$
(10.2.3)

The sign minus appears in the last equality in front of the integral because the induced orientation on the space $\{u_1 = 0\}$ as the boundary of the upper-half space $\{u_1 \geq 0\}$ is opposite to the orientation defined by the volume form $du_2 \wedge \cdots \wedge du_n$. Comparing the expressions (10.2.2) and (10.2.3) we conclude that $\int_A d\omega = \int_{\partial A} \omega$, as required.

10.3 Integration of functions over submanifolds

In order to integrate functions over a submanifold we need a notion of volume for subsets of the submanifold.

Let $A \subset V$ be an oriented k-dimensional submanifold, $0 \leq k \leq n$. By definition, the volume form $\sigma = \sigma_A$ of A (or the area form if k = 2, or the length form if k = 1) is a differential k-form on A whose value on any k tangent vectors $v_1, \ldots, v_k \in T_x A$ equals the oriented volume of the parallelepiped generated by these vectors.

Given a function $f:A\to\mathbb{R}$ we define its integral over A by the formula

$$\int_{A} f dV = \int_{A} f \sigma_{A}, \qquad (10.3.1)$$

and, in particular,

$$\operatorname{Vol} A = \int_A \sigma_A$$
.

Notice that the integral $\int_A f dV$ is independent of the orientation of A. Indeed, changing the orientation we also change the sign of the form σ_A , and hence the integral remains unchanged. This allows us to define the integral $\int_A f dV$ even for a non-orientable A. Indeed, we can cover A by coordinate charts, find a subordinated partition of unity and split correspondingly the function $f = \sum_1^N f_j$ in such a way that each function f_j is supported in a coordinate neighborhood. By orienting in arbitrary ways each of the coordinate neighborhoods we can compute each of the integrals $\int_A f_j dV$, $j = 1, \ldots, N$. It is straightforward to see that the integral $\int_A f dV = \sum_j \int_A f_j dV$ is independent of the choice of the partition of unity.

Let us study in some examples how the form σ_A can be effectively computed.

Example 10.3. Volume form of a hypersurface. Let us fix a Cartesian coordinates in V. Let $A \subset V$ is given by the equation

$$A = \{F = 0\}$$

for some function $F: V \to \mathbb{R}$ which has no critical points on A. The vector field ∇F is orthogonal to A, and

$$\mathbf{n} = \frac{\nabla F}{||\nabla F||}$$

is the unit normal vector field to A. Assuming A to be co-oriented by \mathbf{n} we can write down the volume form of A as the contraction of \mathbf{n} with the volume form $\Omega = dx_1 \wedge \cdots \wedge dx_n$ of \mathbb{R}^n , i.e.

$$\sigma_A = \mathbf{n} \, \rfloor \Omega = \frac{1}{||\nabla F||} \sum_{i=1}^{n} (-1)^{i-1} \frac{\partial F}{\partial x_i} dx_1 \wedge \overset{i}{\vee} \cdot \wedge dx_n.$$

In particular, if n=3 we get the following formula for the area form of an implicitely given 2-dimensional surface $A = \{F = 0\} \subset \mathbb{R}^3$:

$$\sigma_{A} = \frac{1}{\sqrt{\left(\frac{\partial F}{\partial x_{1}}\right)^{2} + \left(\frac{\partial F}{\partial x_{2}}\right)^{2} + \left(\frac{\partial F}{\partial x_{3}}\right)^{2}}} \left(\frac{\partial F}{\partial x_{1}} dx_{2} \wedge dx_{3} + \frac{\partial F}{\partial x_{2}} dx_{3} \wedge dx_{1} + \frac{\partial F}{\partial x_{3}} dx_{1} \wedge dx_{2}\right) . \quad (10.3.2)$$

Example 10.4. Length form of a curve.

Let $\Gamma \subset \mathbb{R}^n$ be an oriented curve given parametrically by a map $\gamma : [a,b] \to \mathbb{R}^n$. Let $\sigma = \sigma_{\Gamma}$ be the length form. Let us compute the form $\gamma^*\sigma_{\Gamma}$. Denoting the coordinate in [a,b] by t and the unit vector field on [a,b] by t we have

$$\gamma^* \sigma_{\Gamma} = f(t) dt,$$

where

$$f(t) = \gamma^* \sigma_{\Gamma}(e) = \sigma_{\Gamma} (\gamma'(t)) = ||\gamma'(t)||.$$

In particular the length of Γ is equal to

$$\int_{\Gamma} \sigma_{\Gamma} = \int_{a}^{b} ||\gamma'(t)|| dt = \int_{a}^{b} \sqrt{\sum_{i=1}^{n} (x_i'(t))^2} dt,$$

where

$$\gamma(t) = (x_1(t), \dots, x_n(t)).$$

Similarly, given any function $f:\Gamma\to\mathbb{R}$ we have

$$\int_{\Gamma} f ds = \int_{a}^{b} f(\gamma(t)) ||\gamma'(t)|| dt.$$

Example 10.5. Area form of a surface given parametrically.

Suppose a surface $S \subset \mathbb{R}^n$ is given parametrically by a map $\Phi : U \to \mathbb{R}^n$ where U in the plane \mathbb{R}^2 with coordinates (u, v).

Let us compute the pull-back form $\Phi^*\sigma_S$. In other words, we want to express σ_S in coordinates u, v. We have

$$\Phi^* \sigma_S = f(u, v) du \wedge dv.$$

To determine f(u,v) take a point $z=(u,v)\in\mathbb{R}^2$ and the standard basis $e_1,e_2\in\mathbb{R}^2$. Then

$$(\Phi^* \sigma_S)_z(e_1, e_2) = f(u, v) du \wedge dv(e_1, e_2).$$
(10.3.3)

On the other hand, by the definition of the pull-back form we have

$$(\Phi^* \sigma_S)_z(e_1, e_2) = (\sigma_S)_{\Phi(z)}(d_z \Phi(e_1), d_z \Phi(e_2)).$$
(10.3.4)

But $d_z\Phi(e_1)=\frac{\partial\Phi}{\partial u}(z)=\Phi_u(z)$ and $d_z\Phi(e_2)=\frac{\partial\Phi}{\partial v}(z)=\Phi_v(z)$. Hence from (10.3.3) and (10.3.4) we get

$$f(u,v) = \sigma_S(\Phi_u, \Phi_v). \tag{10.3.5}$$

The value of the form σ_S on the vectors Φ_u, Φ_v is equal to the area of the parallelogram generated by these vectors, because the surface is assumed to be oriented by these vectors, and hence $\sigma_S(\Phi_u, \Phi_v) > 0$. Denoting the angle between Φ_u and Φ_v by α we get $\sigma_S(\Phi_u, \Phi_v) = ||\Phi_u|| \, ||\Phi_v|| \sin \alpha$. Hence

$$\sigma_S(\Phi_u,\Phi_v)^2 = ||\Phi_u||^2 \, ||\Phi_v||^2 \sin^2\alpha = ||\Phi_u||^2 \, ||\Phi_v||^2 (1-\cos^2\alpha) = ||\Phi_u||^2 \, ||\Phi_v||^2 - (\Phi_u \cdot \Phi_v)^2 \,,$$

and therefore,

$$f(u,v) = \sigma_S(\Phi_u, \Phi_v) = \sqrt{||\Phi_u||^2 ||\Phi_v||^2 - (\Phi_u \cdot \Phi_v)^2}.$$

²See a computation in a more general case below in Example 10.6.

It is traditional to introduce the notation

$$E = ||\Phi_u||^2$$
, $F = \Phi_u \cdot \Phi_v$, $G = ||\Phi_v||^2$.

so that we get

$$\Phi^* \sigma_S = \sqrt{EG - F^2} du \wedge dv \,,$$

and hence we get for any function $f: S \to \mathbb{R}$

$$\int_{S} f dS = \int_{S} f \sigma_{S} = \int_{U} f(\Phi(u, v)) \sqrt{EG - F^{2}} du \wedge dv = \int_{U} \int_{U} f(\Phi(u, v)) \sqrt{EG - F^{2}} du dv. \quad (10.3.6)$$

Consider a special case when the surface S defined as a graph of a function ϕ over a domain $D \subset \mathbb{R}^2$. Namely, suppose

$$S = \{z = \phi(x, y), (x, y) \in D \subset \mathbb{R}^2\}.$$

The surface S as parametrized by the map

$$(x,y) \stackrel{\Phi}{\mapsto} (x,y,\phi(x,y)).$$

Then

$$E = ||\Phi_x||^2 = 1 + \phi_x^2, \ G = ||\Phi_y||^2 = 1 + \phi_y^2, \ F = \Phi_x \cdot \Phi_y = \phi_x \phi_y \,,$$

and hence

$$EG - F^2 = (1 + \phi_x^2)(1 + \phi_y^2) - \phi_x^2 \phi_y^2 = 1 + \phi_x^2 + \phi_y^2.$$

Therefore, the formula (10.3.6) takes the form

$$\int_{S} f dS = \int_{D} \int_{D} f(\Phi(x, y)) \sqrt{EG - F^{2}} dx \wedge dy =$$

$$\int_{D} \int_{D} f(x, y, \phi(x, y)) \sqrt{1 + \phi_{x}^{2} + \phi_{y}^{2}} dx dy. \tag{10.3.7}$$

Note that the formula (10.3.7) can be also deduced from (10.3.2). Indeed, the surface

$$S = \{ z = \phi(x, y), (x, y) \in D \subset \mathbb{R}^2 \},$$

can also be defined implicitly by the equation

$$F(x, y, z) = z - \phi(x, y) = 0, (x, y) \in D.$$

We have

$$\nabla F = (-\frac{\partial \phi}{\partial x}, -\frac{\partial \phi}{\partial y}, 1),$$

and, therefore,

$$\int_{S} f dS = \int_{S} \frac{f(x, y, z)}{||\nabla F||} \left(\frac{\partial F}{\partial x} dy \wedge dz + \frac{\partial F}{\partial y} dz \wedge dx + \frac{\partial F}{\partial z} dx \wedge dy \right)
= \int_{D} \frac{f(x, y, \phi(x, y))}{\sqrt{1 + \left(\frac{\partial \phi}{\partial x}\right)^{2} + \left(\frac{\partial \phi}{\partial y}\right)^{2}}} \left(-\frac{\partial \phi}{\partial x} dy \wedge d\phi - \frac{\partial \phi}{\partial y} d\phi \wedge dx + dx \wedge dy \right)
= \int_{D} \frac{f(x, y, \phi(x, y))}{\sqrt{1 + \left(\frac{\partial \phi}{\partial x}\right)^{2} + \left(\frac{\partial \phi}{\partial y}\right)^{2}}} \left(1 + \left(\frac{\partial \phi}{\partial x}\right)^{2} + \left(\frac{\partial \phi}{\partial y}\right)^{2} \right) dx dy
= \int_{D} f(x, y, \phi(x, y)) \sqrt{1 + \left(\frac{\partial \phi}{\partial x}\right)^{2} + \left(\frac{\partial \phi}{\partial y}\right)^{2}} dx dy .$$

Example 10.6. Integration over a parametrically given k-dimensional submanfold.

Consider now a more general case of a parametrically given k-submanifold A in an n-dimensional Euclidean space V. We fix a Cartersian coordinate system in V and thus identify V with \mathbb{R}^n with the standard dot-product.

Let $U \subset \mathbb{R}^k$ be a compact domain with boundary and $\phi: U \to \mathbb{R}^n$ be an embedding. We assume that the submanifold with boundary $A = \phi(U)$ is oriented by this parameterization. Let σ_A be the volume form of A. We will find an explicit expression for $\phi^*\sigma_A$. Namely, denoting coordinates in \mathbb{R}^k by (u_1, \ldots, u_k) we have $\phi^*\sigma_A = f(u)du_1 \wedge \cdots \wedge du_k$, and our goal is to compute the function f.

By definition, we have

$$f(u) = \phi^*(\sigma_A)_u(e_1, \dots, e_k) = (\sigma_A)_u(d_u\phi(e_1), \dots, d_u\phi(e_k)) = (\sigma_A)_u\left(\frac{\partial\phi}{\partial u_1}(u), \dots, \frac{\partial\phi}{\partial u_k}(u)\right) = \operatorname{Vol}_k\left(P\left(\frac{\partial\phi}{\partial u_1}(u), \dots, \frac{\partial\phi}{\partial u_k}(u)\right)\right)$$
(10.3.8)

In Section 4.2 we proved two formulas for the volume of a parallelepiped. Using formula (4.2.1) we get

$$\operatorname{Vol}_{k}\left(P\left(\frac{\partial \phi}{\partial u_{1}}(u), \dots, \frac{\partial \phi}{\partial u_{k}}(u)\right)\right) = \sqrt{\sum_{1 \leq i_{1} < \dots < i_{k} \leq n} Z_{i_{1} \dots i_{k}}^{2}}.$$
(10.3.9)

where

$$Z_{i_{1}...i_{k}} = \begin{vmatrix} \frac{\partial \phi_{i_{1}}}{\partial u_{1}}(u) & \dots & \frac{\partial \phi_{i_{1}}}{\partial u_{k}}(u) \\ \dots & \dots & \dots \\ \frac{\partial \phi_{i_{k}}}{\partial u_{1}}(u) & \dots & \frac{\partial \phi_{i_{k}}}{\partial u_{k}}(u) \end{vmatrix}.$$

$$(10.3.10)$$

Thus

$$\int_A f dV = \int_A f \sigma_A = \int_U f(\phi(u)) \sqrt{\sum_{1 \le i_1 < \dots < i_k \le n} Z_{i_1 \dots i_k}^2} du_1 \wedge \dots \wedge du_k.$$

Rewriting this formula for k = 2 we get

$$\int_{A} f dV = \int_{U} f(\phi(u)) \sqrt{\sum_{1 \le i < j \le n} \left| \frac{\partial \phi_{i}}{\partial u_{1}} \frac{\partial \phi_{i}}{\partial u_{2}} \right|^{2}} du_{1} \wedge du_{2}.$$

$$(10.3.11)$$

Alternatively we can use formula (??). Then we get

$$\operatorname{Vol}_{k}\left(P\left(\frac{\partial\phi}{\partial u_{1}}(u),\ldots,\frac{\partial\phi}{\partial u_{k}}(u)\right)\right) = \sqrt{\det\left((D\phi)^{T}D\phi\right)},\tag{10.3.12}$$

where

$$D\phi = \begin{pmatrix} \frac{\partial \phi_1}{\partial u_1} & \dots & \frac{\partial \phi_1}{\partial u_k} \\ \dots & \dots & \dots \\ \frac{\partial \phi_n}{\partial u_1} & \dots & \frac{\partial \phi_n}{\partial u_k} \end{pmatrix}$$

is the Jacobi matrix of ϕ . Thus using this expression we get

$$\int_{A} f dV = \int_{A} f \sigma_{A} = \int_{U} f(\phi(u)) \sqrt{\det\left((D\phi)^{T} D\phi\right)} dV.$$
 (10.3.13)

Exercise 10.7. In case n = 3, k = 2 show explicitly equivalence of formulas (10.3.6), (10.3.11) and (10.3.13).

Exercise 10.8. Integration over an implicitly defined k-dimensional submanfold. Suppose that $A = \{F_1 = \cdots = F_{n-k} = 0\}$ and the differentials of defining functions are linearly independent at points of A. Show that

$$\sigma_A = \frac{* (dF_1 \wedge \dots \wedge dF_{n-k})}{|| dF_1 \wedge \dots \wedge dF_{n-k} ||}.$$

The symmetric matrix $(D\phi)^T D\phi$ is the Gram matrix of vectors $\frac{\partial \phi}{\partial u_1}(u), \dots, \frac{\partial \phi}{\partial u_k}(u)$, and its entrees are pairwise scalar products of these vectors, see Remark ??.

Example 10.9. Let us compute the volume of the unit 3-sphere $S^3 = \{x_1^2 + x_2^2 + x_3^2 + x_4^2 = 1\}.$

By definition, $\operatorname{Vol}(S^3) = \int_{S^3} \mathbf{n} \, d\Omega$, where $\Omega = dx_1 \wedge dx_2 \wedge dx_3 \wedge dx_4$, and \mathbf{n} is the outward unit normal vector to the unit ball B_4 . Here S^3 should be co-oriented by the vector field \mathbf{n} . Then using Stokes' theorem we have

Introducing polar coordinates (r, ϕ) and (ρ, θ) in the coordinate planes (x_1, x_2) and (x_3, x_4) and using Fubini's theorem we get

$$\int_{B^4} = \int_0^{2\pi} \int_0^{2\pi} \int_0^1 \int_0^{\sqrt{1-r^2}} r\rho d\rho dr d\theta d\phi = 2\pi^2 \int_0^1 (r-r^3) dr = \frac{\pi^2}{2}.$$
 (10.3.15)

Hence, $Vol(S^3) = 2\pi^2$.

Exercise 10.10. Find the ratio $\frac{\operatorname{Vol}_n(B_R^n)}{\operatorname{Vol}_{n-1}(S_R^{n-1})}$.

10.4 Work and Flux

We introduce in this section two fundamental notions of vector analysis: a work of a vector field along a curve, and a flux of a vector field through a surface. Let Γ be an oriented smooth curve in a Euclidean space V and T the unit tangent vector field to Γ . Let \mathbf{v} be another vector field, defined along Γ . The function $\langle \mathbf{v}, \mathbf{T} \rangle$ equals the projection of the vector field \mathbf{v} to the tangent directions to the curve. If the vector field \mathbf{v} is viewed as a force field, then the integral $\int_{\Gamma} \langle \mathbf{v}, \mathbf{T} \rangle ds$ has the meaning of a work Work Γ (\mathbf{v}) performed by the field \mathbf{v} to transport a particle of mass 1 along the curve Γ in the direction determined by the orientation. If the curve Γ is closed then this integral is sometimes called the circulation of the vector field \mathbf{v} along Γ and denoted by $\oint_{\Gamma} \langle \mathbf{v}, \mathbf{T} \rangle ds$. As we

already indicated earlier, the sign \oint in this case has precisely the same meaning as \int , and it is used only to stress the point that we are integrating along a *closed* curve.

Consider now a co-oriented hypersurface $\Sigma \subset V$ and denote by \mathbf{n} the unit normal vector field to Σ which determines the given co-orientation of Σ . Given a vector field \mathbf{v} along Σ we will view it as the velocity vector field of a flow of a fluid in the space. Then we can interpret the integral

$$\int_{\Sigma} \langle \mathbf{v}, \mathbf{n} \rangle dV$$

as the $flux \operatorname{Flux}_{\Sigma}(\mathbf{v})$ of \mathbf{v} through Σ , i.e. the volume of fluid passing through Σ in the direction of \mathbf{n} in time 1.

Lemma 10.11. 1. For any co-oriented hypersurface Σ and a vector field \mathbf{v} given in its neighborhood we have

$$\langle \mathbf{v}, \mathbf{n} \rangle \sigma_{\Sigma} = (\mathbf{v} \, \rfloor \, \Omega)_{\Sigma},$$

where Ω is the volume form in V.

2. For any oriented curve Γ and a vector field \mathbf{v} near Γ we have

$$\langle \mathbf{v}, \mathbf{T} \rangle \sigma_{\Gamma} = \mathcal{D}(\mathbf{v})|_{\Gamma}.$$

Proof. 1. For any n-1 vectors $T_1, \ldots, T_{n-1} \in T_x \Sigma$ we have

$$\mathbf{v} \sqcup \Omega(T_1, \dots, T_{n-1}) = \Omega(\mathbf{v}, T_1, \dots, T_{n-1}) = \text{Vol}P(\mathbf{v}, T_1, \dots, T_{n-1}).$$

Using (3.3.1) we get

$$\operatorname{Vol}P(\mathbf{v}, T_1, \dots, T_{n-1}) = \langle \mathbf{v}, \mathbf{n} \rangle \operatorname{Vol}_{n-1}P(T_1, \dots, T_{n-1}) = \langle \mathbf{v}, \mathbf{n} \rangle \operatorname{Vol}P(\mathbf{n}, T_1, \dots, T_{n-1}) = \langle \mathbf{v}, \mathbf{n} \rangle \sigma_{\Sigma}(T_1, \dots, T_{n-1}).$$
(10.4.1)

2. The tangent space $T_x\Gamma$ is generated by the vector \mathbf{T} , and hence we just need to check that $\langle \mathbf{v}, \mathbf{T} \rangle \sigma_{\Gamma}(\mathbf{T}) = \mathcal{D}(\mathbf{v})(\mathbf{T})$. But $\sigma_{\Gamma}(\mathbf{T}) = \text{Vol}(\mathbf{n}, \mathbf{T}) = 1$, and hence

$$\mathcal{D}(\mathbf{v})(\mathbf{T}) = \langle \mathbf{v}, \mathbf{T} \rangle = \langle \mathbf{v}, \mathbf{T} \rangle \sigma_{\Gamma}(\mathbf{T}).$$

Note that if we are given a Cartesian coordinate system in V and $\mathbf{v} = \sum_{1}^{n} a_{j} \frac{\partial}{\partial x_{j}}$, then

$$\mathbf{v} \, \rfloor \, \Omega = \sum_{1}^{n-1} (-1)^{i-1} a_i dx_1 \stackrel{i}{\overset{i}{\vee}} \, dx_n, \quad \mathcal{D}(\mathbf{v}) = \sum_{1}^{n} a_i dx_i \, .$$

Thus, we have

Corollary 10.12.

$$\operatorname{Flux}_{\Sigma}(\mathbf{v}) = \int_{\Sigma} \langle \mathbf{v}, \mathbf{n} \rangle dV = \int_{\Sigma} (\mathbf{v} \, \mathbf{n}) = \int_{\Sigma} \sum_{1}^{n} (-1)^{i-1} a_{i} dx_{1} \stackrel{i}{\overset{i}{\vee}} dx_{n};$$

$$\operatorname{Work}_{\Gamma}(\mathbf{v}) = \int_{\Sigma} \langle \mathbf{v}, \mathbf{T} \rangle = \int_{\Sigma} \mathcal{D}(\mathbf{v}) = \int_{\Sigma} \sum_{1}^{n} a_{i} dx_{i}.$$

In particular if n = 3 we have

$$\operatorname{Flux}_{\Sigma}(\mathbf{v}) = \int_{\Sigma} a_1 dx_2 \wedge dx_3 + a_2 dx_3 \wedge dx_1 + a_3 dx_1 \wedge dx_2.$$

Let us also recall that in a Euclidean space V we have $v \, \lrcorner \, \Omega = *\mathcal{D}(v)$. Hence, the equation $\omega = v \, \lrcorner \, \Omega$ is equivalent to the equation

$$v = \mathcal{D}^1 (*^{-1}\omega) = (-1)^{n-1} \mathcal{D}^{-1} (*\omega).$$

In particular, when n=3 we get $\mathbf{v}=\mathcal{D}^{-1}(*\omega)$. Thus we get

Corollary 10.13. For any differential (n-1)-form ω and an oriented compact hypersurface Σ we have

$$\int_{\Sigma} \omega = \operatorname{Flux}_{\Sigma} \mathbf{v},$$

where $\mathbf{v} = (-1)^{n-1} \mathcal{D}^{-1} (*\omega)$.

Integration of functions along curves and surfaces can be interpreted as the work and the flux of appropriate vector fields. Indeed, suppose we need to compute an integral $\int_{\Gamma} f ds$. Consider the tangent vector field $\mathbf{v}(x) = f(x)\mathbf{T}(x)$, $x \in \Gamma$, along Γ . Then $\langle \mathbf{v}, \mathbf{T} \rangle = f$ and hence the integral $\int_{\Gamma} f ds$ can be interpreted as the work Work Γ (\mathbf{v}). Therefore, we have

$$\int_{\Gamma} f ds = \operatorname{Work}_{\Gamma}(\mathbf{v}) = \int_{\Gamma} \mathcal{D}(\mathbf{v})$$

Note that we can also express \mathbf{v} through ω by the formula $\mathbf{v} = \mathcal{D}^{-1} \star \omega$, see Section 8.7.

Similarly, to compute an integral $\int_{\Sigma} f dS$ let us co-orient the surface Σ with a unit normal to Σ vector field $\mathbf{n}(x)$, $x \in \Sigma$ and set $\mathbf{v}(x) = f(x) \mathbf{n}(x)$. Then $\langle \mathbf{v}, \mathbf{n} \rangle = f$, and hence

$$\int_{\Gamma} f dS = \int_{\Gamma} \langle \mathbf{v}, \mathbf{n} \rangle dS = \operatorname{Flux}_{\Sigma}(\mathbf{v}) = \int_{\Gamma} \omega,$$

where $\omega = \mathbf{v} \, \mathsf{J} \, \Omega$.

10.5 Integral formulas of vector analysis

Using Corollary 10.12 we can reformulate Stokes' theorem for domains in a \mathbb{R}^3 as follows.

Theorem 10.14. Let \mathbf{v} be a smooth vector field in a domain $U \subset \mathbb{R}^3$ with a smooth (or piece-wise) smooth boundary Σ . Suppose that Σ is co-oriented by an outward normal vector field. Then we have

$$\operatorname{Flux}_{\Sigma} \mathbf{v} = \int \int_{U} \int \operatorname{div} \mathbf{v} dx dy dz.$$

Indeed, div $\mathbf{v} = *d\omega$. Hence we have

$$\int_{U} \operatorname{div} \mathbf{v} dV = \int_{U} (\star d\omega) dx \wedge dy \wedge dz = \int_{U} d\omega = \int_{\Sigma} \omega = \int_{\Sigma} \mathbf{v} \, \exists \, \Omega = \operatorname{Flux}_{\Sigma} \mathbf{v} \, .$$

This theorem clarifies the meaning of $\operatorname{div} \mathbf{v}$:

Let $B_r(x)$ be the ball of radius r centered at a point $x \in \mathbb{R}^3$, and $S_r(x) = \partial B_r(x)$ be its boundary sphere co-oriented by the outward normal vector field. Then

$$\operatorname{div} \mathbf{v}(x) = \lim_{r \to 0} \frac{\operatorname{Flux}_{S_r(x)} \mathbf{v}}{\operatorname{Vol}(B_r(x))}.$$

Theorem 10.15. Let Σ be a piece-wise smooth compact oriented surface in \mathbb{R}^3 with a piece-wise smooth boundary $\Gamma = \partial \Sigma$ oriented respectively. Let \mathbf{v} be a smooth vector field defined near Σ . Then

$$\operatorname{Flux}_{\Sigma}(\operatorname{curl} \mathbf{v}) = \int\limits_{\Sigma} \langle \operatorname{curl} \mathbf{v}, \, \mathbf{n} \rangle dV = \oint\limits_{\Gamma} \, \mathbf{v} \cdot \mathbf{T} ds = \operatorname{Work}_{\Gamma} \mathbf{v} \,.$$

To prove the theorem we again use Stokes' theorem and the connection between integrals of functions and differential forms. Set $\alpha = \mathcal{D}(\mathbf{v})$. Then $\operatorname{curl} \mathbf{v} = \mathcal{D}^{-1} \star d\alpha$. We have

$$\oint_{\Gamma} \mathbf{v} \cdot \mathbf{T} ds = \int_{\Gamma} \alpha = \int_{\Sigma} d\alpha = \operatorname{Flux}_{\Sigma}(\mathcal{D}^{-1} \star (d\alpha)) = \operatorname{Flux}_{\Sigma}(\operatorname{curl} \mathbf{v}).$$

Again, similar to the previous case, this theorem clarifies the meaning of curl. Indeed, let us denote by $D_r(x, \mathbf{w})$ the 2-dimensional disc of radius r in \mathbb{R}^3 centered at a point $x \in \mathbb{R}^3$ and orthogonal to a unit vector $\mathbf{w} \in \mathbb{R}^3_x$. Set

$$c(x, \mathbf{w}) = \lim_{r \to 0} \frac{\operatorname{Work}_{\partial D_r(x, \mathbf{w})} \mathbf{v}}{\pi r^2}.$$

Then

$$c(x, \mathbf{w}) = \lim_{r \to 0} \frac{\operatorname{Flux}_{D_r(x, \mathbf{w})}(\operatorname{curl} \mathbf{v})}{\pi r^2} = \lim_{r \to 0} \frac{\int\limits_{D_r(x, \mathbf{w})} \langle \operatorname{curl} \mathbf{v}, \mathbf{w} \rangle}{\pi r^2} = \langle \operatorname{curl} \mathbf{v}, \mathbf{w} \rangle.$$

Hence, $||\operatorname{curl} \mathbf{v}(x)|| = \max_{\mathbf{w}} c(x, \mathbf{w})$ and direction of $\operatorname{curl} \mathbf{v}(x)$ coincides with the direction of the vector \mathbf{w} for which the maximum value of $c(x, \mathbf{w})$ is achieved.

10.6 Expressing div and curl in curvilinear coordinates

Let us show how to compute div \mathbf{v} and curl \mathbf{v} of a vector field \mathbf{v} in \mathbb{R}^3 given in a curvilinear coordinates u_1, u_2, u_3 , i.e. expressed through the coordinate vector fields $\frac{\partial}{\partial u_1}, \frac{\partial}{\partial u_2}$ and $\frac{\partial}{\partial u_3}$. Let

$$\Omega = f(u_1, u_2, u_3) du_1 \wedge du_2 \wedge du_3$$

be the volume form $dx_1 \wedge dx_2 \wedge dx_3$ expressed in coordinates u_1, u_2, u_3 .

Let us first compute $\star (du_1 \wedge du_2 \wedge du_3)$. We have

$$\star (du_1 \wedge du_2 \wedge du_3) = \star \left(\frac{1}{f} dx_1 \wedge dx_2 \wedge dx_3\right) = \frac{1}{f}.$$

Let

$$\mathbf{v} = a_1 \frac{\partial}{\partial u_1} + a_2 \frac{\partial}{\partial u_2} + a_3 \frac{\partial}{\partial u_3}.$$

Then we have

$$\operatorname{div} \mathbf{v} = \star d(\mathbf{v} \sqcup \Omega)$$

$$= \star d\left(\left(\sum_{1}^{3} a_{i} \frac{\partial}{\partial u_{i}}\right) \sqcup f d u_{1} \wedge d u_{2} \wedge d u_{3}\right)$$

$$= \star d\left(f a_{1} d u_{2} \wedge d u_{3} + f a_{2} d u_{3} \wedge d u_{1} + f a_{3} d u_{1} \wedge d u_{2}\right)$$

$$= \star \left(\left(\frac{\partial (f a_{1})}{\partial u_{1}} + \frac{\partial (f a_{2})}{\partial u_{2}} + \frac{\partial (f a_{3})}{\partial u_{3}}\right) d u_{1} \wedge d u_{2} \wedge d u_{3}\right)$$

$$= \frac{1}{f} \left(\frac{\partial (f a_{1})}{\partial u_{1}} + \frac{\partial (f a_{2})}{\partial u_{2}} + \frac{\partial (f a_{3})}{\partial u_{3}}\right)$$

$$= \frac{\partial a_{1}}{\partial u_{1}} + \frac{\partial a_{2}}{\partial u_{2}} + \frac{\partial a_{3}}{\partial u_{3}} + \frac{1}{f} \left(\frac{\partial f}{\partial u_{1}} a_{1} + \frac{\partial f}{\partial u_{2}} a_{2} + \frac{\partial f}{\partial u_{3}} a_{3}\right).$$

In particular, we see that the divergence of a vector field is expressed by the same formulas as in the cartesian case if and only if the volume form is proportional to the form $du_1 \wedge du_2 \wedge du_3$ with a constant coefficient.

For instance, in the spherical coordinates the volume form can be written as

$$\Omega = r^2 \sin \varphi dr \wedge d\varphi \wedge d\theta,^4$$

and hence the divergence of a vector field

$$\mathbf{v} = a\frac{\partial}{\partial r} + b\frac{\partial}{\partial \theta} + c\frac{\partial}{\partial \varphi}$$

can be computed by the formula

$$\operatorname{div} \mathbf{v} = \frac{\partial a}{\partial r} + \frac{\partial b}{\partial \theta} + \frac{\partial c}{\partial \varphi} + \frac{2a}{r} + c \cot \varphi.$$

The general formula for $\operatorname{curl} \mathbf{v}$ in curvilinear coordinates looks pretty complicated. So instead of deriving the formula we will explain here how it can be obtained in the general case, and then illustrate this procedure for the spherical coordinates.

By the definition we have

⁴Note that the spherical coordinates ordered as (r, ϕ, θ) determine the same orientation of \mathbb{R}^3 as the cartesian coordinates (x, y, z).

$$\operatorname{curl} \mathbf{v} = D^{-1} \left(\star \left(d \left(D(\mathbf{v}) \right) \right) \right) .$$

Hence we first need to compute $D(\mathbf{v})$.

To do this we need to introduce a symmetric matrix

$$G = \begin{pmatrix} g_{11} & g_{12} & g_{13} \\ g_{21} & g_{22} & g_{23} \\ g_{31} & g_{32} & g_{33} \end{pmatrix},$$

where

$$g_{ij} = \langle \frac{\partial}{\partial u_i}, \frac{\partial}{\partial u_j} \rangle, \ i, j = 1, 2, 3.$$

The matrix G is called the $Gram\ matrix$.

Notice, that if $D(\mathbf{v}) = A_1 du_1 + B du_2 + C du_3$ then for any vector $h = h_1 \frac{\partial}{\partial u_1} + h_2 \frac{\partial}{\partial u_2} + h_3 \frac{\partial}{\partial u_3}$ we have

$$D(\mathbf{v})(h) = A_1 h_1 + A_2 h_2 + A_3 h_3 = \begin{pmatrix} A_1 & A_2 & A_3 \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = \langle \mathbf{v}, h \rangle.$$

But

$$\langle \mathbf{v}, h \rangle = \begin{pmatrix} a_1 & a_2 & a_3 \end{pmatrix} G \begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix}.$$

Hence

$$\begin{pmatrix} A_1 & A_2 & A_3 \end{pmatrix} = \begin{pmatrix} a_1 & a_2 & a_3 \end{pmatrix} G,$$

or, equivalently, because the Gram matrix G is symmetric we can write

$$\begin{pmatrix} A_1 \\ A_2 \\ A_3 \end{pmatrix} = G \begin{pmatrix} a_1 \\ a_2 \\ a_3 \end{pmatrix},$$

and therefore,

$$A_i = \sum_{i=1}^n g_{ij} a_j, \quad i = 1, 2, 3.$$

After computing

$$\omega = d(D(\mathbf{v})) = B_1 du_2 \wedge du_3 + B_2 du_3 \wedge du_1 + B_3 du_1 \wedge du_2$$

we compute curl \mathbf{v} by the formula = curl $\mathbf{v} = D^{-1}$ ($\star \omega$). Let us recall (see Proposition 4.15 above) that for any vector field \mathbf{w} the equality $D\mathbf{w} = \star \omega$ is equivalent to the equality $\mathbf{w} \, \exists \, \Omega = \omega$, where $\Omega = f du_1 \wedge du_2 \wedge du_3$ is the volume form. Hence, if

$$\operatorname{curl} \mathbf{v} = c_1 \frac{\partial}{\partial u_1} + c_2 \frac{\partial}{\partial u_2} + c_3 \frac{\partial}{\partial u_3}$$

then we have

$$\mathbf{w} \, \rfloor \, \Omega = f c_1 du_2 \wedge du_3 + f c_2 du_3 \wedge du_1 + f c_3 du_1 \wedge du_2,$$

and therefore,

$$\operatorname{curl} \mathbf{v} = \frac{B_1}{f} \frac{\partial}{\partial u_1} + \frac{B_2}{f} \frac{\partial}{\partial u_2} + \frac{B_3}{f} \frac{\partial}{\partial u_3}.$$

Let us use the above procedure to compute $\operatorname{curl} \mathbf{v}$ of the vector field

$$\mathbf{v} = a\frac{\partial}{\partial r} + b\frac{\partial}{\partial \phi} + c\frac{\partial}{\partial \theta}$$

given in the spherical coordinates. The Gram matrix in this case is the diagonal matrix

$$G = \begin{pmatrix} 1 & 0 & 0 \\ 0 & r^2 & 0 \\ 0 & 0 & r^2 \sin^2 \varphi \end{pmatrix} .$$

Hence,

$$D\mathbf{v} = adr + br^2d\varphi + cr^2\sin^2\varphi d\theta,$$

and

$$\begin{split} \omega &= d(D\,\mathbf{v}) &= da \wedge dr + d(br^2) \wedge d\varphi + d(cr^2\sin^2\varphi) \wedge d\theta \\ &= \left(-r\frac{\partial b}{\partial \theta} + r^2c\sin 2\varphi + r^2\sin^2\varphi\frac{\partial c}{\partial \varphi}\right)d\varphi \wedge d\theta \\ &+ \left(-\frac{\partial a}{\partial \varphi} + r^2\frac{\partial b}{\partial r} + 2br\right)dr \wedge d\varphi + \left(-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}\right)d\theta \wedge dr \,. \end{split}$$

Finally, we get the following expression for $\operatorname{curl} \mathbf{v}$:

$$\operatorname{curl} \mathbf{v} = \frac{-r\frac{\partial b}{\partial \theta} + r^2c\sin 2\varphi + r^2\sin^2\varphi\frac{\partial c}{\partial \varphi}}{r^2\cos\varphi}\frac{\partial}{\partial r} + \frac{-\frac{\partial a}{\partial \varphi} + r^2\frac{\partial b}{\partial r} + 2br}{r^2\cos\varphi}\frac{\partial}{\partial \varphi} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} \cdot \frac{\partial c}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2rc\sin^2\varphi - r^2\sin^2\varphi\frac{\partial c}{\partial r} + \frac{\partial a}{\partial \theta}}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \phi} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \theta} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \phi} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \phi}\frac{\partial}{\partial \phi} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \phi}\frac{\partial}{\partial \phi} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \phi}\frac{\partial}{\partial \phi} + \frac{-2r\cos\varphi}{r^2\cos\varphi}\frac{\partial}{\partial \phi}\frac{\partial}{\partial \phi}$$

Chapter 11

Applications of Stokes' formula

11.1 Integration of closed and exact forms

Let us recall that a differential k-form ω is called *closed* if $d\omega = 0$, and that it is called *exact* if there exists a (k-1)-form α , called *primitive* of ω , such that $\omega = d\alpha$.

Any exact form is closed, because $d(d\alpha) = 0$. Any *n*-form in a *n*-dimensional space is closed.

Proposition 11.1. a) For a closed k-form ω defined near a (k+1)-dimensional submanifold Σ with boundary $\partial \Sigma$ we have

$$\int_{\partial \Sigma} \omega = 0.$$

b) If ω is exact k-form defined near a closed k-dimensional submanifold S then

$$\int_{S} \omega = 0.$$

The proof immediately follows from Stokes' formula. Indeed, in case a) we have

$$\int_{\partial \Sigma} \omega = \int_{\Sigma} d\omega = 0.$$

In case b) we have $\omega = d\alpha$ and $\partial S = \emptyset$. Thus

$$\int_{S} d\alpha = \int_{\emptyset} \alpha = 0.$$

Proposition 11.1b) gives a necessary condition for a closed form to be exact.

Example 11.2. The differential 1-form $\alpha = \frac{1}{x^2+y^2}(xdy-ydx)$ defined on the punctured plane $\mathbb{R}^2 \setminus 0$ is closed but not exact.

Indeed, it is straightforward to check that α is exact (one can simplify computations by passing to polar coordinates and computing that $\alpha = d\varphi$). To check that it is not exact we compute the integral $\int_S \alpha$, where S in the unit circle $\{x^2 + y^2 = 1\}$. We have

$$\int_{S} \alpha = \int_{0}^{2\pi} d\varphi = 2\pi \neq 0.$$

More generally, an (n-1)-form

$$\theta_n = \sum_{i=1}^n (-1)^{i-1} \frac{x_i}{r^n} dx_1 \wedge \dots \wedge dx_n \quad (dx_i \text{ is missing})$$

is closed in $\mathbb{R}^n \setminus 0$. However, it is not exact. Indeed, let us show that $\int_{S^{n-1}} \theta_n \neq 0$, where S^{n-1} is the unit sphere oriented as the boundary of the unit ball. Let us recall that the volume form $\sigma_{S^{n-1}}$ on the unit sphere is defined as

$$\sigma_{S^{n-1}} = \mathbf{n} \, \rfloor \, \Omega = \sum_{i=1}^n (-1)^{i-1} \frac{x_i}{r} dx_1 \wedge \dot{\tilde{x}} \cdot \wedge dx_n.$$

Notice that $\theta_n|_{S^{n-1}} = \sigma_{S^{n-1}}$, and hence

$$\int_{S^{n-1}} \theta_n = \int_{S^{n-1}} \sigma_{S^{n-1}} = \int_{S^{n-1}} dV = \text{Vol}(S^{n-1}) > 0.$$

11.2 Approximation of continuous functions by smooth ones

Theorem 11.3. Let $C \subset V$ be a compact domain with smooth boundary. Then any continuous function $f: C \to \mathbb{R}$ can be C^0 -approximated by C^∞ - smooth functions, i.e. for any $\epsilon > 0$ there exists a C^∞ -smooth function $g: C \to \mathbb{R}$ such that $|f(x) - g(x)| < \epsilon$ for any $x \in C$. Moreover, if the function f is already C^∞ -smooth in a neighborhood of a closed subset $B \subset \operatorname{Int} C$, then one can arrange that the function g coincides with f over B.

Lemma 11.4. There is a continuous extension of f to V.

Sketch of the proof. Let **n** be the outward normal vector field to the boundary ∂C . If the boundary is C^{∞} -smooth then so is the vector field **n**. Consider a map $\nu: \partial C \times [0,1] \to V$ given by the formula $\nu(x,t) = x + t$ **n**, $x \in \partial C$, $t \in [0,1]$. The differential of ν at the points of $\partial C \times 0$ has rank n. (Exercise: prove this.) Hence by the inverse function theorem for a sufficiently small $\epsilon > 0$ the map ν is a diffeomorphism of $\partial C \times [0,\epsilon)$ onto $U \setminus \text{Int } C$ for some open neighborhood $U \supset C$. Consider a function $F: \partial C \times [0,\epsilon) \to \mathbb{R}$, defined by the formula

$$F(x,t) = \left(1 - \frac{2t}{\epsilon}\right)f(x)$$

if $t \in [0, \frac{\epsilon}{2}]$ and f(x, t) = 0 if $t \in (\frac{\epsilon}{2}, \epsilon)$. Now we can extend f to U by the formula $f(y) = F\nu^{-1}(y)$ if $y \in U \setminus C$, and setting it to 0 outside U.

Consider the function

$$\Psi_{\epsilon} = \frac{1}{\int\limits_{D_{\epsilon}(0)} \psi_{0,\epsilon} dV} \psi_{0,\epsilon},$$

where $\psi_{0,\sigma}$ is a bump function defined above in (9.1.2). It is supported in the disc $D_{\sigma}(0)$, non-negative, and satisfies

$$\int_{D_{\sigma}(0)} \Psi_{\sigma} dV = 1.$$

. Given a continuous function $f:V\to\mathbb{R}$ we define a function $f_\sigma:V\to\mathbb{R}$ by the formula

$$f_{\sigma}(x) = \int f(x-y)\Psi_{\sigma}(y)d^{n}y. \tag{11.2.1}$$

Then

Lemma 11.5. 1. The function f_{σ} is C^{∞} -smooth.

2. For any $\epsilon > 0$ there exists $\delta > 0$ such that for all $x \in C$ we have $|f(x) - f_{\sigma}(x)| < \epsilon$ provided that $\sigma < \delta$.

Proof. 1. By the change of variable formula we have, replacing the variable y by u = y - x:

$$f_{\sigma}(x) = \int_{D_{\epsilon}(0)} f(x-y)\Psi_{\sigma}(y)d^{n}y = \int_{D_{\epsilon}(-x)} f(-u)\Psi_{\sigma}(x+u)d^{n}u = \int_{V} f(-u)\Psi_{\sigma}(x+u)d^{n}u.$$

But the expression under the latter integral depends on x C^{∞} -smoothly as a parameter. Hence, by the theorem about differentiating integral over a parameter, we conclude that the function f_{ϵ} in C^{∞} -smooth.

2. Fix some $\sigma_0 > 0$. The function f is uniformly continuous in $\overline{U_{\sigma_0}(C)}$. Hence there exists $\delta > 0$ such that $x, x' \in \overline{U_{\sigma_0}(C)}$ and $||x - x'|| < \delta$ we have $|f(x) - f(x')| < \epsilon$. Hence, for $\sigma < \min(\sigma_0, \delta)$ and for $x \in C$ we have

$$|f_{\sigma}(x) - f(x)| = |\int_{D_{\epsilon}(0)} f(x - y)\Psi_{\sigma}(y)d^{n}y - \int_{D_{\epsilon}(0)} f(x)\Psi_{\sigma}(y)d^{n}y| \leq \int_{D_{\epsilon}(0)} |f(x - y) - f(x)|\Psi_{\sigma}(y)d^{n}y \leq \epsilon \int_{D_{\epsilon}(0)} \Psi_{\sigma}(y)d^{n}y = \epsilon.$$

$$(11.2.2)$$

Proof of Theorem 11.3. Lemma 11.5 implies that for a sufficiently small σ the function $g = f_{\sigma}$ is the required C^{∞} -smooth ϵ -approximation of the continuous function f. To prove the second part of the theorem let us assume that f is already C^{∞} -smooth on a neighborhood U, $B \subset U \subset C$. Let us choose a cut-off function $\sigma_{B,U}$ constructed in Lemma 9.2 and define the required approximation g by the formula $f_{\sigma} + (f - f_{\sigma})\sigma_{B,U}$.

Theorem 11.3 implies a similar theorem form continuous maps $C \to \mathbb{R}^n$ by applying it to all coordinate functions.

11.3 Homotopy

Let A, B be any 2 subsets of vector spaces V and W, respectively. Two continuous maps $f_0, f_1 : A \to B$ are called *homotopic* if there exists a continuous map $F : A \times [0,1] \to B$ such that $F(x,0) = f_0(x)$ and $F(x,1) = f_1(x)$ for all $t \in [0,1]$. Notice that the family $f_t : A \to B$, $t \in [0,1]$, defined by the formula $f_t(x) = F(x,t)$ is a continuous deformation connecting f_0 and f_1 . Conversely, any such continuous deformation $\{f_t\}_{t \in [0,1]}$ provides a homotopy between f_0 and f_1 .

Given a subset $C \subset A$, we say that a homotopy $\{f_t\}_{t \in [0,1]}$ is fixed over C if $f_t(x) = f_0(x)$ for all $x \in C$ and all $t \in [0,1]$.

A set A is called *contractible* if there exists a point $a \in A$ and a homotopy $f_t : A \to A$, $t \in [0, 1]$, such that $f_1 = \text{Id}$ and f_0 is a constant map, i.e. $f_1(x) = x$ for all $x \in A$ and $f_0(x) = a \in A$ for all $x \in A$.

Example 11.6. Any star-shaped domain A in V is contractible. Indeed, assuming that it is star-shaped with respect to the origin, the required homotopy $f_t: A \to A$, $t \in [0,1]$, can be defined by the formula $f_t(x) = tx$, $x \in A$.

Remark 11.7. In what follows we will always assume all homotopies to be smooth. According to Theorem 11.3 this is not a serious constraint. Indeed, any continuous map can be C^0 -approximated by smooth ones, and any homotopy between smooth maps can be C^0 -approximated by a smooth homotopy between the same maps.

Lemma 11.8. Let $U \subset V$ be an open set, A a compact oriented manifold (possibly with boundary) and α a smooth closed differential k-form on U. Let $f_0, f_1 : A \to U$ be two maps which are homotopic relative to the boundary ∂A . Then

$$\int_A f_0^* \alpha = \int_A f_1^* \alpha.$$

Proof. Let $F: A \times [0,1] \to U$ be the homotopy map between f_0 and f_1 . By assumption $d\alpha = 0$, and hence $\int_{A \times [0,1]} F^* d\alpha = 0$. Then, using Stokes' theorem we have

$$0 = \int\limits_{A \times [0,1]} F^* d\alpha = \int\limits_{A \times [0,1]} dF^* \alpha = \int\limits_{\partial (A \times [0,1])} F^* \alpha = \int\limits_{\partial A \times [0,1]} F^* \alpha + \int\limits_{A \times 1} F^* \alpha + \int\limits_{A \times 0} F^* \alpha$$

where the boundary $\partial(A \times [0,1]) = (A \times 1) \cap (A \times 0) \cap (\partial A \times [0,1])$ is oriented by an outward normal vector field \mathbf{n} . Note that $\mathbf{n} = \frac{\partial}{\partial t}$ on $A \times 1$ and $\mathbf{n} = -\frac{\partial}{\partial t}$ on $A \times 0$, where we denote by t the coordinate corresponding to the factor [0,1]. First, we notice that $F^*\alpha|_{\partial A \times [0,1]} = 0$ because the map F is independent of the coordinate t, when restricted to $\partial A \times [0,1]$. Hence $\int\limits_{\partial A \times [0,1]} F^*\alpha = 0$. Consider the inclusion maps $A \to A \times [0,1]$ defined by the formulas $j_0(x) = (x,0)$ and $j_1(x) = (x,1)$. Note that j_0, j_1 are diffeomorphisms $A \to A \times 0$ and $A \to A \times 1$, respectively. Note that the map j_1 preserves the orientation while j_0 reverses it. We also have $F \circ j_0 = f_0$, $F \circ j_1 = f_1$. Hence, $\int\limits_{A \times 1} F^*\alpha = \int\limits_A f_1^*\alpha$ and $\int\limits_{A \times 0} F^*\alpha = -\int\limits_A f_0^*\alpha$. Thus,

$$0 = \int\limits_{\partial (A \times [0,1])} F^*\alpha = \int\limits_{\partial A \times [0,1]} F^*\alpha + \int\limits_{A \times 1} F^*\alpha + \int\limits_{A \times 0} F^*\alpha = \int\limits_A f_1^*\alpha - \int\limits_A f_0^*\alpha.$$

Lemma 11.9. Let A be an oriented m-dimensional manifold, possibly with boundary. Let $\Omega(A)$ denote the space of differential forms on A and $\Omega(A \times [0,1])$ denote the space of differential forms on the product $A \times [0,1]$. Let $j_0, j_1 : A \to A \times [0,1]$ be the inclusion maps $j_0(x) = (x,0) \in A \times [0,1]$ and $j_1(x) = (x,1) \in A \times [0,1]$. Then there exists a linear map $K : \Omega(A \times [0,1]) \to \Omega(A)$ such that

- If α is a k-form, k = 1, ..., m then $K(\alpha)$ is a (k-1)-form;
- $d \circ K + K \circ d = j_1^* j_0^*$, i.e. for each differential k-form $\alpha \in \Omega^k(A \times [0,1])$ one has $dK(\alpha) + K(d\alpha) = j_1^*\alpha j_0^*\alpha$.

Remark 11.10. Note that the first d in the above formula denotes the exterior differential $\Omega^k(A) \to \Omega^k(A)$, while the second one is the exterior differential $\Omega^k(A \times [0,1]) \to \Omega^k(A \times [0,1])$.

Proof. Let us write a point in $A \times [0,1]$ as (x,t), $x \in A$, $t \in [0,1]$. To construct $K(\alpha)$ for a given $\alpha \in \Omega^k(A \times [0,1])$ we first contract α with the vector field $\frac{\partial}{\partial t}$ and then integrate the resultant form with respect to the t-coordinate. More precisely, note that any k-form α on $A \times [0,1]$ can be written as $\alpha = \beta(t) + dt \wedge \gamma(t)$, $t \in [0,1]$, where for each $t \in [0,1]$

$$\beta(t) \in \Omega^k(A), \gamma(t) \in \Omega^{k-1}(A).$$

Then $\frac{\partial}{\partial t} \, \exists \, \alpha = \gamma(t)$ and we define $K(\alpha) = \int_{0}^{1} \gamma(t) dt$.

If we choose a local coordinate system (u_1, \ldots, u_m) on A then $\gamma(t)$ can be written as $\gamma(t) = \sum_{1 \leq i_1 < \cdots < i_{\leq m}} h_{i_1 \dots i_k}(t) du_{i_1} \wedge \cdots \wedge du_{i_k}$, and hence

$$K(\alpha) = \int_{0}^{1} \gamma(t)dt = \sum_{1 \le i_1 < \dots < i \le m} \left(\int_{0}^{1} h_{i_1 \dots i_k}(t)dt \right) du_{i_1} \wedge \dots \wedge du_{i_k}.$$

Clearly, K is a linear operator $\Omega^k(A \times I) \to \Omega^{k-1}(A)$.

Note that if $\alpha = \beta(t) + dt \wedge \gamma(t) \in \Omega^k(A \times [0,1])$ then

$$j_0^* \alpha = \beta(0), \ j_1^* \alpha = \beta(1).$$

We further have

$$K(\alpha) = \int_{0}^{1} \gamma(t)dt;$$

$$d\alpha = d_{A \times I}\alpha = d_U\beta(t) + dt \wedge \dot{\beta}(t) - dt \wedge d_U\gamma(t) = d_A\beta(t) + dt \wedge (\dot{\beta}(t) - d_A\gamma(t)),$$

where we denoted $\dot{\beta}(t) := \frac{\partial \beta(t)}{\partial t}$ and I = [0, 1]. Here the notation $d_{A \times I}$ stands for exterior differential on $\Omega(A \times I)$ and d_A denotes the exterior differential on $\Omega(A)$. In other words, when we write $d_A\beta(t)$ we view $\beta(t)$ as a form on A depending on t as a parameter. We do not write any subscript for d when there could not be any misunderstanding.

Hence,

$$K(d\alpha) = \int_{0}^{1} (\dot{\beta}(t) - d_A \gamma(t)) dt = \beta(1) - \beta(0) - \int_{0}^{1} d_A \gamma(t) dt;$$
$$d(K(\alpha)) = \int_{0}^{1} d_A \gamma(t) dt.$$

Therefore,

$$d(K(\alpha)) + K(d(\alpha)) = \beta(1) - \beta(0) - \int_{0}^{1} d_{A}\gamma(t)dt + \int_{0}^{1} d_{A}\gamma(t)dt$$
$$= \beta(1) - \beta(0) = j_{1}^{*}(\alpha) - j_{0}^{*}(\alpha).$$

Theorem 11.11. (Poincaré's Lemma) Let U be a contractible domain in V. Then any closed form in U is exact. More precisely, let $F: U \times [0,1] \to U$ be the contraction homotopy to a point $a \in U$, i.e. F(x,1) = x, F(x,0) = a for all $x \in U$. Then if ω a closed k-form in U then

$$\omega = dK(F^*\omega),$$

where $K: \Omega^{k+1}(U \times [0,1]) \to \Omega^k(U)$ is an operator constructed in Lemma 11.9.

Proof. Consider a contraction homotopy $F: U \times [0,1] \to U$. Then $F \circ j_0(x) = a \in U$ and $F \circ j_1(x) = x$ for all $x \in U$. Consider an operator $K: \Omega U \to \Omega(U)$ constructed above. Thus

$$K \circ d + d \circ K = j_1^* - j_0^*.$$

Let ω be a closed k-form on U. Denote $\alpha := F^*\omega$. Thus α is a k-form on $U \times [0,1]$. Note that $d\alpha = dF^*\omega = F^*d\omega = 0$, $j_1^*\alpha = (F \circ j_1)^*\omega = \omega$ and $j_0^*\alpha = (F \circ j_0)^*\omega = 0$. Then, using Lemma 11.9 we have

$$K(d\alpha) + dK(\alpha) = dK(\alpha) = j_1^* \alpha - j_0^* \alpha = \omega, \tag{11.3.1}$$

i.e. $\omega = dK(F^*\omega)$.

In particular, any closed form is locally exact.

Example 11.12. Let us work out explicitly the formula for a primitive of a closed 1-form in a star-shaped domain $U \subset \mathbb{R}^n$. We can assume that U is star-shaped with respect to the origin. Let $\alpha = \sum_{i=1}^{n} f_i dx_i$ be a closed 1-form. Then according to Theorem 11.11 we have $\alpha = dF$, where $F = K(\Phi^*\alpha)$, where $\Phi : U \times [0,1] \to U$ is a contraction homotopy, i.e. $\Phi(x,1) = x, \Phi(x,1) = 0$ for $x \in U$. Φ can be defined by the formula $\Phi(x,t) = tx, x \in U, t \in [0,1]$. Then

$$\Phi^*\alpha = \sum_{1}^{n} f_i(tx)d(tx_i) = \sum_{1}^{n} tf_i(tx)dx_i + \sum_{1}^{n} x_i f_i(tx)dt.$$

Hence,

$$K(\alpha) = \int_{0}^{1} \frac{\partial}{\partial t} \, d\Phi^* \alpha = \int_{0}^{1} \left(\sum_{1}^{n} x_i f_i(tx) \right) dt.$$

Note that this expression coincides with the expression in formula (9.4.1) in Section 9.4.

Exercise 11.13. Work out an explicit expression for a primitive of a closed 2-form $\alpha = Pdy \wedge dz + Qdz \wedge dx + Rdx \wedge dy$ on a star-shaped domain $U \subset \mathbb{R}^3$.

Example 11.14. 1. $\mathbb{R}^n_+ = \{x_1 \geq 0\} \subset \mathbb{R}^n$ is not diffeomorphic to \mathbb{R}^n . Indeed, suppose there exists such a diffeomorphism $f: \mathbb{R}^n_+ \to \mathbb{R}^n$. Denote a := f(0). Without loss of generality we can assume that a = 0. Then $\widetilde{f} = f|_{\mathbb{R}^n_+ \setminus 0}$ is a diffeomorphism $\mathbb{R}^n_+ \setminus 0 \to \mathbb{R}^n \setminus 0$. But $\mathbb{R}^n_+ \setminus 0$ is star-shaped with respect to any point with positive coordinate x_1 , and hence it is contractible. In particular any closed form on $\mathbb{R}^n_+ \setminus 0$ is exact. On the other hand, we exhibited above in 11.2 a closed (n-1)-form on $\mathbb{R}^n \setminus 0$ which is not exact.

2. Borsuk's theorem: There is no continuous map $D^n \to \partial D^n$ which is the identity on ∂D^n . We denote here by D^n the unit disc in \mathbb{R}^n and by ∂D^n its boundary (n-1)-sphere.

Proof. Suppose that there is such a map $f:D^n\to\partial D^n$. One can assume that f is smooth. Indeed, according to Theorem 11.3 one can approximate f by a smooth map, keeping it fixed on the boundary where it is the identity map, and hence smooth. Take the closed non-exact form θ_n from Example 11.2 on ∂D^n . Then $\Theta_n = f^*\theta_n$ is a closed (n-1)-form on D^n which coincides with

 θ_n on ∂D^n . D^n is star-shaped, and therefore Θ_n is exact, $\Theta_n = d\omega$. But then $\theta_n = d(\omega|_{\partial D^n})$ which is a contradiction.

3. Brouwer's fixed point theorem: Any continuous map $f:D^n\to D^n$ has at least 1 fixed point. **Proof.** Suppose $f:D^n\to D^n$ has no fixed points. Let us define a map $F:D^n\to \partial D^n$ as follows. For each $x\in D^n$ take a ray r_x from the point f(x) which goes through x till it intersects ∂D^n at a point which we will denote F(x). The map is well defined because for any x the points x and f(x) are distinct. Note also that if $x\in \partial D^n$ then the ray r_x intersects ∂D^n at the point x, and hence F(x)=x in this case. But existence of such F is ruled out by Borsuk' theorem.

k-connected manifolds

A subset $A \subset V$ is called k-connected, k = 0, 1, ..., if for any $m \leq k$ any two continuous maps of discs $f_0, f_1 : D^m \to A$ which coincide along ∂D^m are homotopic relative to ∂D^m . Thus, 0-connectedness is equivalent to path-connectedness. 1-connected submanifolds are also called *simply connected*.

Exercise 11.15. Prove that k-connectedness can be equivalently defined as follows: A is k-connected if any map $f: S^m \to A$, $m \le k$ is homotopic to a constant map.

- **Example 11.16.** 1. If A is contractible then it is k-connected for any k. For some classes of subsets, e.g. submanifolds, the converse is also true (J.H.C Whitehead's theorem) but this is a quite deep and non-trivial fact.
 - 2. The n-sphere S^n is (n-1)-connected but not n-connected. Indeed, to prove that S^{n-1} simply connected we will use the second definition. Consider a map $f: S^k \to S^n$. We first notice that according to Theorem 11.3 we can assume that the map f is smooth. Hence, according to Corollary 9.21 $\operatorname{Vol}_n f(S^k) = 0$ provided that k < n. In particular, there exists a point $p \in S^n \setminus f(S^k)$. But the complement of a point p in S^n is diffeomorphic to S^n vis the stereographic projection from the point p. But R^n is contractible, and hence f is homotopic to a constant map. On the other hand, the identity map $\operatorname{Id}: S^n \to S^n$ is not homotopic to a constant map. Indeed, we know that there exists a closed n-form on S^n , say the form θ_n

from Example 11.2, such that $\int_{S^n} \theta_n \neq 0$. Hence, $\int_{S^n} \mathrm{Id}^* \theta_n \neq 0$. On the other hand if Id were homotopic to a constant map this integral would vanish.

Exercise 11.17. Prove that $\mathbb{R}^{n+1} \setminus 0$ is (n-1)-connected but not n-connected.

Proposition 11.18. Let $U \subset V$ be a m-connected domain. Then for any $k \leq m$ any closed differential k-form α in U is exact.

Proof. We will prove here only the case m=1. Though the general case is not difficult, it requires developing certain additional tools. Let α be a closed differential 1-form. Choose a reference point $b \in U$. By assumption, U is path-connected. Hence, any other point x can be connected to b by a path $\gamma_x: [0,1] \to U$, i.e. $\gamma_x(0) = b$, $\gamma_x(1) = x$. Let us define the function $F: U \to \mathbb{R}$ by the formula $F(x) = \int_{\gamma_x} \alpha$. Note that due to the simply-connectedness of the domain U, any $\delta: [0,1] \to U$ connecting b and x is homotopic to γ_x relative its ends, and hence according to Lemma 11.8 we have $\int_{\gamma_x} \alpha = \int_{\delta} \alpha$. Thus the above definition of the function F is independent of the choice of paths γ_x . We claim that the function F is differentiable and $dF = \alpha$. Note that if the primitive of α exists than it has to be equal to F up to an additive constant. But we know that in a sufficiently small ball $B_{\epsilon}(a)$ centered at any point $a \in U$ there exists a primitive G of α . Hence, G(x) = F(x) + const, and the the differentiability of G implies differentiability of F and we have $dF = dG = \alpha$.

11.4 Winding and linking numbers

Given a loop $\gamma: S^1 \to \mathbb{R}^2 \setminus 0$ we define its winding number around 0 as the integral

$$w(\gamma) = \frac{1}{2\pi} \int_{S}^{1} \theta_2 = \frac{1}{2\pi} \int_{S^1} \frac{xdy - ydx}{x^2 + y^2},$$

where we orient S^1 as the boundary of the unit disc in \mathbb{R}^2 . For instance, if $j: S^1 \hookrightarrow \mathbb{R}^2$ is the inclusion map then w(j) = 1. For the loop γ_n parameterized by the map $t \mapsto (\cos nt, \sin nt), t \in [0, 2\pi]$ we have $w(\gamma_n) = n$.

Proposition 11.19. 1. For any loop γ the number $w(\gamma)$ is an integer.

2. If loops $\gamma_0, \gamma_1 : S^1 \to \mathbb{R}^2 \setminus 0$ are homotopic in $\mathbb{R}^2 \setminus 0$ then $w(\gamma_0) = w(\gamma_1)$.

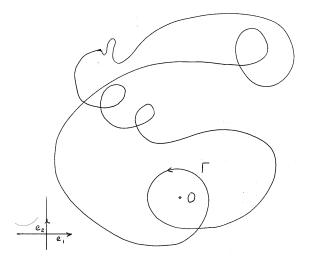


Figure 11.1: $w(\Gamma) = 2$.

3. $w(\gamma) = n$ then the loop γ is homotopic (as a loop in $\mathbb{R}^2 \setminus 0$) to the loop $\zeta_n : [0,1] \to \mathbb{R}^2 \setminus 0$ given by the formula $\zeta_n(t) = (\cos 2\pi nt, \sin 2\pi nt)$.

Proof. 1. Let us define the loop γ parametrically in polar coordinates:

$$r = r(s), \phi = \phi(s), s \in [0, 1],$$

where r(0) = r(1) and $\phi(1) = \phi(0) + 2n\pi$. The form θ_2 in polar coordinates is equal to $d\phi$, and hence

$$\int_{\gamma} \alpha = \frac{1}{2\pi} \int_{0}^{1} \phi'(s) ds = \frac{\phi(1) - \phi(0)}{2\pi} = n.$$

- 2. This is an immediate corollary of Proposition 11.8.
- 3. Let us write both loops γ and ζ_n in polar coordinates. Respectively,we have $r=r(t), \phi=\phi(s)$ for γ and $r=1, \phi=2\pi ns$ for $\zeta_n, s\in[0,1]$. The condition $w(\gamma)=n$ implies, in view of part 1, that $\phi(1)=\phi(0)+2n\pi$. Then the required homotopy $\gamma_t, t\in[0,1]$, connecting the loops $\gamma_0=\gamma$ and $\gamma_1=\zeta_n$ can be defined by the parametric equations $r=(1-t)r(s)+t, \phi=\phi_t(s)=(1-t)\phi(s)+2n\pi st$. Note that for all $t\in[0,1]$ we have $\phi_t(1)=\phi_t(0)+2n\pi$. Therefore, γ_t is a loop for all $t\in[0,1]$.

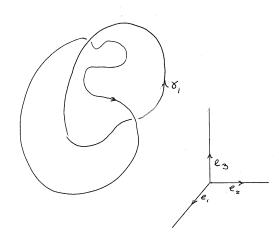


Figure 11.2: $l(\gamma_1, \gamma_2) = 1$.

Given two disjoint loops $\gamma, \delta: S^1 \to \mathbb{R}^3$ (i.e. $\gamma(s) \neq \delta(t)$ for any $s, t \in S^1$) consider a map $F_{\gamma,\delta}: T^2 \to \mathbb{R}^3 \setminus 0$, where $T^2 = S^1 \times S^1$ is the 2-torus, defined by the formula

$$F_{\gamma,\delta}(s,t) = \gamma(s) - \delta(t).$$

Then the number

$$l(\gamma, \delta) := \frac{1}{4\pi} \int\limits_{T}^{2} F_{\gamma, \delta}^* \theta_3 = \frac{1}{4\pi} \int\limits_{T}^{2} F_{\gamma, \delta}^* \left(\frac{xdy \wedge dz + ydz \wedge dx + zdx \wedge dy}{(x^2 + y^2 + z^2)^{\frac{3}{2}}} \right)$$

is called the $\mathit{linking}\ \mathit{number}$ of loops $\gamma, \delta.^1$

Exercise 11.20. Prove that

- 1. The number $l(\gamma, \delta)$ remains unchanged if one continuously deforms the loops γ, δ keeping them disjoint;
- 2. The number $l(\gamma, \delta)$ is an integer for any disjoint loops γ, δ ;

¹This definition of the linking number is due to Carl Friedrich Gauss.

- 3. $l(\gamma, \delta) = l(\delta, \gamma)$;
- 4. Let $\gamma(s) = (\cos s, \sin s, 0), s \in [0, 2\pi]$ and $\delta(t) = (-1 + \frac{1}{2}\cos t, 0, \frac{1}{2}\sin t), t \in [0, 2\pi]$. Then $l(\gamma, \delta) = 1$.

11.5 Properties of k-forms on k-dimensional manifolds

A k-form α on k-dimensional submanifold is always closed. Indeed, $d\alpha$ is a (k+1)-form and hence it is identically 0 on a k-dimensional manifold.

Remark 11.21. Given a k-dimensional submanifold $A \subset V$, and a k-form α on V, the differential $d_x \alpha$ does not need to vanish at a point $x \in A$. However, $d\alpha_x|_{T_x(A)}$ does vanish.

The following theorem is the main result of this section.

Theorem 11.22. Let $A \subset V$ be an orientable compact connected k-dimensional submanifold, possibly with boundary, and α a differential k-form on A.

- 1. Suppose that $\partial A \neq \emptyset$. Then α is exact, i.e. there exists a (k-1)-form β on A such that $d\beta = \alpha$.
- 2. Suppose that A is closed, i.e. $\partial A = \emptyset$. Then α is exact if and only if $\int_A \alpha = 0$.

To prove Theorem 11.22 we will need a few lemmas.

Lemma 11.23. Let I^k be the k-dimensional cube $\{-1 \le x_j \le 1, j = 1, \ldots, k\}$.

1. Let α be a differential k-form on I^k such that

$$\operatorname{Supp}(\alpha)\cap \left(0\times I^{k-1}\cup [0,1]\times \partial I^{k-1}\right)=\varnothing.$$

Then there exists a (k-1)-form β such that $d\beta = \alpha$ and such that $\operatorname{Supp}(\beta) \cap \left(-1 \times I^{k-1} \cup [-1,1] \times \partial I^{k-1}\right) = \emptyset$.

2. Let α be a differential k-form on I^k such that $\operatorname{Supp}(\alpha) \subset \operatorname{Int} I^k$ and $\int_{I^k} = 0$. Then there exists $a\ (k-1)$ -form β such that $d\beta = \alpha$ and $\operatorname{Supp}(\beta) \subset \operatorname{Int} I^k$.

Proof. We have

$$\alpha = f(x_1, \dots, x_k) dx_1 \wedge \dots \wedge dx_k,$$

In the first case of the lemma the function f vanishes on $0 \times I^{k-1} \cup [-1,1] \times \partial I^{k-1}$. We will look for β in the form

$$\beta = q(x_1, \dots, x_k) dx_2 \wedge \dots \wedge dx_k.$$

Then

$$d\beta = \frac{\partial g}{\partial x_1}(x_1, \dots, x_k)dx_1 \wedge dx_2 \wedge \dots \wedge dx_k.$$

and hence the equation $d\beta = \alpha$ is equivalent to

$$\frac{\partial g}{\partial x_1}(x_1,\ldots,x_k) = f(x_1,\ldots,x_k).$$

Hence, if we define

$$g(x_1,\ldots,x_k) := \int_{-1}^{x_1} f(u,x_2,\ldots,x_k) du,$$

then the form $\beta = g(x_1, \dots, x_k)dx_2 \wedge \dots \wedge dx_k$ has the required properties.

The second part of the lemma we will prove here only for the case k=2. The general case can be handled similarly by induction over k. We have in this case $\operatorname{Supp}(f)\subset\operatorname{Int} I^2$ and $\int\limits_{I^2}fdS=0$. Let us denote $h(x_2):=\int\limits_{-1}^1f(x_1,x_2)dx_1$. Note that h(u)=0 if u is sufficiently close to -1 or 1. According to Fubini's theorem, $\int\limits_{-1}^1h(x_2)dx_2=0$. We can assume that $f(x_1,x_2)=0$ for $x_1\geq 1-\epsilon$, and hence $\int\limits_{-1}^uf(x_1,\ldots,x_{k-1},t)dt=h(x_1,\ldots,x_{k-1})$ for $u\in[1-\epsilon,1]$. Consider any non-negative C^∞ -function $\theta:[1-\epsilon,1]\to\mathbb{R}$ such that $\theta(u)=1$ for $u\in[1-\epsilon,1-\frac{2\epsilon}{3}]$ and $\theta(u)=0$ for $u\in[1-\frac{\epsilon}{3},1]$. Define a function $g_1:I^2\to\mathbb{R}$ by the formula

$$g_1(x_1x_2) = \begin{cases} \int_{-1}^{x_1} f(u, x_2) du, & x_1 \in [-1, 1 - \epsilon], \\ h(x_2)\theta(x_1), & x_1 \in (1 - \epsilon, 1]. \end{cases}$$

Denote $\beta_1 = g_1(x_1, x_2)dx_2$. Then $d\beta = \alpha$ on $[-1, 1 - \epsilon] \times [0, 1]$ and $d\beta_1 = h(x_2)\theta'(x_1)dx_1 \wedge dx_2$ on $[1 - \epsilon, 1] \times [0, 1]$. Note that $\text{Supp}(\beta_1) \subset \text{Int } I^2$.

Let us define

$$g_2(x_1, x_2) := \begin{cases} 0, & x_1 \in [-1, 1 - \epsilon], \\ \theta'(x_1) \int_{-1}^{x_2} h(u) du, & x_1 \in (1 - \epsilon, 1] \end{cases}$$

and denote $\beta_2 = g_2(x_1, x_2)dx_1$. Then $d\beta_2 = 0$ on $[-1, 1 - \epsilon] \times [-1, 1]$ and

$$d\beta_2 = -h(x_2)\theta'(x_1)dx_1 \wedge dx_2$$

on $[1-\epsilon,1]\times[-1,1]$. Note that $g_2(x_1,1)=-\theta'(x_1)\int\limits_{-1}^1h(u)du=0$. Taking into account that h(u)=0 when u is sufficiently close to -1 or 1 we conclude that $h(x_1,x_2)=0$ near ∂I^2 , i.e. $\mathrm{Supp}(\beta_2)\subset\mathrm{Int}\,I^2$.

The following lemma is a special case of the, so-called, tubular neighborhood theorem.

Lemma 11.24. Let $A \subset V$ be a compact k-dimensional submanifold with boundary. Let $\phi: [-1,1] \to A$ be an embedding such that $\phi(1) \in \partial A$, $\phi'(1) \perp T_{\phi(1)}(\partial A)$ and $\phi([0,1)) \subset \operatorname{Int} A$. Then the embedding ϕ can be extended to an embedding $\Phi: [-1,1] \times I^{k-1} \to A$ such that

- $\Phi(t,0) = \phi(t)$, for $t \in [-1,1]$, $0 \in I^{k-1}$;
- $\Phi(1 \times I^{k-1}) \subset \partial A$, $\Phi([-1,1) \times I^{k-1}) \subset \operatorname{Int} A$;
- $\frac{\partial \Phi}{\partial t}(1,x) \notin T(\partial A)$ for all $x \in I^{k-1}$.

There are many ways to prove this lemma. We will explain below one of the arguments.

Proof. Step 1. We first construct k-1 ortonormal vector fields ν_1, \ldots, ν_k along $\Gamma = \phi([-1, 1])$ which are tangent to A and normal to Γ . To do that let us denote by N_u the normal (k-1)-dimensional space N_u to $T_u\Gamma$ in T_uA . Let us observe that in view of compactness of Γ there is an $\epsilon > 0$ with the following property: for any two points $u = \phi(t), u' = \phi(t') \in \Gamma$, $t, t' \in [-1, 1]$, such that $|t-t'| \leq \epsilon$ the orthogonal projection $N_u \to N_{u'}$ is non-degenerate (i.e. is an isomorphism). Choose $N < \frac{1}{2\epsilon}$ and consider points $u_j = \phi(t_j)$, where $t_j = -1 + \frac{2j}{N}, j = 1, \ldots N$. Choose any orthonormal basis $\nu_1(0), \ldots, \nu_k(0) \in N_{u_0}$, parallel transport these vectors to all points of the arc $\Gamma_1 = \phi([-1, t_1])$, project them orthogonally to the normal spaces N_u in these points, and then orthonormalize the resulted bases via the Gram-Schmidt process. Thus we constructed orthonormal vector fields $\nu_1(t), \ldots \nu_k(t) \in N_{\phi(t)}, t \in [-1, t_1]$. Now we repeat this procedure beginning with the

basis $\nu_1(t_1), \dots \nu_k(t_1) \in N_{\phi(t_1)} = N_{u_1}$ and extend the vector fields ν_1, \dots, ν_k to $\Gamma_2 = \phi([t_1, t_2])$. Continuing this process we will construct the orhonormal vector fields ν_1, \dots, ν_k along the whole curve Γ .²

Step 2. Consider a map $\Psi: [-1,1] \times I^{k-1} \to V$ given by the formula

$$\Psi(t, x_1, \dots, x_{k-1}) = \phi(t) + \sigma \sum_{j=1}^{k-1} x_j \nu_j(t), \ t, x_1, \dots, x_{k-1} \in [-1, 1],$$

where a small positive number σ will be chosen later. The map Ψ is an embedding if σ is chosen small enough.³ Unfortunately the image $\Psi([-1,1] \times I^{k-1})$ is not contained in A. We will correct this in the next step.

STEP 3. Take any point $a \in A$ and denote by π_a the orthogonal projection $V \to T_a A$. Let us make the following additional assumption (in the next step we will show how to get rid of it): there exists a neighborhood $U \ni a = \phi(1)$ in ∂A such that $\pi_a(U) \subset N_a \subset T_a A$. Given $\epsilon > 0$ let us denote by $B_{\epsilon}(a)$ the (k-1)-dimensional ball of radius ϵ in the space $N_{\alpha} \subset T_a A$. In view of compactness of A one can choose an $\epsilon > 0$ such that for all points $a \in \Gamma$ there exists an embedding $e_a : B_{\epsilon}(a) \to A$ such that $\pi_a \circ e_a = \operatorname{Id}$. Then for a sufficiently small $\sigma < \frac{\epsilon}{\sqrt{k-1}}$ the map $\widetilde{\Psi} : [-1,1] \times I^{k-1} \to A$ defined by the formula

$$\widetilde{\Psi}(t,x) = e_{\phi(t)} \circ \Psi(t,x), \ t \in [-1,1], x \in I^{k-1}$$

is an embedding with the required properties.

STEP 4 It remains to show how to satisfy the additional condition at the boundary point $\phi(1) \in \Gamma \cap \partial A$ which were imposed above in Step 3. Take the point $a = \phi(1) \in \Gamma \cap \partial A$. Without loss of generality we can assume that $a = 0 \in V$. Choose an orthonormal basis v_1, \ldots, v_n of V such that $v_1, \ldots, v_k \in N_a$ and v_k is tangent to Γ and pointing inward Γ . Let (y_1, \ldots, y_n) be the corresponding cartesian coordinates in V. Then there exists a neighborhood $U \ni a$ in A which is graphical in these coordinates and can be given by

$$y_j = \theta_j(y_1, \dots, y_k), \ j = k+1, \dots, n, \ \epsilon \ge y_k \ge \theta_k(y_1, \dots, y_{k-1}), |y_i| \le \epsilon, i = 1, \dots, k-1,$$

³Exercise: prove it!

²Strictly speaking, the constructed vector fields only piece-wise smooth, because we did not make any special precautions to ensure smoothness at the points $u_j, j = 1, ..., N-1$. This could be corrected via a standard smoothing procedure.

where all the first partial derivatives of the functions $\theta_k, \ldots, \theta_n$ vanish at the origin. Take a C^{∞} cut-off function $\sigma: [0, \infty)] \to \mathbb{R}$ which is equal to 1 on $[0, \frac{1}{2}]$ and which is supported in [0, 1] (see Lemma 9.2). Consider a map F given by the formula

$$F(y_1, \dots, y_n) = (y_1, \dots, y_{k-1}, y_k - \theta_k(y_1, \dots, y_{k-1})\sigma\left(\frac{||y||}{\epsilon}\right), y_{k+1}, \dots, y_n).$$

For a sufficiently small $\epsilon > 0$ this is a diffeomorphism supported in an ϵ -ball in V centered in the origin. On the other hand, the manifold $\widetilde{A} = F(A)$ satisfies the extra condition of Step 3.

Lemma 11.25. Let $A \subset V$ be a (path)-connected submanifold with a non-empty boundary. Then for any point $a \in A$ there exists an embedding $\phi_a : [-1,1] \to A$ such that $\phi_a(0) = a$, $\phi_a(1) \in \partial A$ and $\phi'_a(1) \perp T_{\phi_a(1)}(\partial A)$.

Sketch of the proof. Because A is path-connected with non-empty boundary, any interior point can be connected by a path with a boundary point. However, this path need not be an embedding. First, we perturb this path to make it an immersion $\psi: [-1,1] \to A$, i.e. a map with non-vanising derivative. This can be done as follows. As in the proof of the previous lemma we consider a sufficiently small partition of the path, so that two neighboring subdivision points lie in a coordinate neighborhood. Then we can connect these points by a straight segment in these coordinate neighborhoods. Finally we can smooth the corners via the standard smoothing procedure. Unfortunately the constructed immersed path ψ may have self-intersection points. First, one can arrange that there are only finitely many intersections, and then "cut-out the loops", i.e. if $\psi(t_1) = \psi(t_2)$ for $t_1 < t_2$ we can consider a new piece-wise smooth path which consists of $\psi|_{[-1,t_1]}$ and $\psi|_{[t_2,1]}$ The new path has less self-intersection points, and thus continuing by induction we will end with a piece-wise smooth embedding. It remains to smooth again the corners.

Proof of Theorem 11.22. 1. For every point $a \in A$ choose an embedding $\phi_a : [-1,1] \to A$, as in Lemma 11.25, and using Lemma 11.24 extend ϕ_a to an embedding $\Phi_a : [-1,1] \times I^{k-1} \to A$ such that

-
$$\Phi_a(t,0) = \phi(t)$$
, for $t \in [-1,1]$, $0 \in I^{k-1}$;

-
$$\Phi_a(1 \times I^{k-1}) \subset \partial A$$
, $\Phi_a([-1,1) \times I^{k-1}) \subset \operatorname{Int} A$;

-
$$\frac{\partial \Phi_a}{\partial t}(1,x) \notin T(\partial A)$$
 for all $x \in I^{k-1}$.

Due to compactness of A we can choose finitely many such embeddings $\Phi_j = \Phi_{a_j}, \ j = 1, \ldots, N$, such that $\bigcup_{1}^{N} \Phi_j((-1,1] \times \operatorname{Int}(I^{k-1})) = A$. Choose a partition of unity subordinated to this covering and split the k-form α as a sum $\alpha = \sum_{1}^{K} \alpha_j$, where each α_i is supported in $\Phi_j((-1,1] \times \operatorname{Int}(I^{k-1}))$ for some $j = 1, \ldots, N$. To simplify the notation we will assume that N = K and each α_j is supported in $\Phi_j((-1,1] \times \operatorname{Int}(I^{k-1})), \ j = 1, \ldots, N$. Consider the pull-back form $\widetilde{\alpha}_j = \Phi_j^* \alpha_j$ on $I^k = [-1,1] \times I^{k-1}$. According to Lemma 11.23.1 there exists a (k-1)-form $\widetilde{\beta}_j$ such that $\operatorname{Supp}(\widetilde{\beta}_j) \subset (-1,1] \times \operatorname{Int}(I^{k-1})$ and $d\widetilde{\beta}_j = \widetilde{\alpha}_j$. Let us transport the form $\widetilde{\beta}_j$ back to A. Namely, set β_j equal to $(\Phi_j^{-1})^*\widetilde{\beta}_j$ on $\Phi_j((-1,1] \times \operatorname{Int}(I^{k-1})) \subset A$ and extend it as 0 elsewhere on A. Then $d\beta_j = \alpha_j$, and hence $d(\sum_{1}^{N} \beta_j) = \sum_{1}^{N} \alpha_j = \alpha$.

2. Choose a point $a \in A$ and parameterize a coordinate neighborhood $U \subset A$ by an embedding $\Phi: I^k \to A$ such that $\Phi(0) = a$. Take a small closed ball $D_{\epsilon}(0) \subset I^k \subset \mathbb{R}^k$ and denote $\widetilde{D} = \Phi(D_{\epsilon}(0))$. Then $\widetilde{A} = A \setminus \operatorname{Int} \widetilde{D}$ is a submanifold with non-empty boundary, and $\partial \widetilde{A} = \partial \widetilde{D}$. Let us use part 1 of the theorem to construct a form $\widetilde{\beta}$ on \widetilde{A} such that $d\widetilde{\beta} = \alpha|_{\widetilde{A}}$. Let us extent the form $\widetilde{\beta}$ in any way to a form, still denoted by $\widetilde{\beta}$ on the whole submanifold A. Then $d\widetilde{\beta} = \alpha + \eta$ where $\operatorname{Supp}(\eta) \subset \widetilde{D} \subset \operatorname{Int} \Phi(I^k)$. Note that

$$\int\limits_{\Phi(I^k)} \eta = \int\limits_A \eta = \int\limits_A \alpha - \int\limits_A d\widetilde{\beta} = 0$$

because $\int_A \alpha = 0$ by our assumption, and $\int_A d\widetilde{\beta} = 0$ by Stokes' theorem. Thus, $\int_I^k \Phi^* \eta = 0$, and hence, we can apply Lemma 11.23.2 to the form $\Phi^* \eta$ on I^k and construct a (k-1)-form λ on I^{k-1} such that $d\lambda = \Phi^* \eta$ and $\operatorname{Supp}(\lambda) \subset \operatorname{Int} I^k$. Now we push-forward the form λ to A, i.e. take the form $\widetilde{\lambda}$ on A which is equal to $(\Phi^{-1})^* \lambda$ on $\Phi(I_k)$ and equal to 0 elsewhere. Finally, we have $d(\widetilde{\beta} + \widetilde{\lambda}) = d\widetilde{\beta} + \eta = \alpha$, and hence $\beta = \widetilde{\beta} + \widetilde{\lambda}$ is the required primitive of α on A.

Corollary 11.26. Let A be an oriented compact connected k-dimensional submanifold with nonempty boundary and α a differential k-form on A from Theorem 11.22. Then for any smooth map $f: A \to A$ such that $f|_{\partial A} = \operatorname{Id}$ we have

$$\int_{A} f^* \alpha = \int_{A} \alpha.$$

Proof. According to Theorem 11.22.1 there exists a form β such that $\alpha = d\beta$. Then

$$\int_A f^* \alpha = \int_A f^* d\beta = \int_A df^* \beta = \int_{\partial A} f^* \beta = \int_A \beta = \int_A \alpha.$$

Degree of a map

Consider two closed connected oriented submanifolds $A \subset V$, $B \subset W$ of the same dimension k. Let ω be an n-form on B such that $\int_B \omega = 1$. Given a smooth map $f: A \to B$ the integer $\deg(f) := \int_A f^* \omega$ is called the degree of the map f.

Proposition 11.27. 1. Given any two k-forms on B such $\int_B \omega = \int_B \widetilde{\omega}$ we have $\int_A f^*\omega = \int_A f^*\widetilde{\omega}$, for any smooth map $f: A \to B$, and thus $\deg(f)$ is independent of the choice of the form ω on B with the property $\int_A \omega = 1$.

- 2. If the maps $f, g: A \to B$ are homotopic then $\deg(f) = \deg(g)$.
- 3. Let $b \in B$ be a regular value of the map f. Let $f^{-1}(b) = \{a_1, \ldots, a_d\}$. Then

$$\deg(f) = \sum_{1}^{d} \operatorname{sign}(\det Df(a_j)).$$

In particular, deg(f) is an integer number.

Proof. The second part follows from Lemma 11.8. To prove the first part, let us write $\widetilde{\omega} = \omega + \eta$, where $\int_B \eta = 0$. Using Theorem 11.22.2 we conclude that $\eta = d\beta$ for some (k-1)-form β on B. Then

$$\int_A f^* \widetilde{\omega} = \int_A f^* \omega + \int_A f^* \eta = \int_A f^* \omega + \int_A df^* \beta = \int_A f^* \omega.$$

Let us prove the last statement of the theorem. By the inverse function theorem there exists a neighborhood $U \ni b$ in B and neighborhoods $U_1 \ni a_1, \ldots, U_d \ni a_d$ in A such that the restrictions of the map f to the neighborhoods U_1, \ldots, U_d are diffeomorphisms $f|_{U_j}: U_j \to U, \ j = 1, \ldots, d$. Let us consider a form ω on B such that $\operatorname{Supp}\omega \subset U$ and $\int_B \omega = \int_U \omega = 1$. Then

$$\deg(f) = \int_A f^*\omega = \sum_1^d \int_{U_i} f^*\omega = \sum_1^d \operatorname{sign}(\det Df(a_j)),$$

because according to Theorem ?? we have

$$\int_{U_j} f^* \omega = \operatorname{sign}(\det Df(a_j)) \int_{U} \omega = \operatorname{sign}(\det Df(a_j)).$$

for each $j = 1, \ldots, d$.

Remark 11.28. Any continuous map $f: A \to B$ can be approximated by a homotopic to f smooth map $A \to B$, and any two such smooth approximations of f are homotopic. Hence this allows us to define the degree of any *continuous* map $f: A \to B$.

Exercise 11.29. 1. Let us view \mathbb{R}^2 as \mathbb{C} . In particular, we view the unit sphere $S^1 = S_1^1(0)$ as the set of complex numbers of modulus 1:

$$S^1 = \{ z \in \mathbb{C}; |z| = 1 \}.$$

Consider a map $h_n: S^1 \to S^1$ given by the formula $h_n(z) = z^n, z \in S^1$. Then $\deg(h_n) = n$.

2. Let $f: S^{n-1} \to S^{n-1}$ be a map of degree d. Let p_{\pm} be the north and south poles of S^{n+1} , i.e. $p_{\pm} = (0, \dots, 0, \pm 1)$. Given any point $x = (x_1, \dots, x_{n+1}) \in S^n \setminus \{p_+, p_-\}$ we denote by $\pi(x)$ the point

$$\frac{1}{\sqrt{\sum_{1}^{n} x_j^2}} (x_1, \dots, x_n) \in S^{n-1}$$

and define a map $\Sigma f:S^n \to S^n$ by the formula

$$\Sigma f(x) = \begin{cases} p_{\pm}, & \text{if } x = p_{\pm}, \\ \left(\sqrt{\sum_{1}^{n} x_{j}^{2}} f(\pi(x)), x_{n+1}\right), & \text{if } x \neq p_{\pm}. \end{cases}$$

Prove that $deg(\Sigma(f)) = d.^4$

- 3. Prove that two maps $f, g: S^n \to S^n$ are homotopic if and only if they have the same degree. In particular, any map of degree n is homotopic to the map h_n . (Hint: For n=1 this follows from Proposition 11.19. For n > 1 first prove that any map is homotopic to a suspension.)
- 4. Give an example of two non-homotopic orientation preserving diffeomorphisms $T^2 \to T^2$. Note that the degree of both these maps is 1. Hence, for manifolds, other than spheres, having the same degree is not sufficient for their homotopy.

⁴The map Σf is called the *suspension* of the map f.

5. Let $\gamma, \delta: S^1 \to \mathbb{R}^3$ be two disjoint loops in \mathbb{R}^3 . Consider a map $\widetilde{F}_{\gamma,\delta}: T^2 \to S^2$ defined by the formula

$$\widetilde{F}_{\gamma,\delta}(s,t) = \frac{\gamma(s) - \delta(t)}{||\gamma(s) - \delta(t)||}, \ s,t \in S^1.$$

Prove that $l(\gamma, \delta) = \deg(\widetilde{F}_{\gamma, \delta})$. Use this to solve Exercise 11.20.4 above.